Black Holes-galaxies co-evolution
Physical Models

V. Antonuccio-Delogu
INAF-Catania Astrophysical Observatory

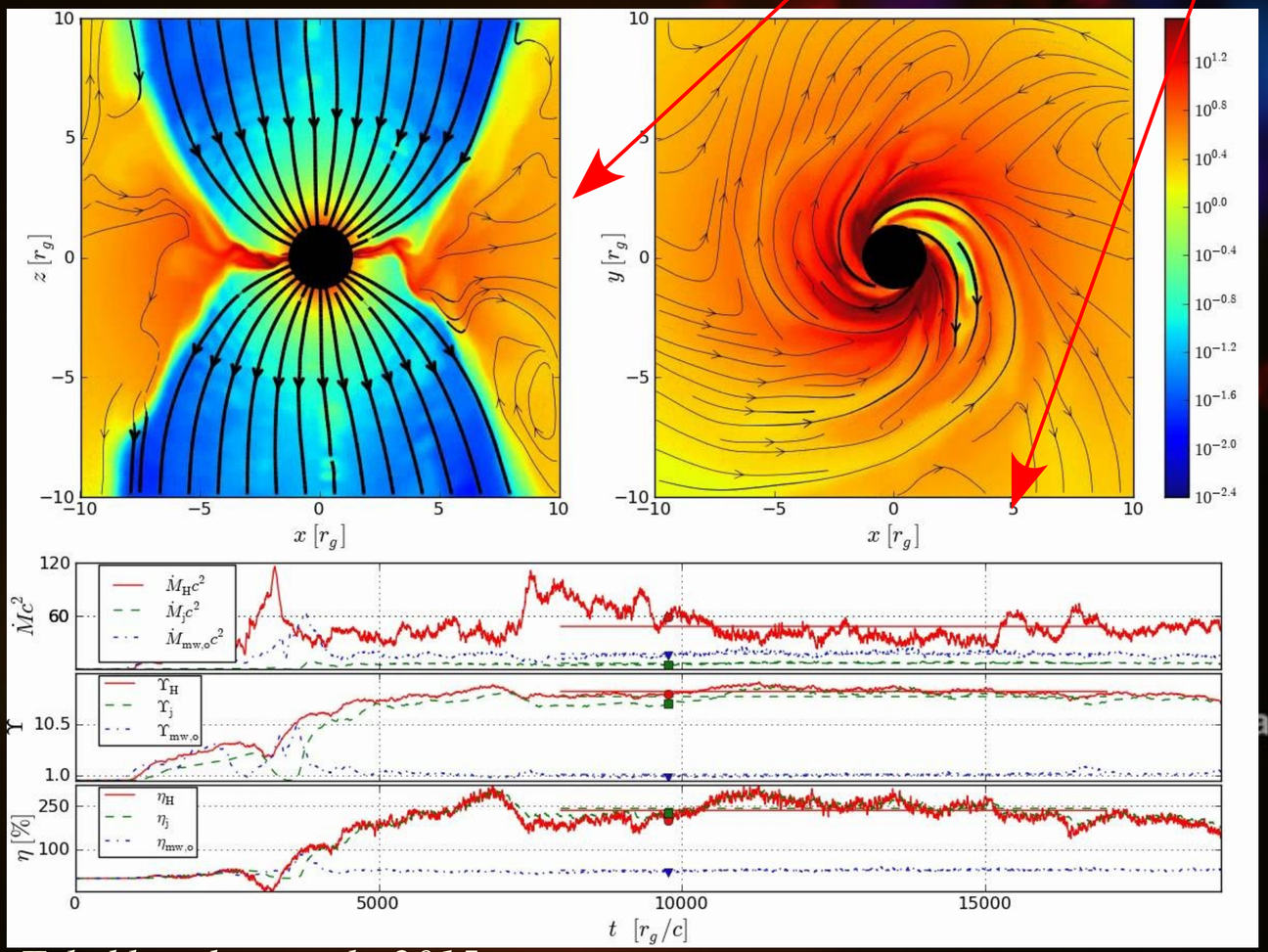
Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

Galaxy-SMBH connection: the numerical gap

Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

Galaxy-SMBH connection: the numerical gap

- Accretion/jet production region: $R_d \sim 10^3\ r_g$ $(\sim 4.78 * 10^{-3}\ \mathrm{pc})$



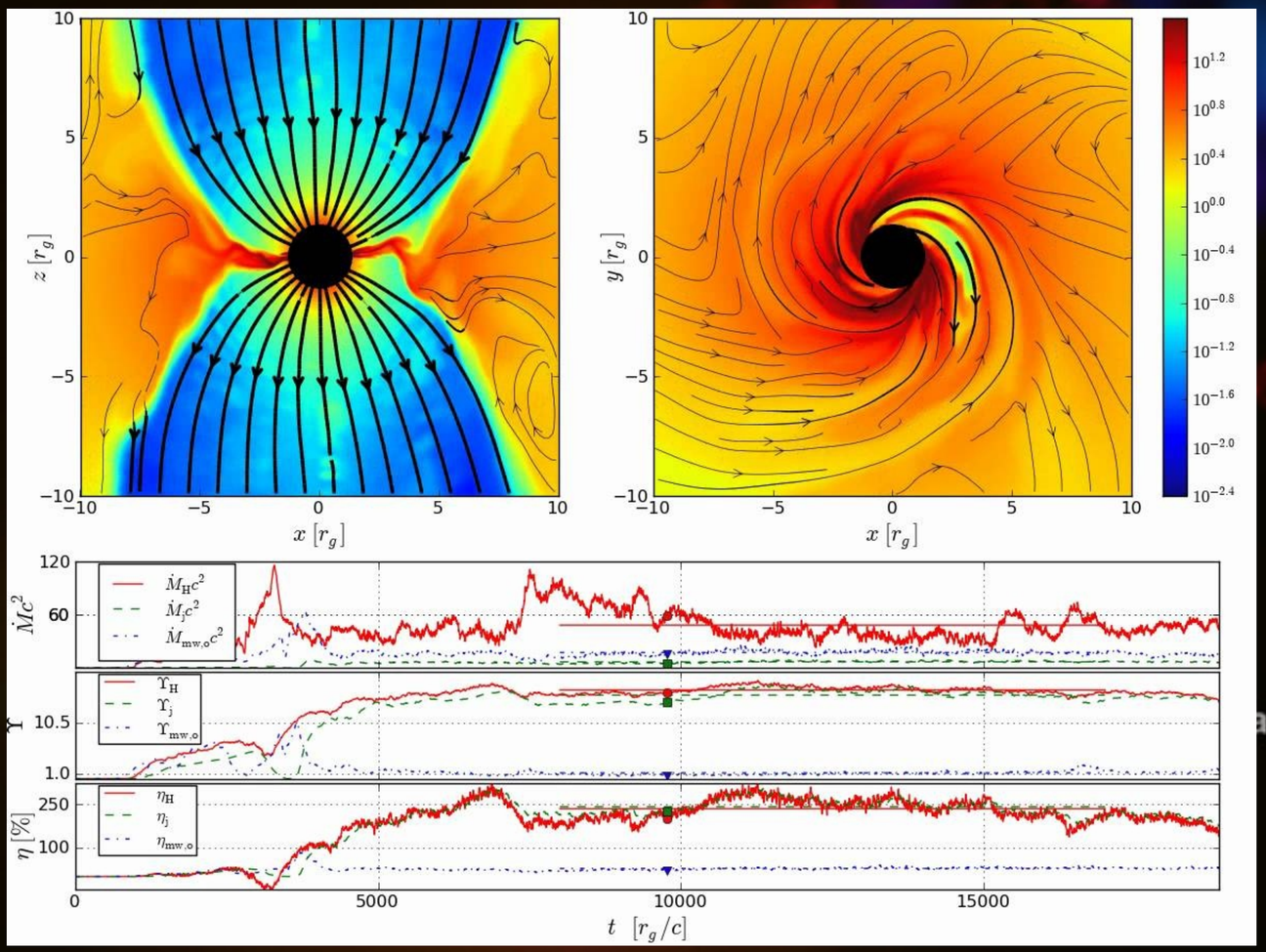
Tchekhovskoy et al., 2015

axy System 3C 321 HUBBLESITE.org

Galaxy-SMBH connection: the numerical gap

- Accretion/jet production region: $R_d \sim 10^3\ r_g$ (~ $4.78 * 10^{-3}$ pc)

AMR (Gaspari, Brüggen, Fong, Bicknell, Wagner, Cielo+V.A.-D.,..):
l_{max}=10-15 → $R = L_{box}/2^{lmax} \simeq$ 1.52 - 48.8 pc, <u>3 orders of magnitude larger than</u> R_d production



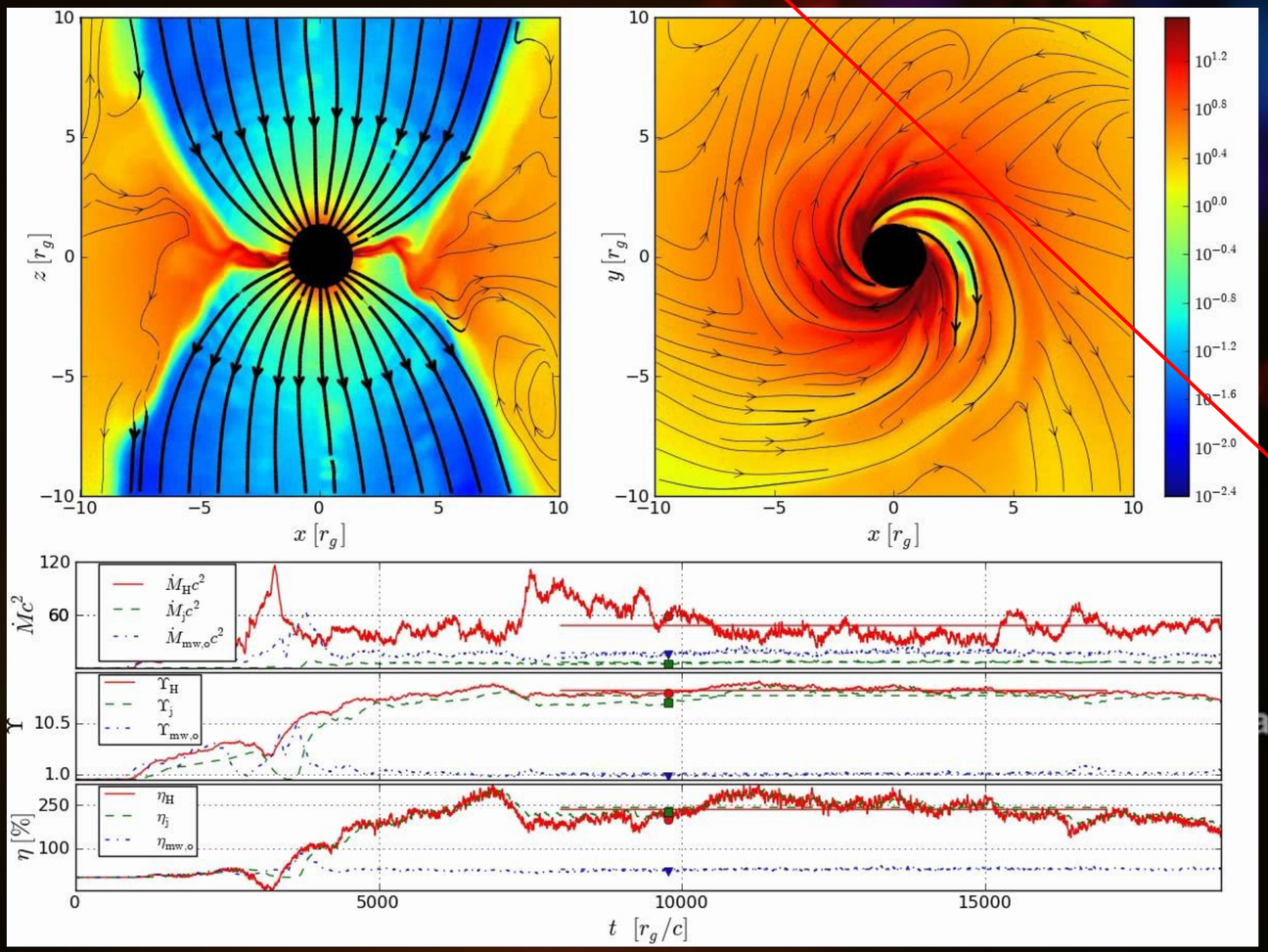
Tchekhovskoy et al., 2015

axy System 3C 321 HUBBLESITE.org

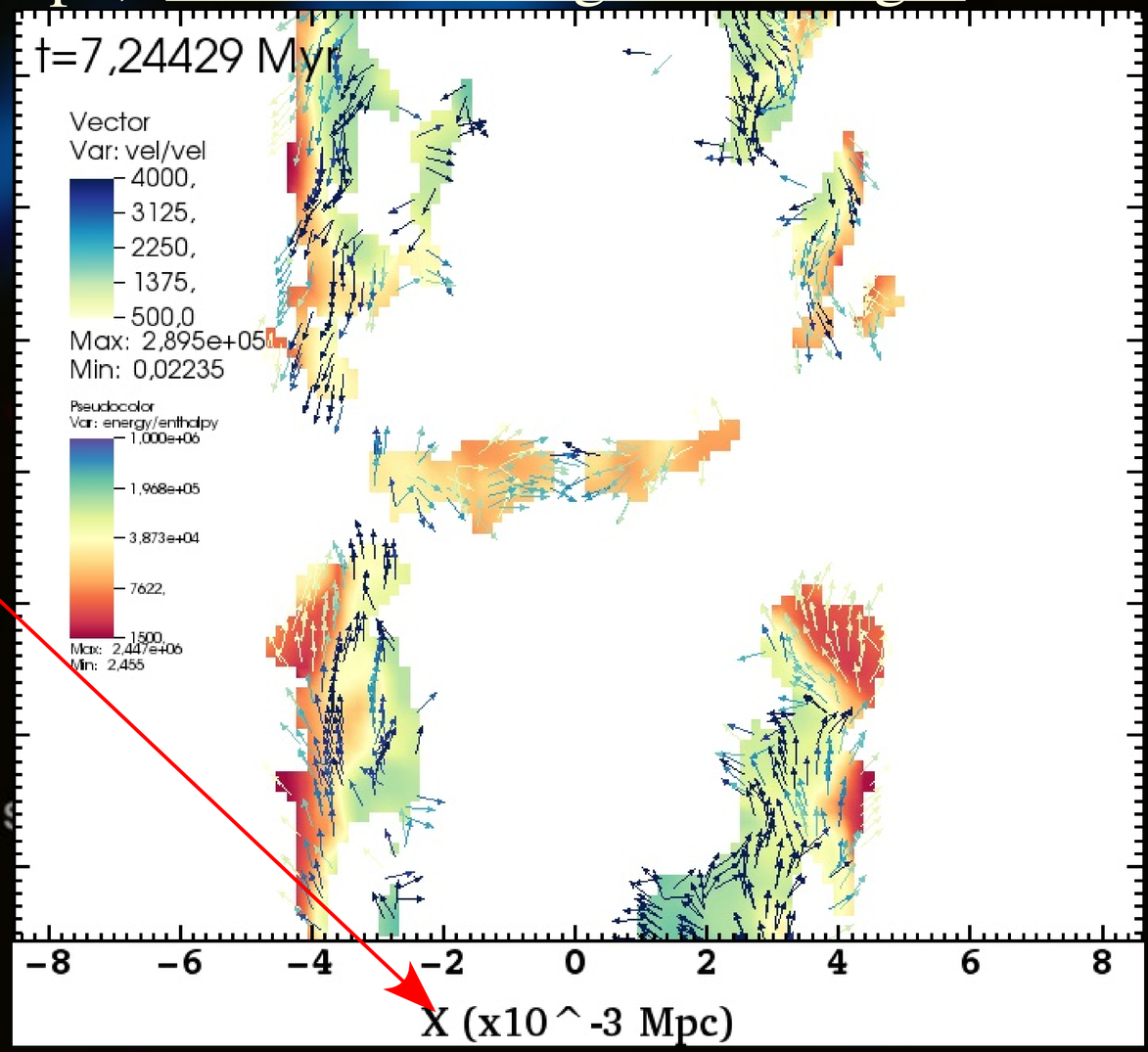
Galaxy-SMBH connection: the numerical gap

- Accretion/jet production region: $R_d \sim 10^3\ r_g$ $(\sim 4.78 * 10^{-3}$ pc)

AMR (Gaspari, Brüggen, Fong, Bicknell, Wagner, Cielo+V.A.-D.,..):
l_{max}=10-15 → $R = L_{box}/2^{lmax} \simeq$ 1.52 - 48.8 pc, <u>3 orders of magnitude larger than</u> R_d production



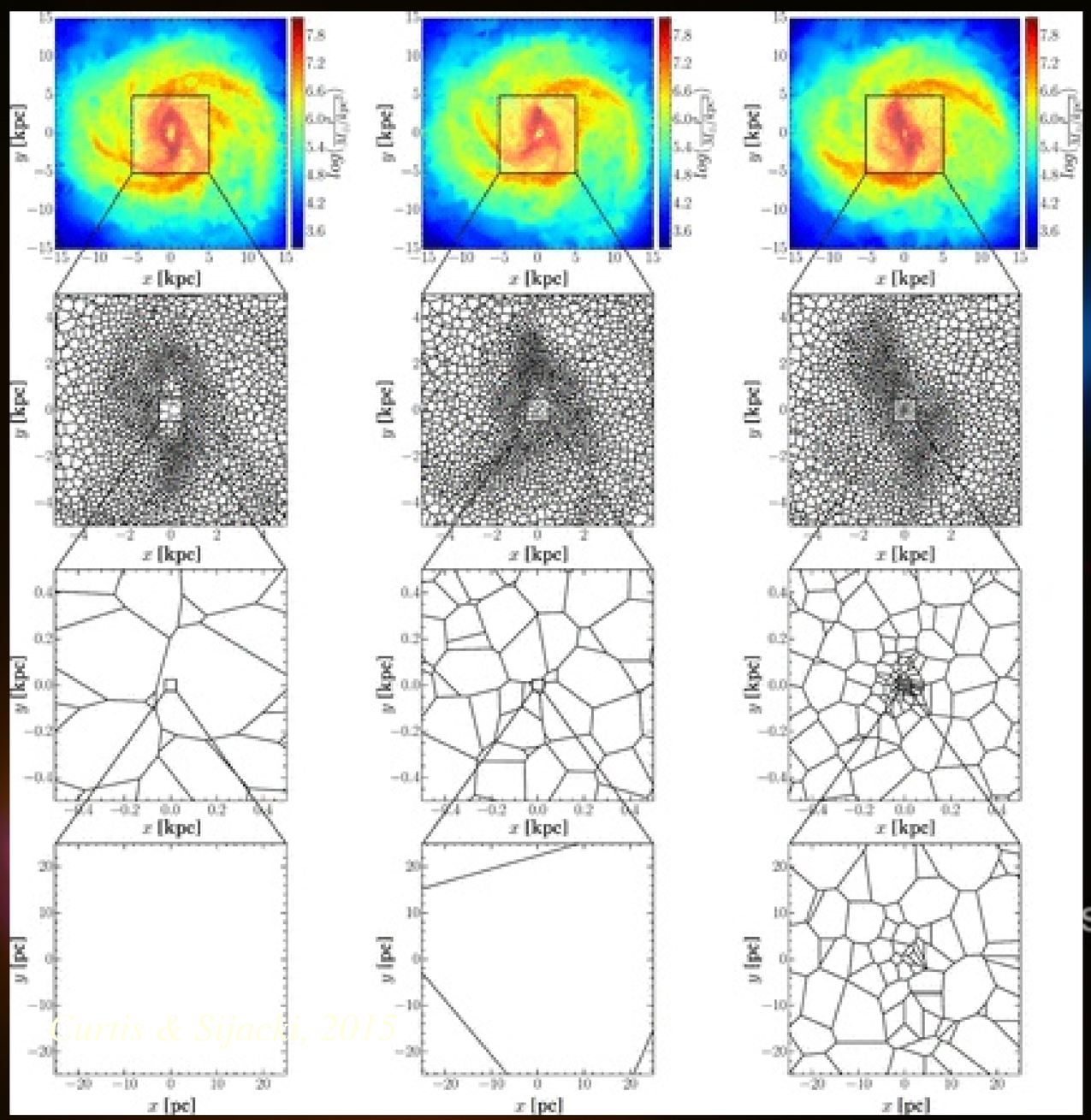
Tchekhovskoy et al., 2015



Cielo et al et al., 2015

Galaxy-SMBH connection: HOW TO FILL the numerical gap?

- AREPO (TreeSPH, *Springel*) + HPC parallel *simulations*



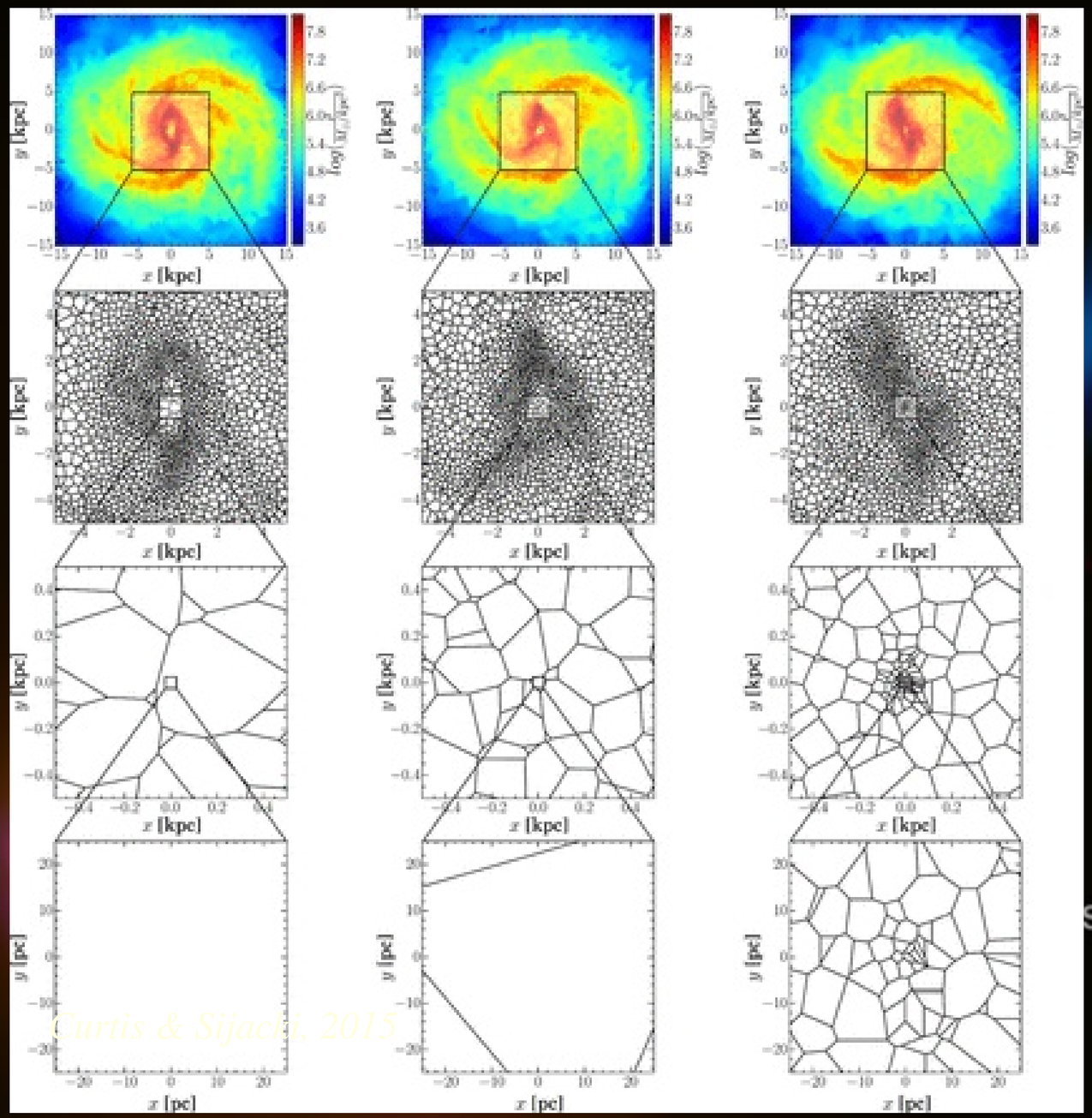
Curtis & Sijacki., 2015

System 3C 321 HUBBLESITE.org

Galaxy-SMBH connection: HOW TO FILL the numerical gap?

- AREPO (TreeSPH, *Springel*) + HPC parallel *simulations*

➔ Resolves accretion region (~$10^4\,r_g$) within a <u>single galaxy</u>



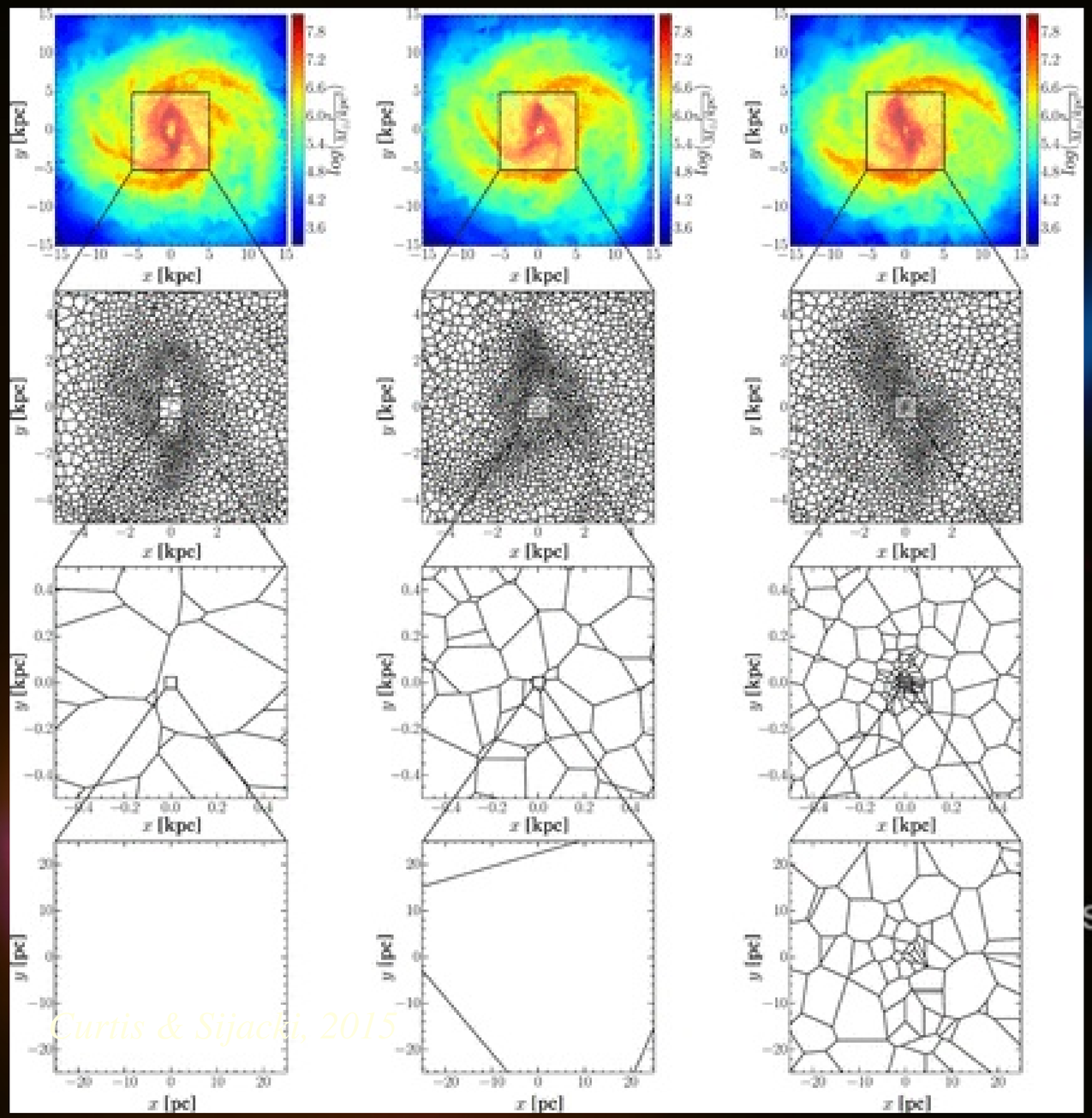
Curtis & Sijacki., 2015

System 3C 321 HUBBLESITE.org

Galaxy-SMBH connection: HOW TO FILL the numerical gap?

- AREPO (TreeSPH, *Springel*) + HPC parallel *simulations*

➔ Resolves accretion region (~10^4 r_g) within a <u>single galaxy</u>



Curtis & Sijacki, 2015

- High resolution → <u>backflows</u> from the large-scale jet are resolved

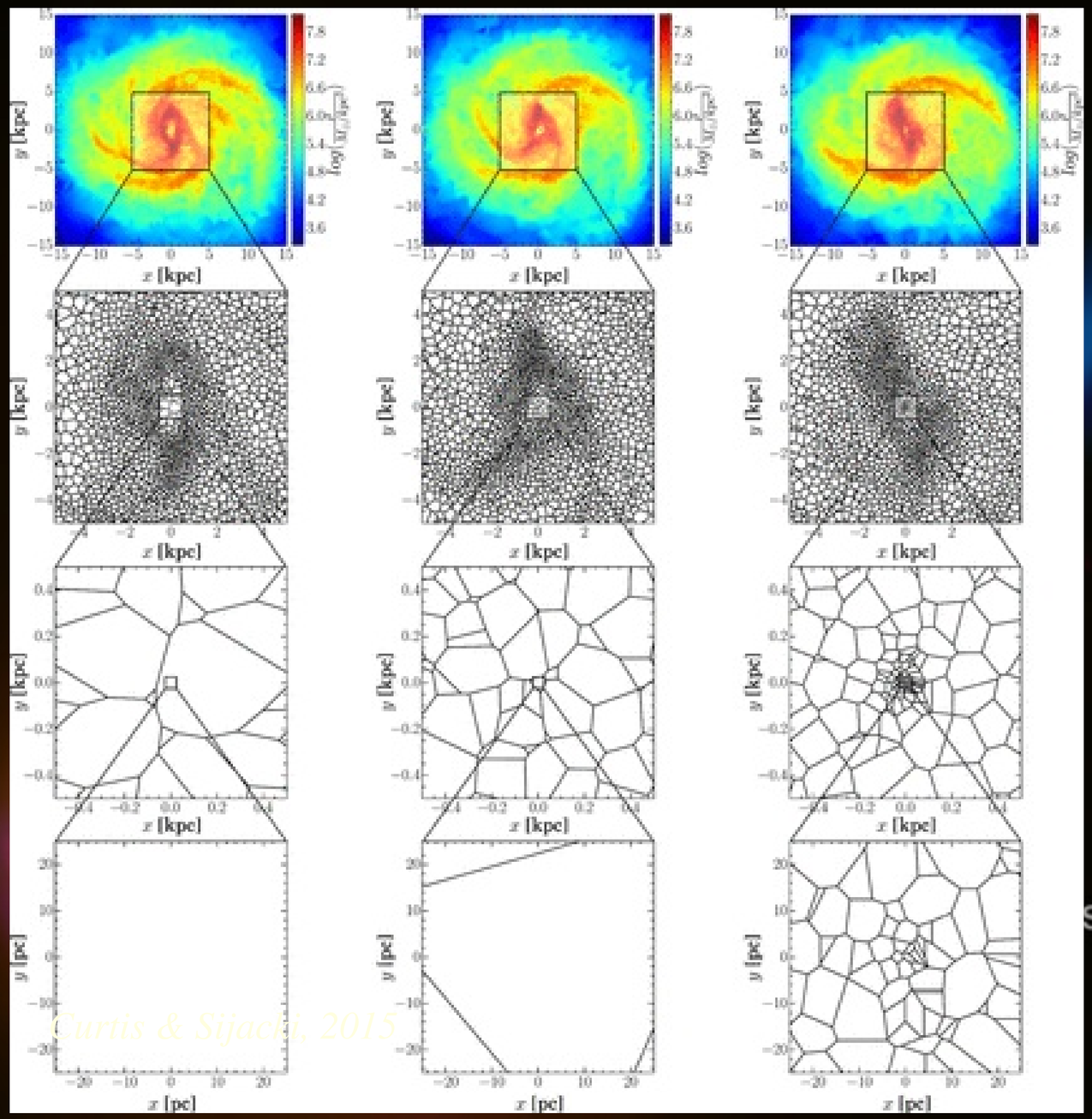
System 3C 321 HUBBLESITE.org

Curtis & Sijacki., 2015

Galaxy-SMBH connection: HOW TO FILL the numerical gap?

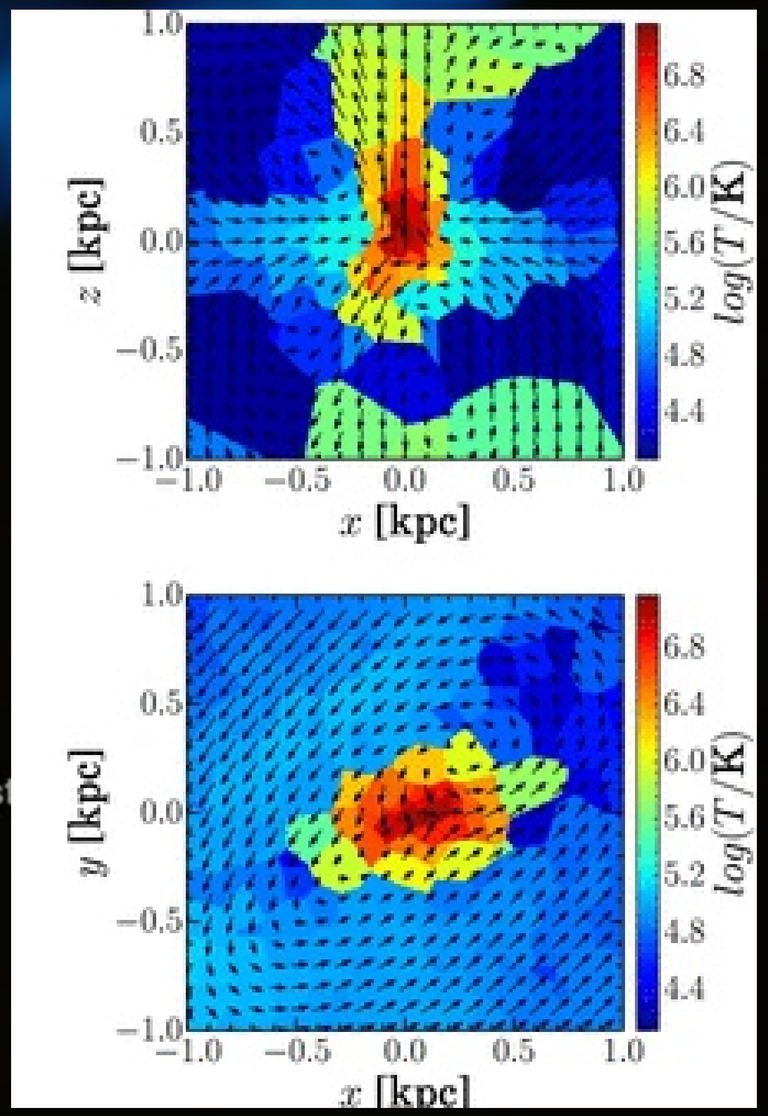
- AREPO (TreeSPH, *Springel*) + HPC parallel *simulations*

➔ Resolves accretion region (~$10^4\,r_g$) within a <u>single galaxy</u>



Curtis & Sijacki., 2015

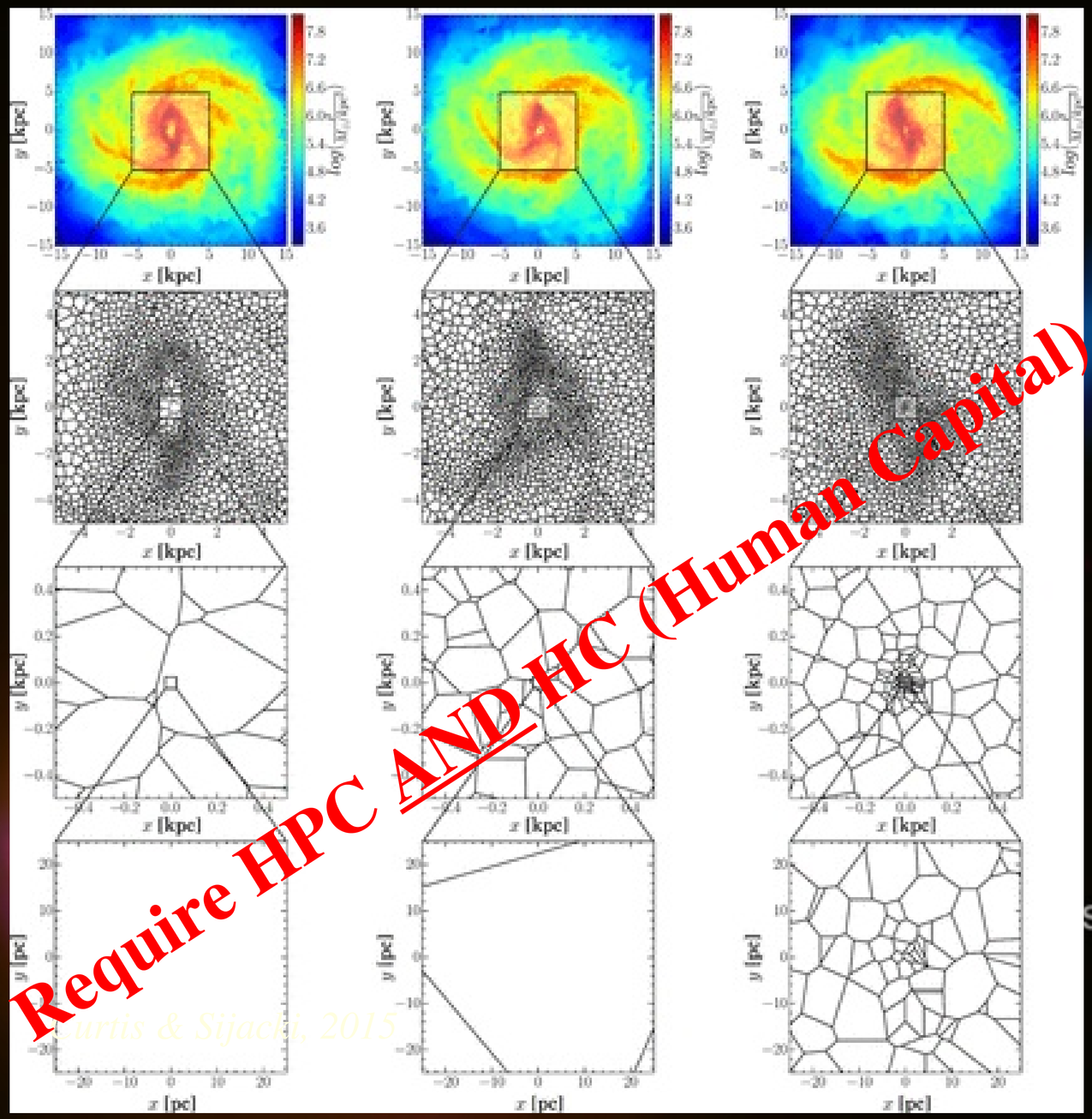
- High resolution → <u>backflows</u> from the large-scale jet are resolved



Galaxy-SMBH connection: HOW TO FILL the numerical gap?

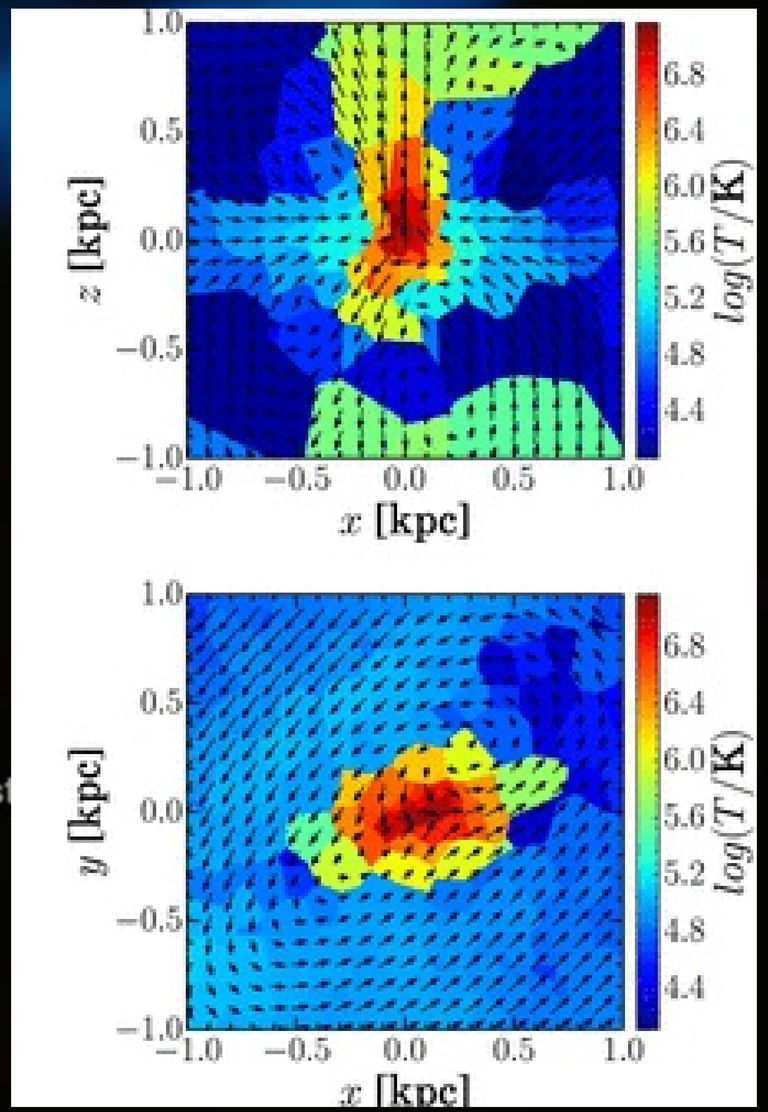
- AREPO (TreeSPH, *Springel*) + HPC parallel *simulations*

➔ Resolves accretion region (~$10^4\,r_g$) within a <u>single galaxy</u>



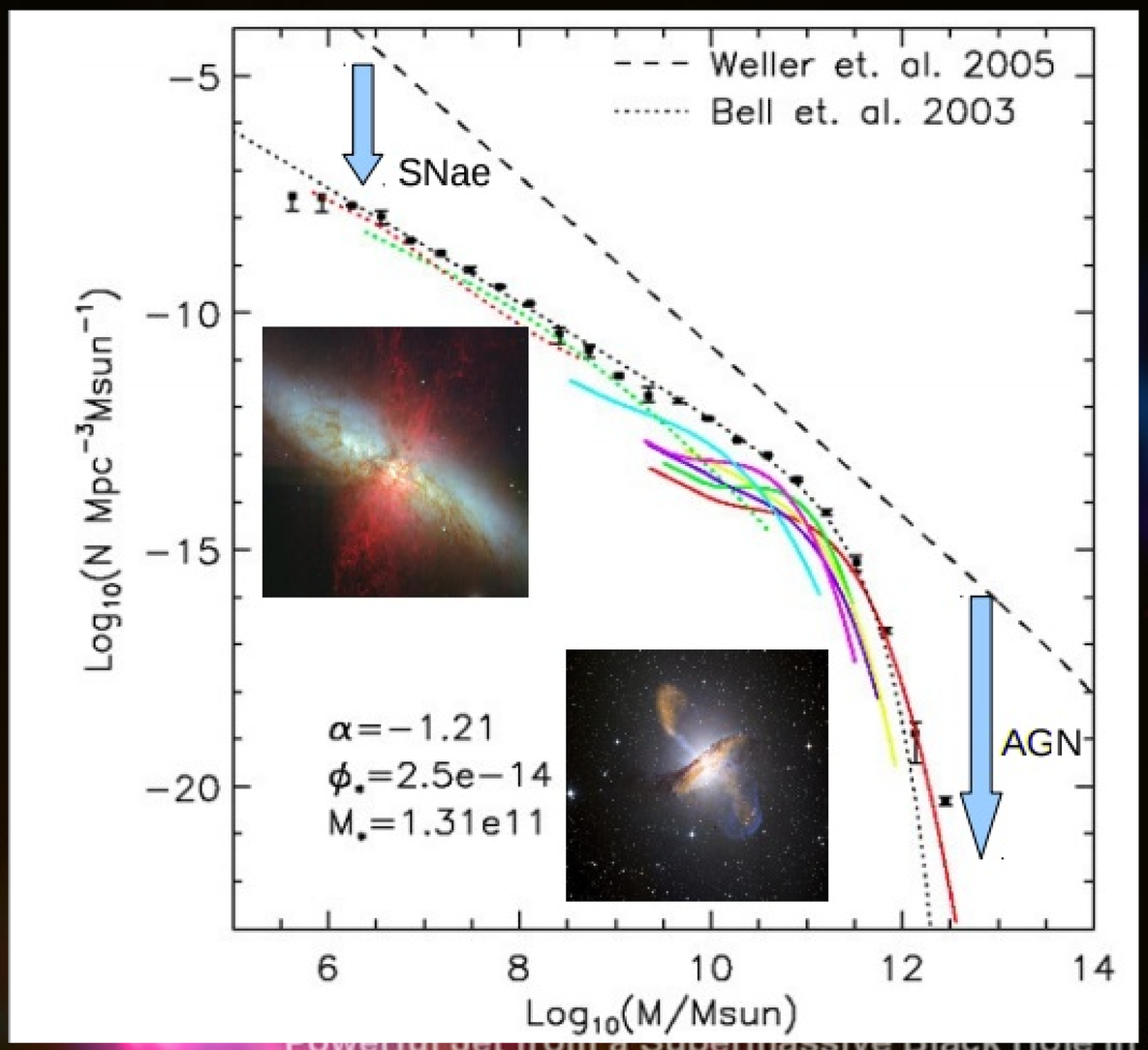
Curtis & Sijacki., 2015

- High resolution → <u>backflows</u> from the large-scale jet are resolved



Require HPC AND HC (Human Capital)

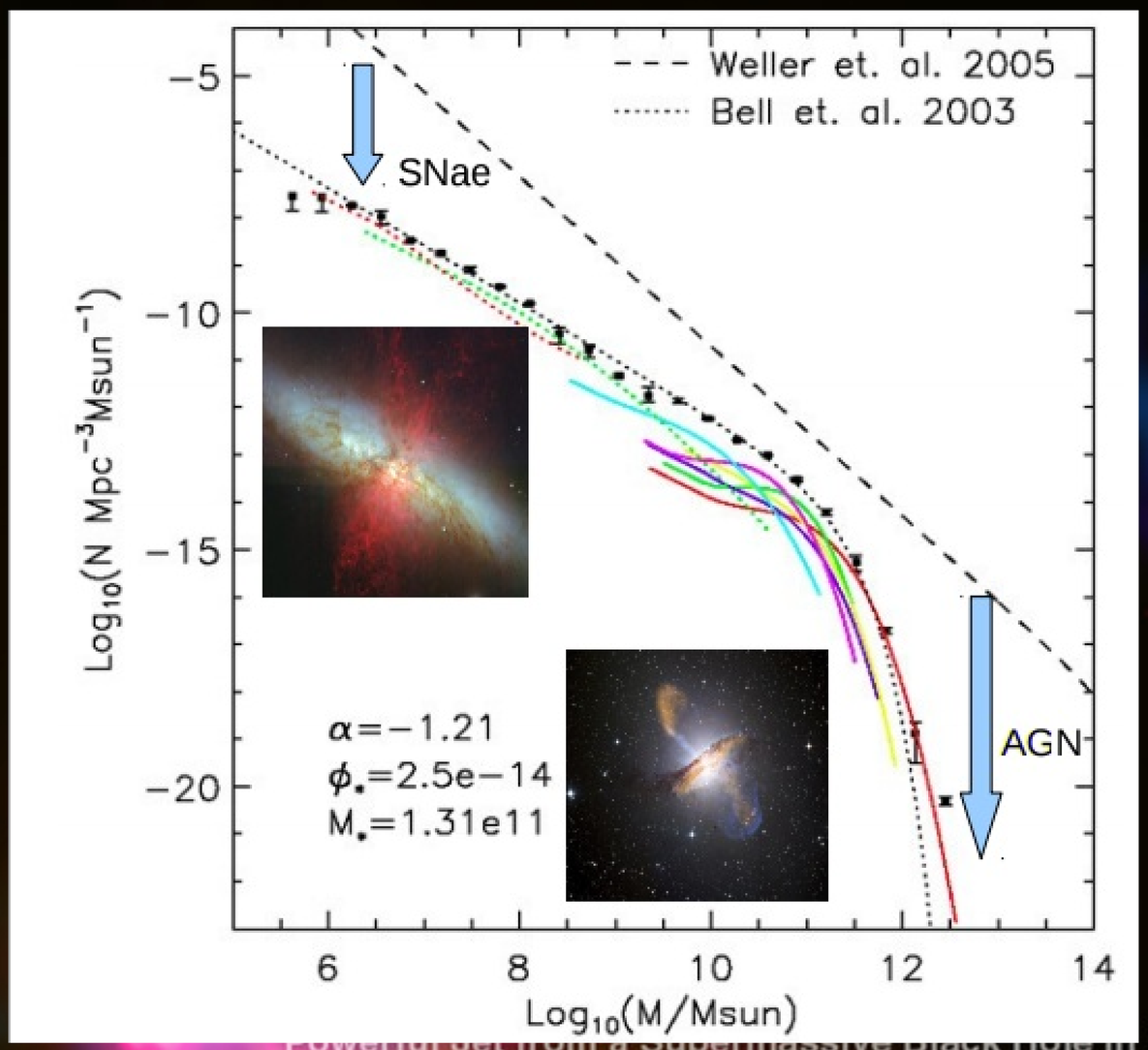
Negative AGN feedback ("*quenching*")



Sijacki, 2015 [from Read & Trentham, 2005]

Negative AGN feedback ("*quenching*")

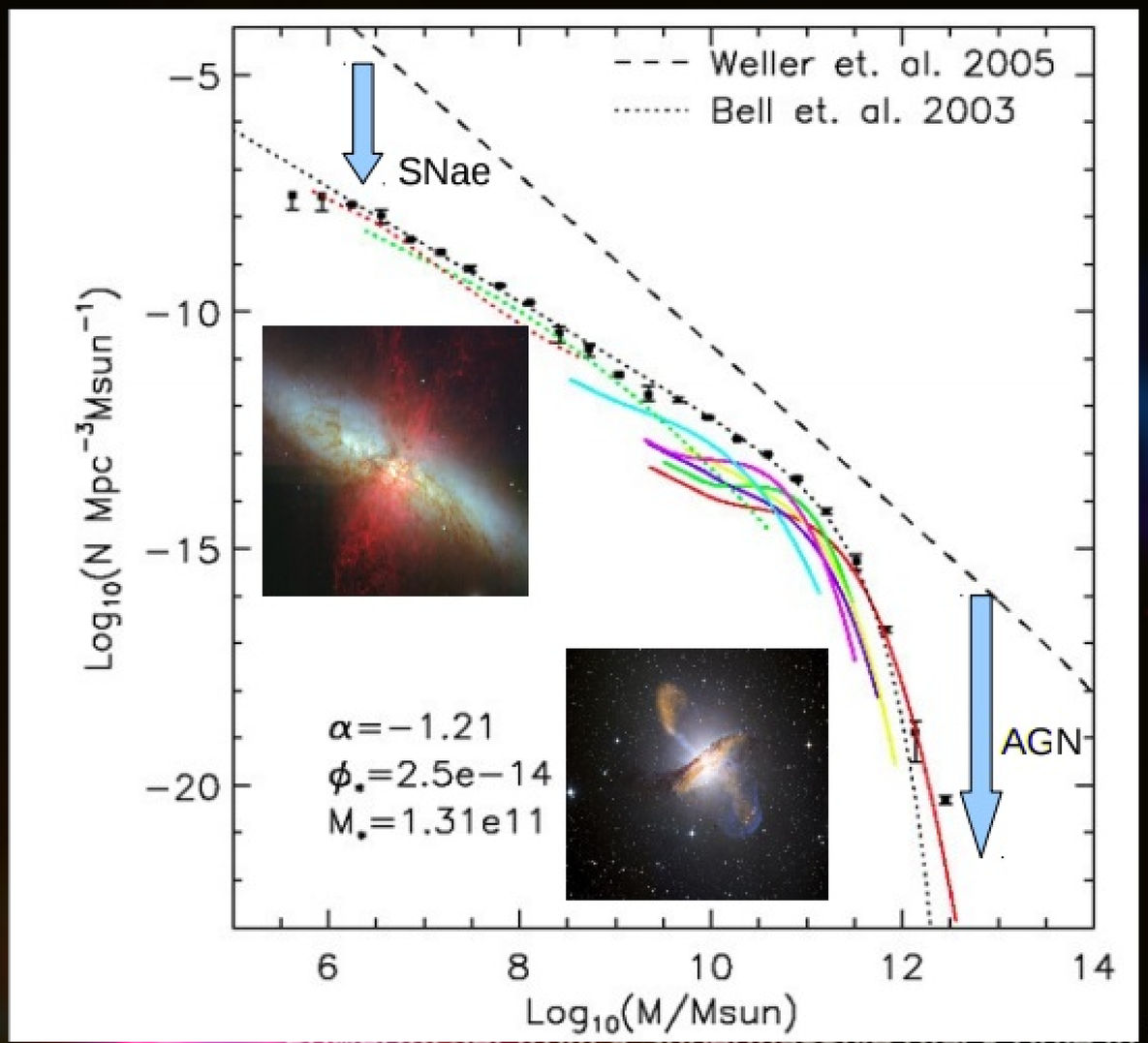
- Baryonic galaxies MF: features from SN/wind and AGN feedback ($M_* \lesssim$ few $10^{11}\,M_\odot$ [z~0])



Sijacki, 2015 [from Read & Trentham, 2005]

Negative AGN feedback ("*quenching*")

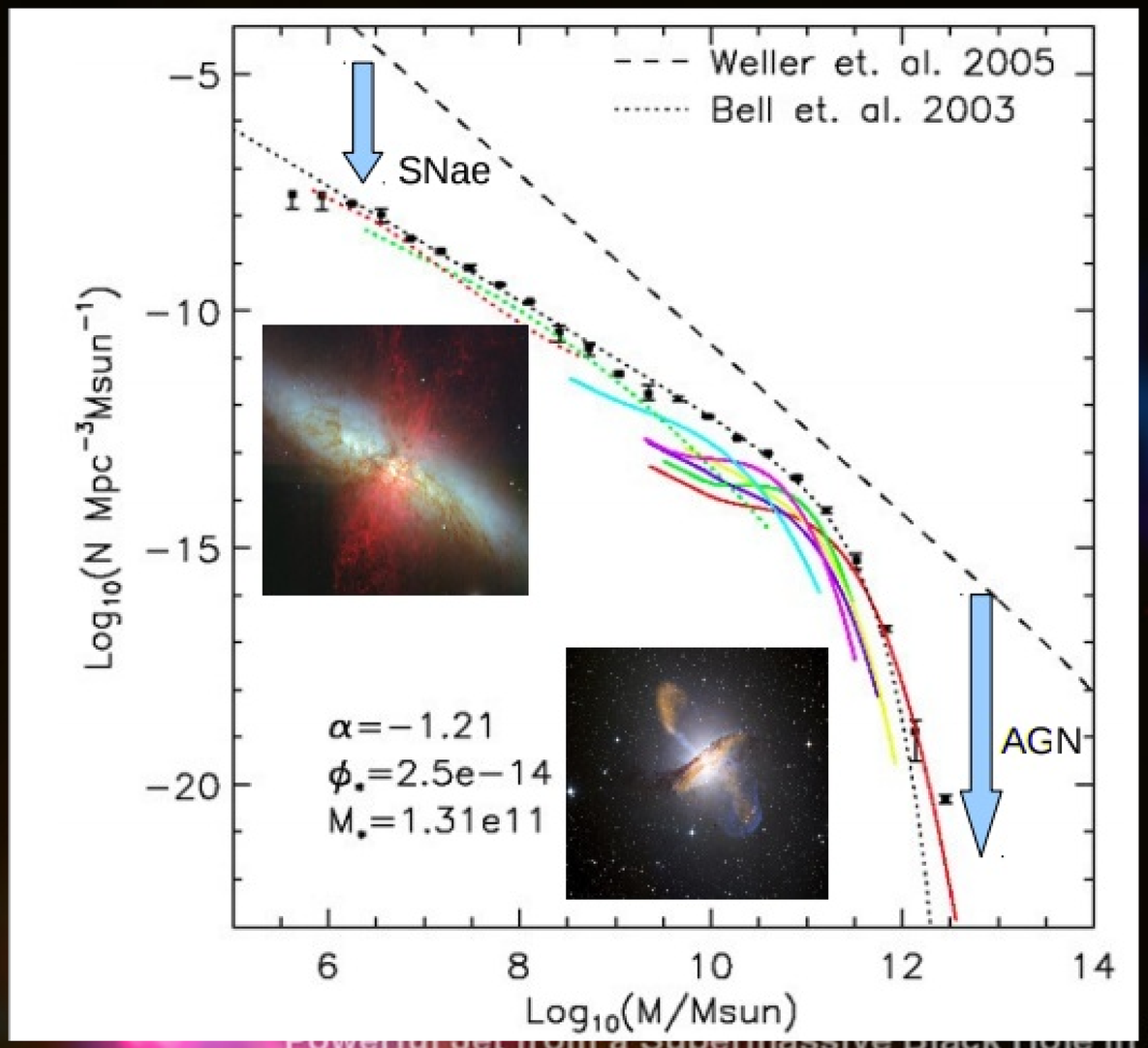
- Baryonic galaxies MF: features from SN/wind and AGN feedback ($M_* \lesssim$ few $10^{11}\,M_\odot$ [z~0])
- Montecarlo (including N-body+hydro) techniques reproduce also the evolution of the MF → fine tuning of feedback parameters



Sijacki, 2015 [from Read & Trentham, 2005]

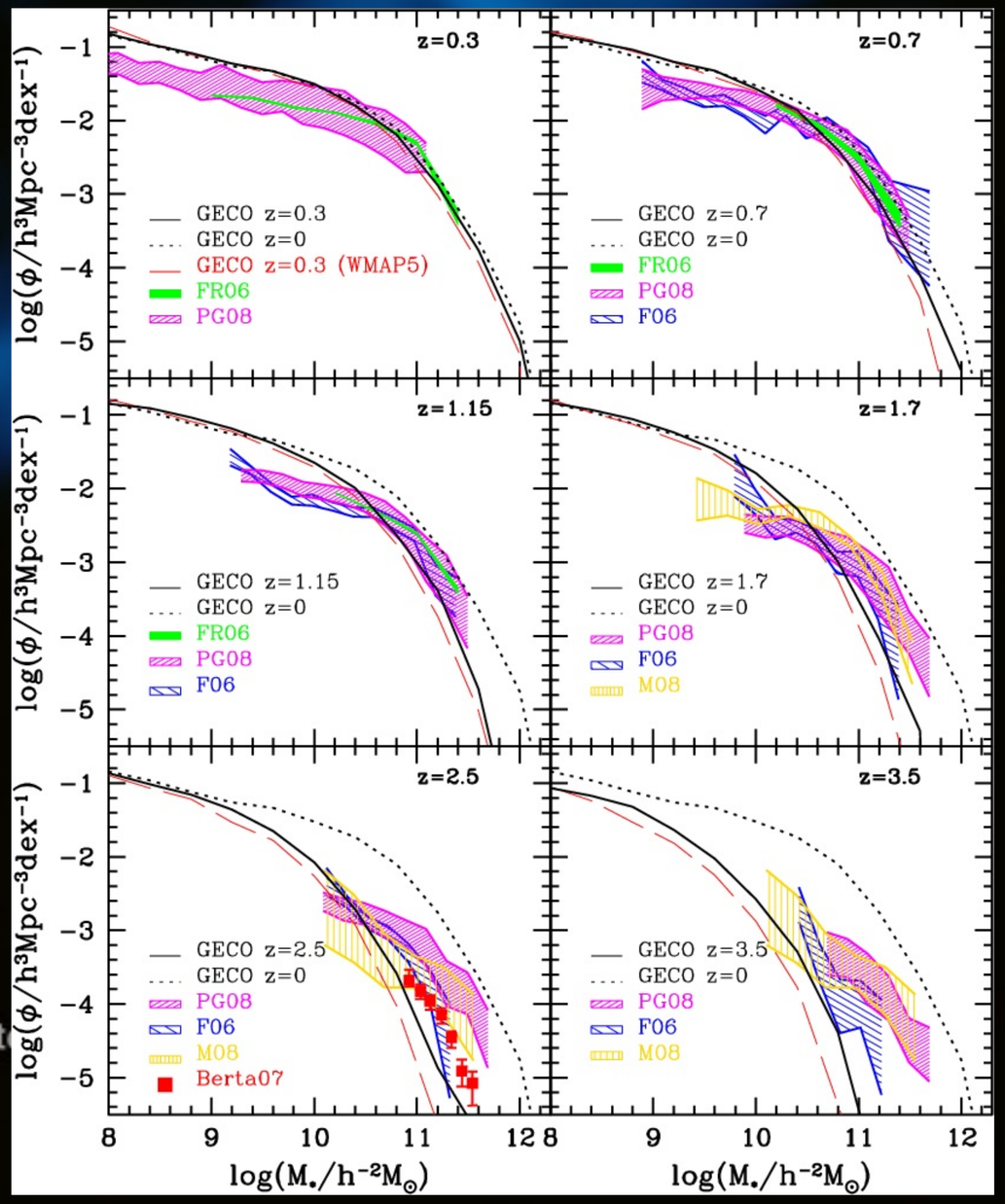
Negative AGN feedback ("*quenching*")

- Baryonic galaxies MF: features from SN/wind and AGN feedback ($M_* \lesssim$ few $10^{11}\,M_\odot$ [z~0])



Sijacki, 2015 [from Read & Trentham, 2005]

- Montecarlo (including N-body+hydro) techniques reproduce also the evolution of the MF → fine tuning of feedback parameters



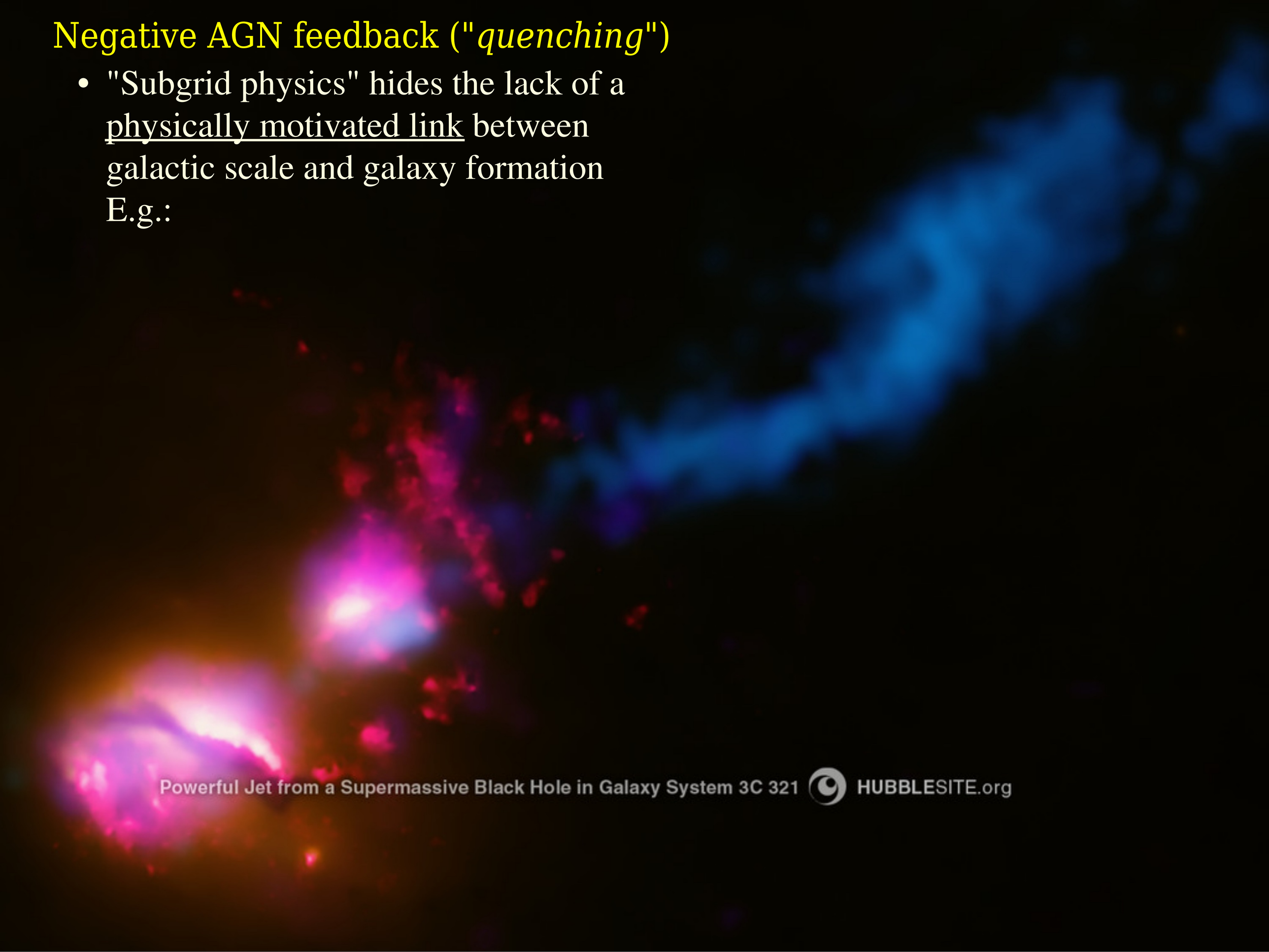
GECO: Ricciardelli & Franceschini, 2010

Negative AGN feedback ("*quenching*")

Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

Negative AGN feedback ("*quenching*")

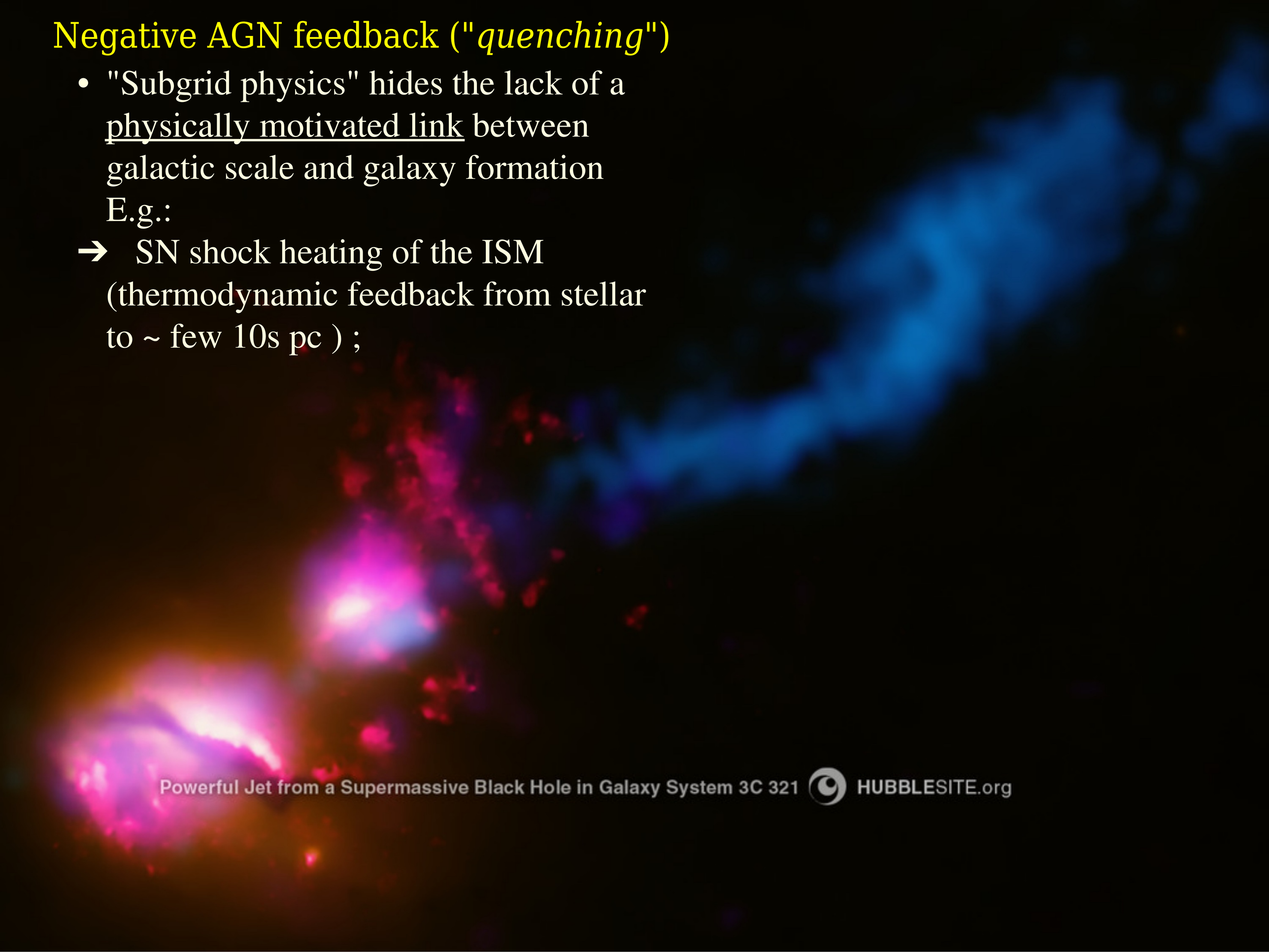
- "Subgrid physics" hides the lack of a physically motivated link between galactic scale and galaxy formation
  E.g.:



Negative AGN feedback ("*quenching*")

- "Subgrid physics" hides the lack of a <u>physically motivated link</u> between galactic scale and galaxy formation
  E.g.:

➔ SN shock heating of the ISM (thermodynamic feedback from stellar to ~ few 10s pc) ;

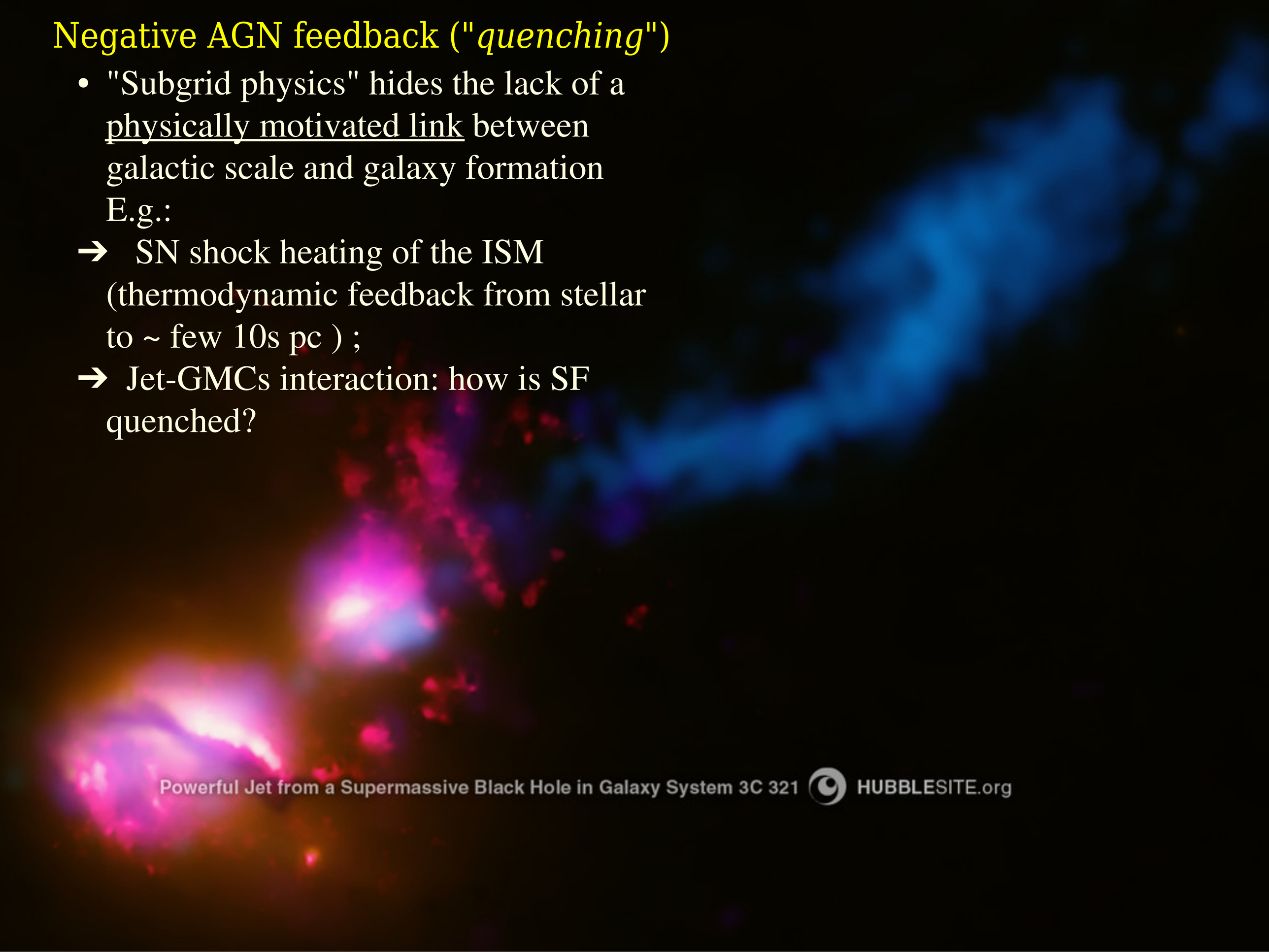


Negative AGN feedback ("*quenching*")

- "Subgrid physics" hides the lack of a physically motivated link between galactic scale and galaxy formation
  E.g.:

➔ SN shock heating of the ISM (thermodynamic feedback from stellar to ~ few 10s pc) ;

➔ Jet-GMCs interaction: how is SF quenched?



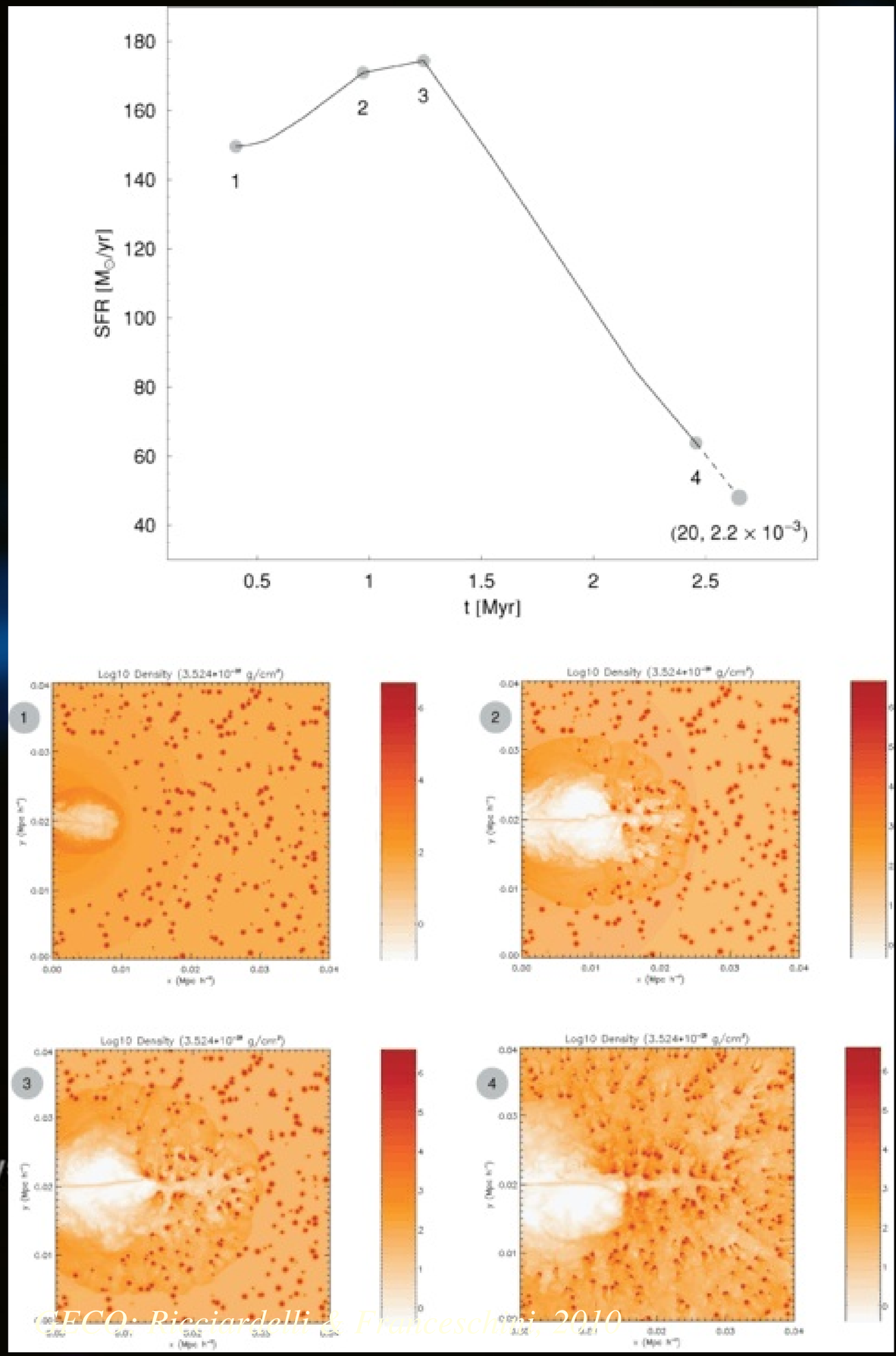
Negative AGN feedback ("*quenching*")

- "Subgrid physics" hides the lack of a <u>physically motivated link</u> between galactic scale and galaxy formation E.g.:

➔ SN shock heating of the ISM (thermodynamic feedback from stellar to ~ few 10s pc) ;

➔ Jet-GMCs interaction: how is SF quenched?

Tortora et al. (2009): <u>positive</u> feedback precedes quenching



Tortora, V.-A.-D. et al., 2009

©ECO: Ricciardelli & Franceschini 2010

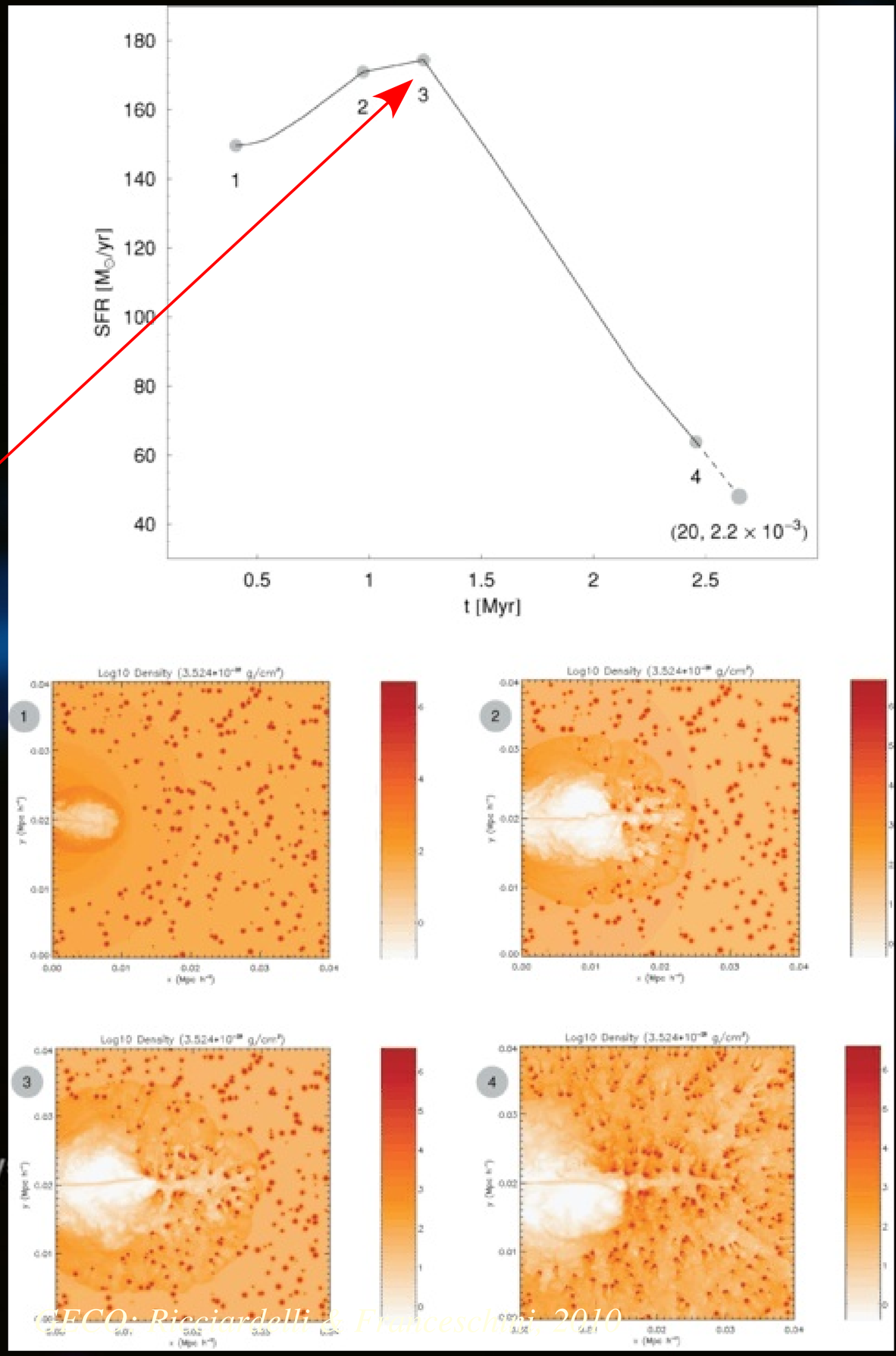
Negative AGN feedback ("*quenching*")

- "Subgrid physics" hides the lack of a <u>physically motivated link</u> between galactic scale and galaxy formation E.g.:

➔ SN shock heating of the ISM (thermodynamic feedback from stellar to ~ few 10s pc) ;

➔ Jet-GMCs interaction: how is SF quenched?

Tortora et al. (2009): <u>positive</u> feedback precedes quenching



Tortora, V.-A.-D. et al., 2009

©ECO: Ricciardelli & Franceschini 2010

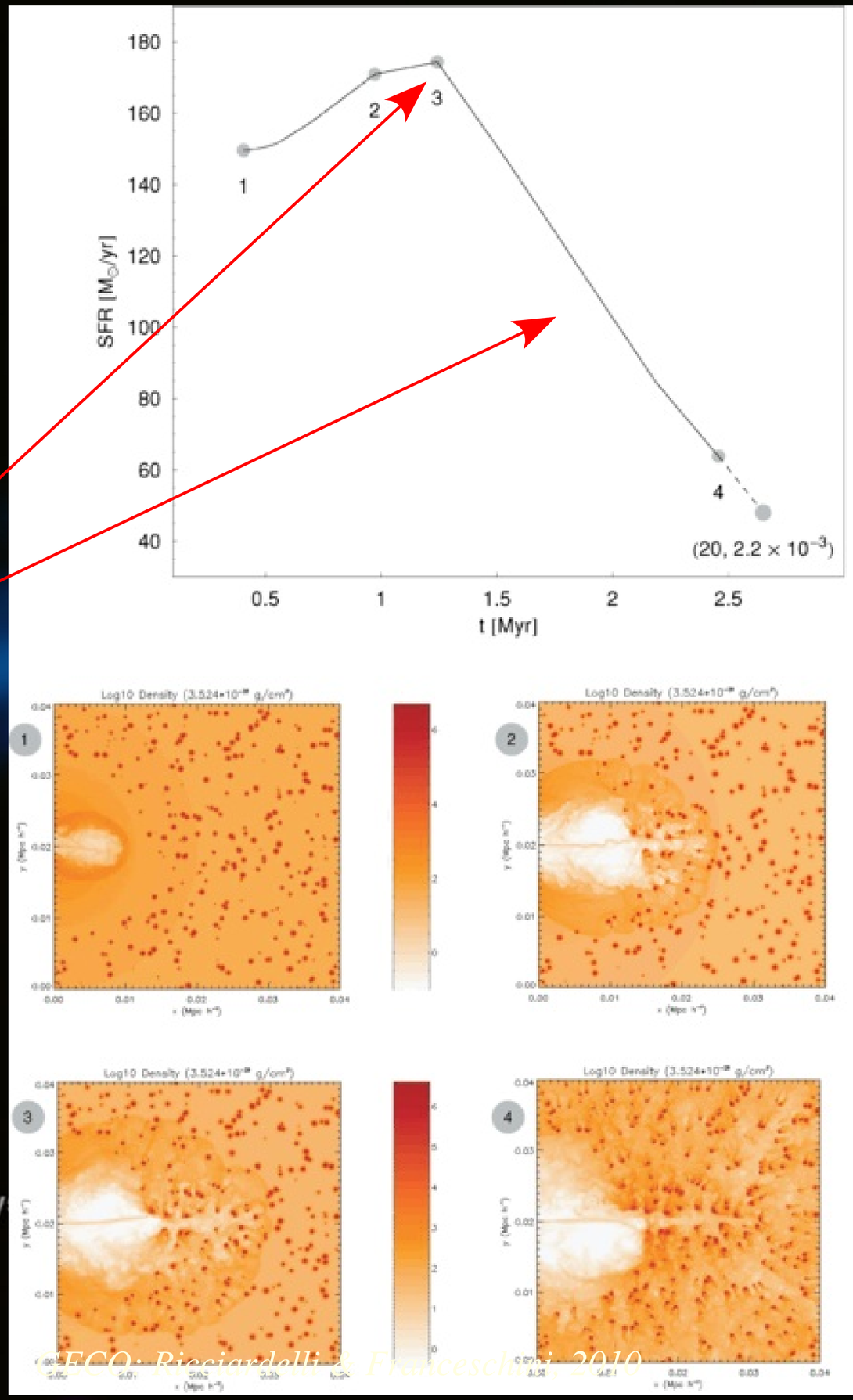
Negative AGN feedback ("*quenching*")

- "Subgrid physics" hides the lack of a <u>physically motivated link</u> between galactic scale and galaxy formation E.g.:

➔ SN shock heating of the ISM (thermodynamic feedback from stellar to ~ few 10s pc) ;

➔ Jet-GMCs interaction: how is SF quenched?

Tortora et al. (2009): <u>positive</u> feedback precedes quenching



Tortora, V.-A.-D. et al., 2009

©ECO: Ricciardelli & Franceschini 2010

Negative AGN feedback ("*quenching*")

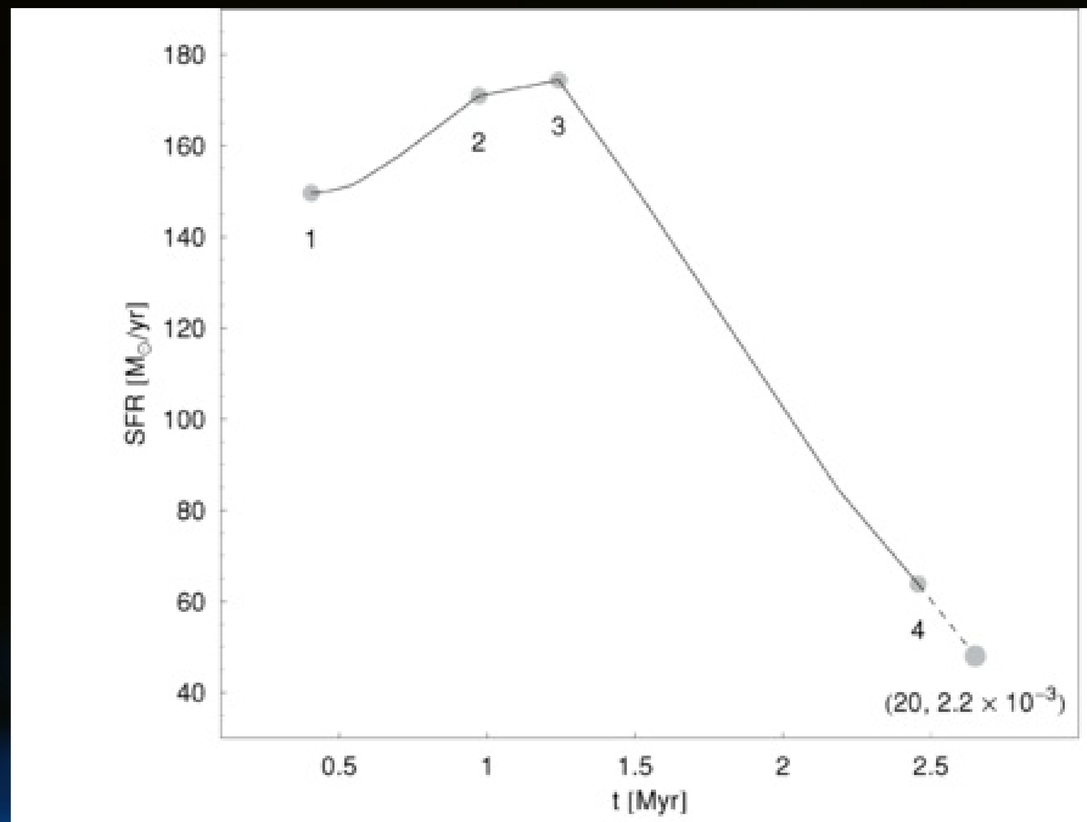
- "Subgrid physics" hides the lack of a <u>physically motivated link</u> between galactic scale and galaxy formation
  E.g.:

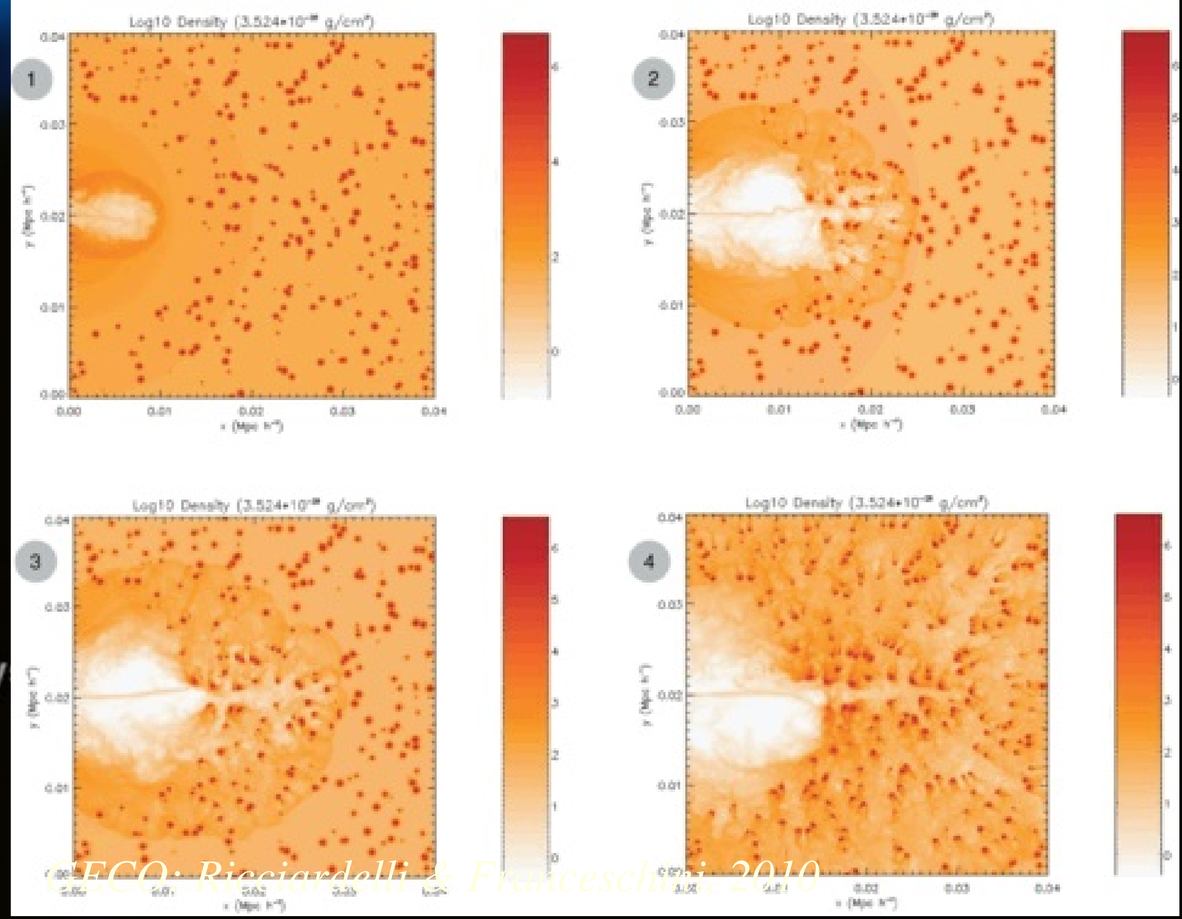
➔ SN shock heating of the ISM (thermodynamic feedback from stellar to ~ few 10s pc) ;

➔ Jet-GMCs interaction: how is SF quenched?

Tortora et al. (2009): <u>positive</u> feedback precedes quenching
Jet-induced SF is localized to the rims of the cocoon's lobes

Powerful Jet from a Supermassive Black Hole in Galaxy Sy





Tortora, V.-A.-D. et al., 2009

©ECO: Ricciardelli & Franceschini 2010

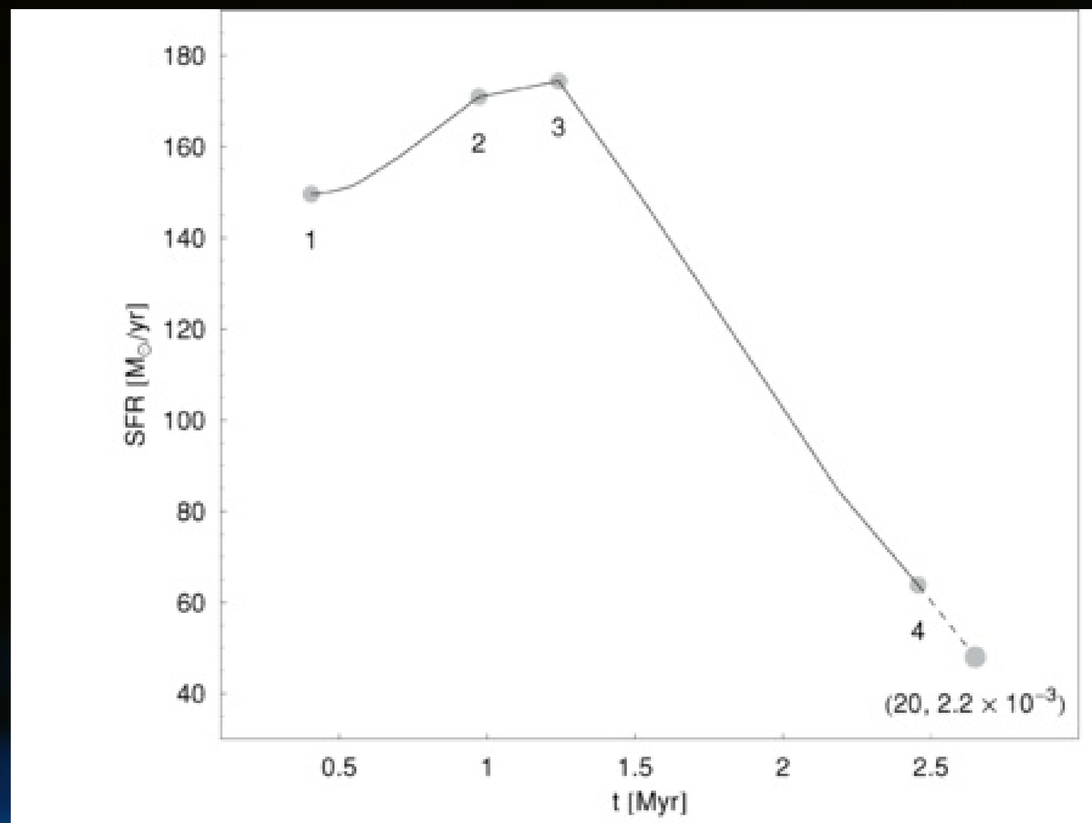
Negative AGN feedback ("*quenching*")

- "Subgrid physics" hides the lack of a <u>physically motivated link</u> between galactic scale and galaxy formation
  E.g.:

➔ SN shock heating of the ISM (thermodynamic feedback from stellar to ~ few 10s pc) ;

➔ Jet-GMCs interaction: how is SF quenched?

Tortora et al. (2009): <u>positive</u> feedback precedes quenching
Jet-induced SF is localized to the rims of the cocoon's lobes



Tortora, V.-A.-D. et al., 2009

©ESO: Ricciardelli & Franceschini, 2010

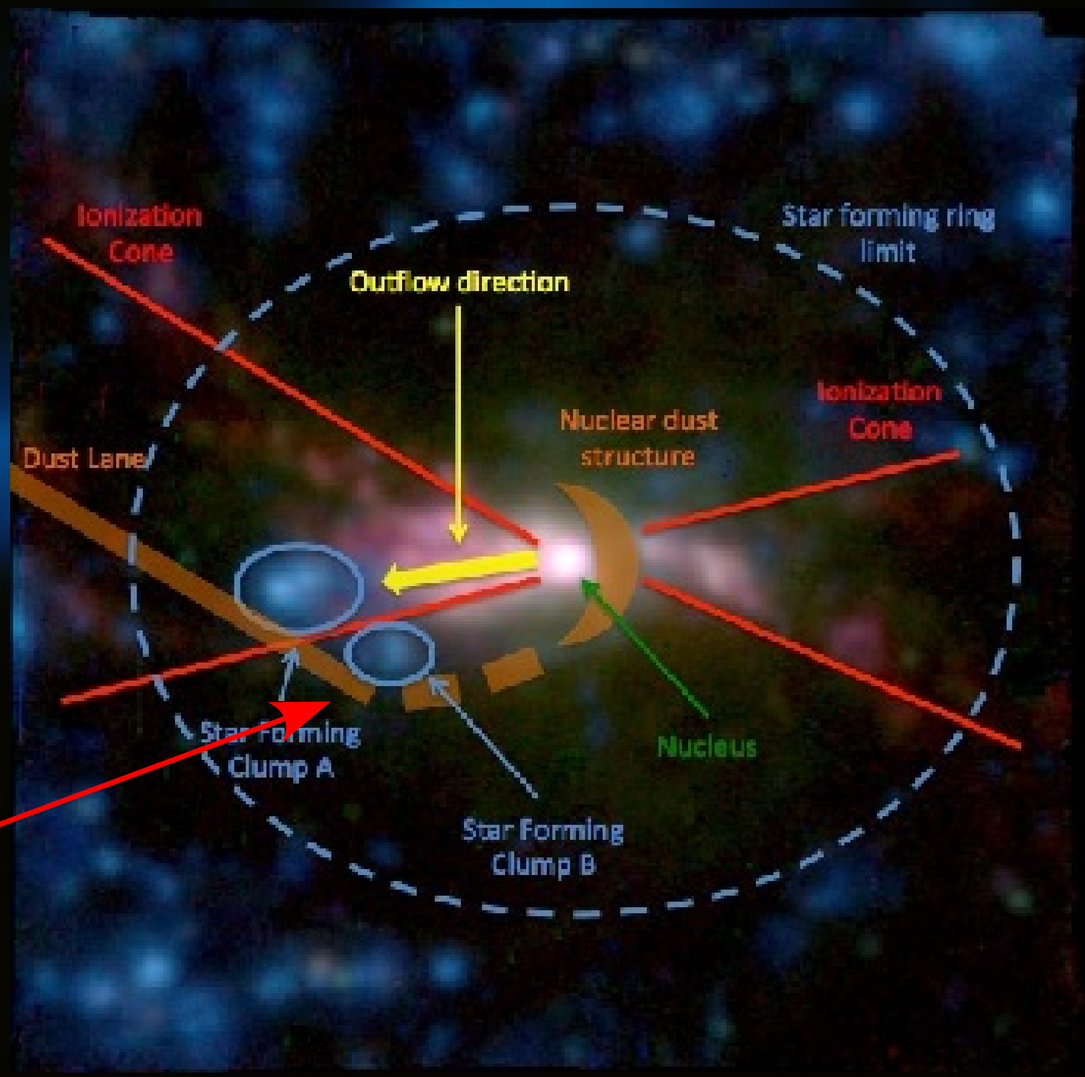
Negative AGN feedback ("*quenching*")

- "Subgrid physics" hides the lack of a <u>physically motivated link</u> between galactic scale and galaxy formation
  E.g.:

➔ SN shock heating of the ISM (thermodynamic feedback from stellar to ~ few 10s pc) ;

➔ Jet-GMCs interaction: how is SF quenched?

Tortora et al. (2009): <u>positive</u> feedback precedes quenching
Jet-induced SF is localized to the rims of the cocoon's lobes
....compare with Carniani et al., 2016; Cresci et al, 2015



Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

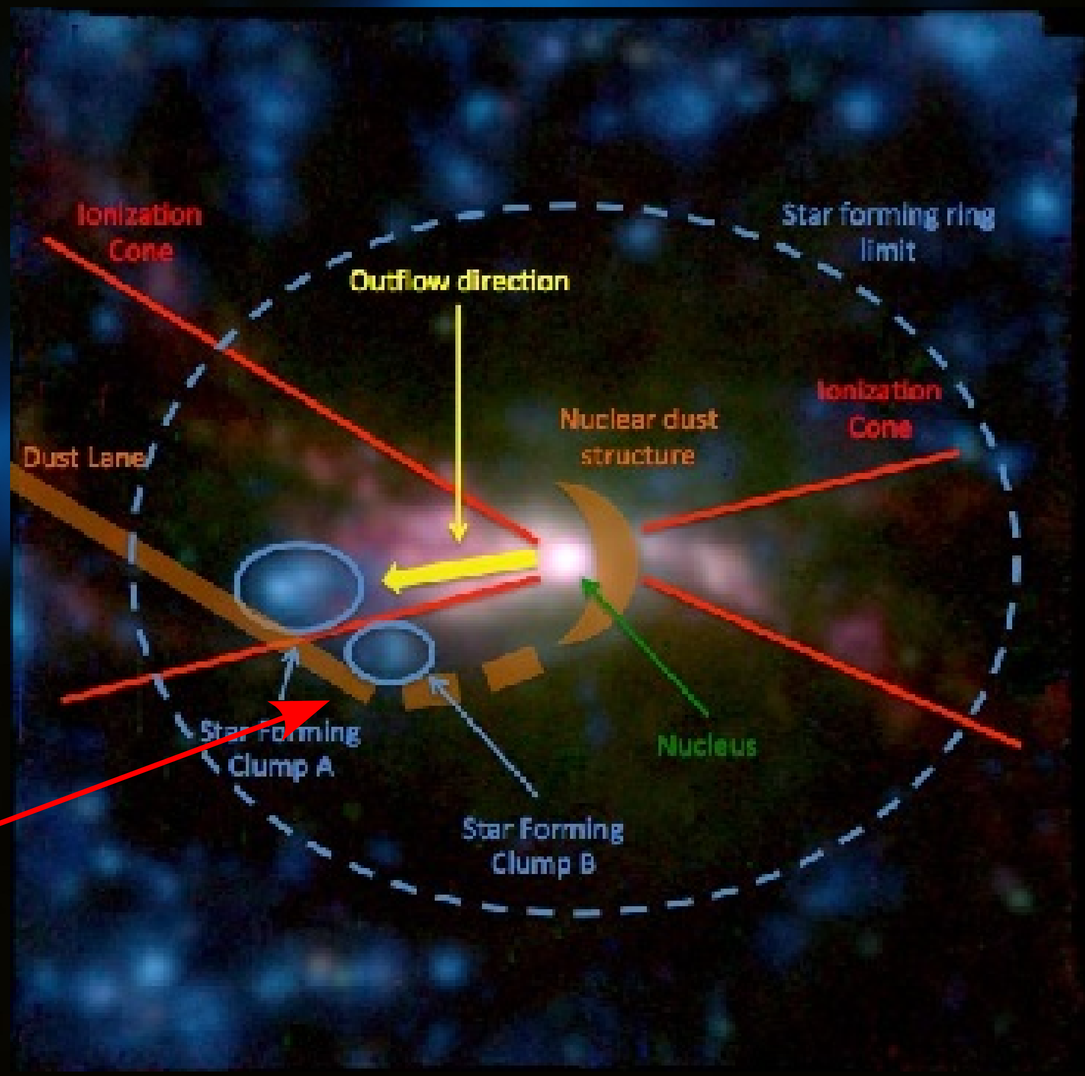
Negative AGN feedback ("*quenching*")

- "Subgrid physics" hides the lack of a <u>physically motivated link</u> between galactic scale and galaxy formation
  E.g.:

➔ SN shock heating of the ISM (thermodynamic feedback from stellar to ~ few 10s pc) ;

➔ Jet-GMCs interaction: how is SF quenched?

Tortora et al. (2009): <u>positive</u> feedback precedes quenching
Jet-induced SF is localized to the rims of the cocoon's lobes
....compare with Carniani et al., 2016; Cresci et al, 2015



Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

Positive feedback lasts lesser than quenching.....

Is positive feedback cosmologically relevant to galaxy stellar formation @ $z \gtrsim 2$?

Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

Is positive feedback cosmologically relevant to galaxy stellar formation @ $z \gtrsim 2$?

Only modelling (incl. *numerical experiments*) can provide us with unbiased , realistic hints.

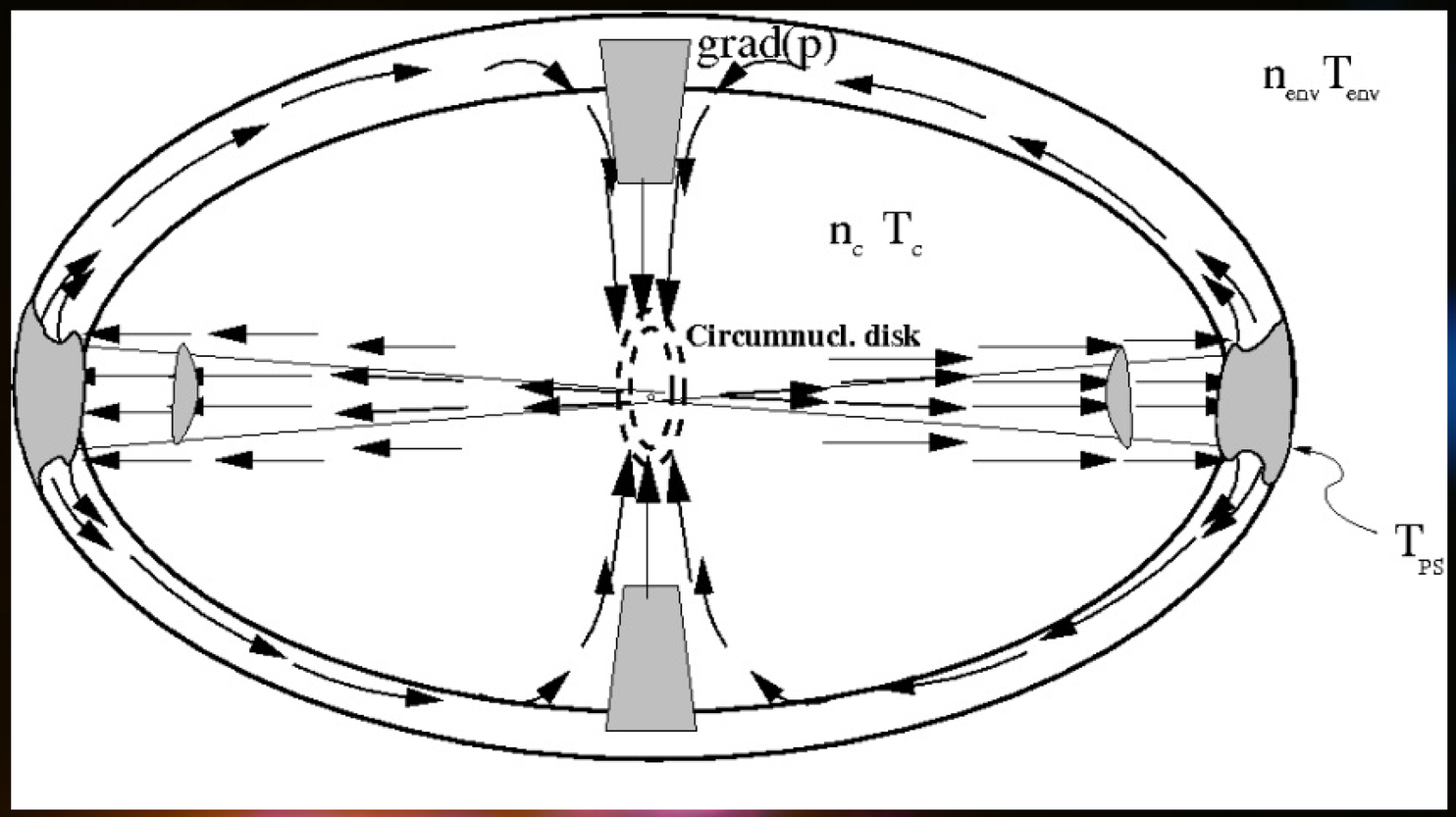
Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

Backflows within AGNs relativistic jets:
The link between large- and subparsec scale feedback

subparsec scale feedback

Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

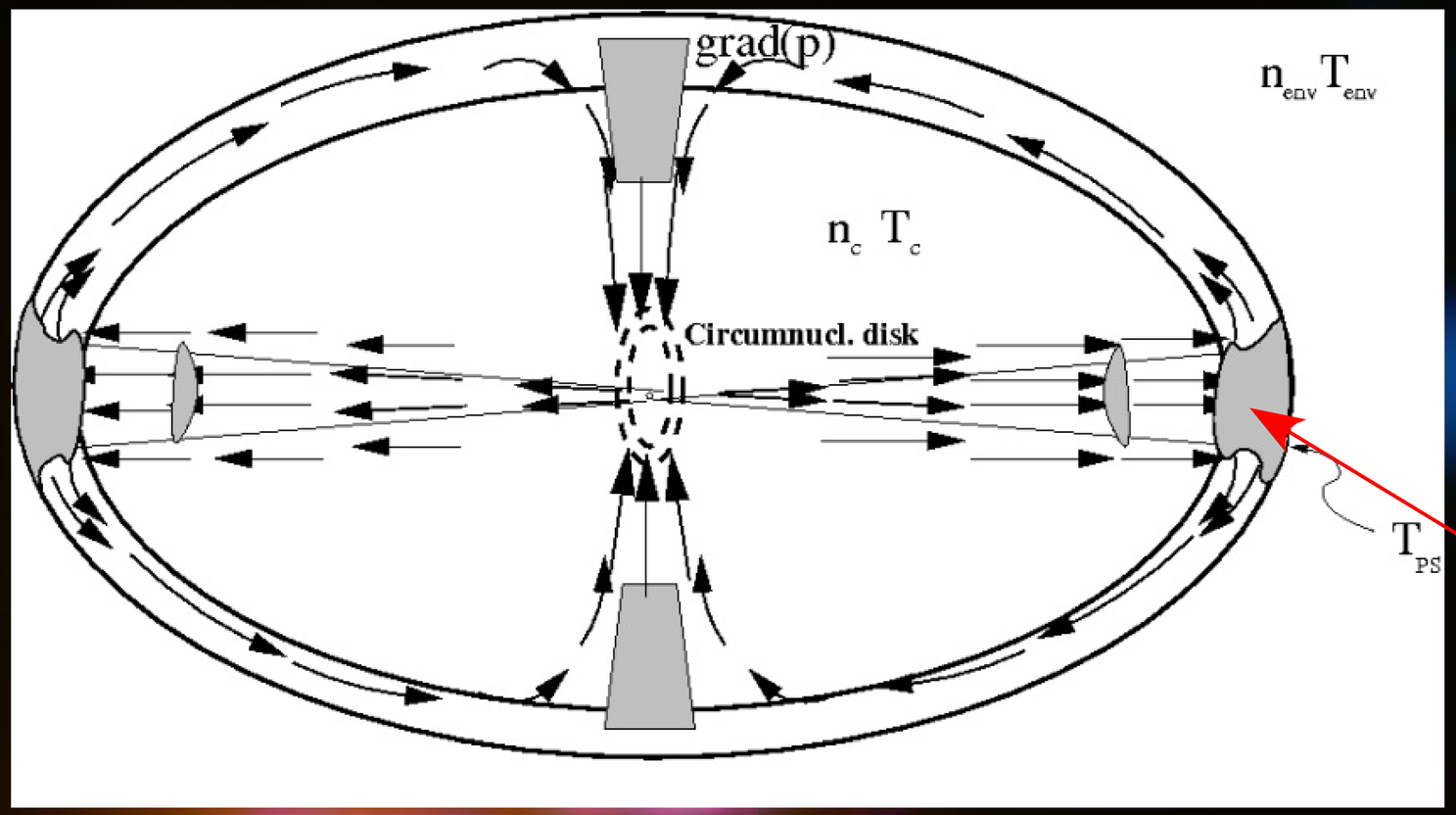
Backflows within AGNs relativistic jets:
The link between large- and subparsec scale feedback



Backflow: gas streaming opposite to and along the boundary of the cocoon.

Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

Backflows within AGNs relativistic jets:
The link between large- and subparsec scale feedback

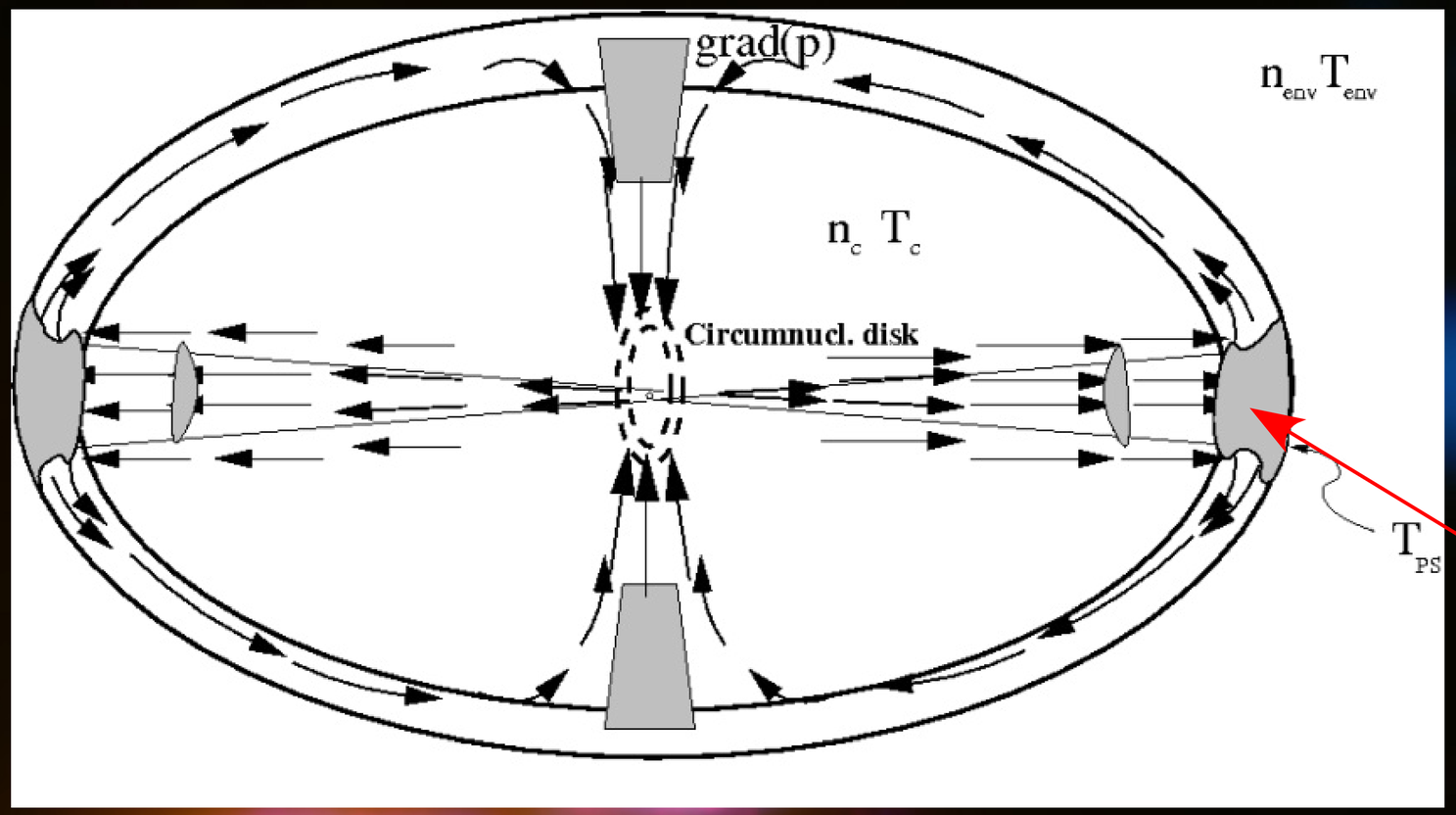


Backflow: gas streaming opposite to and along the boundary of the cocoon.

- Origin is *thermodynamical* : vorticity is produced from jet's gas crossing the shock before the Hot Spot

Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

Backflows within AGNs relativistic jets:
The link between large- and subparsec scale feedback



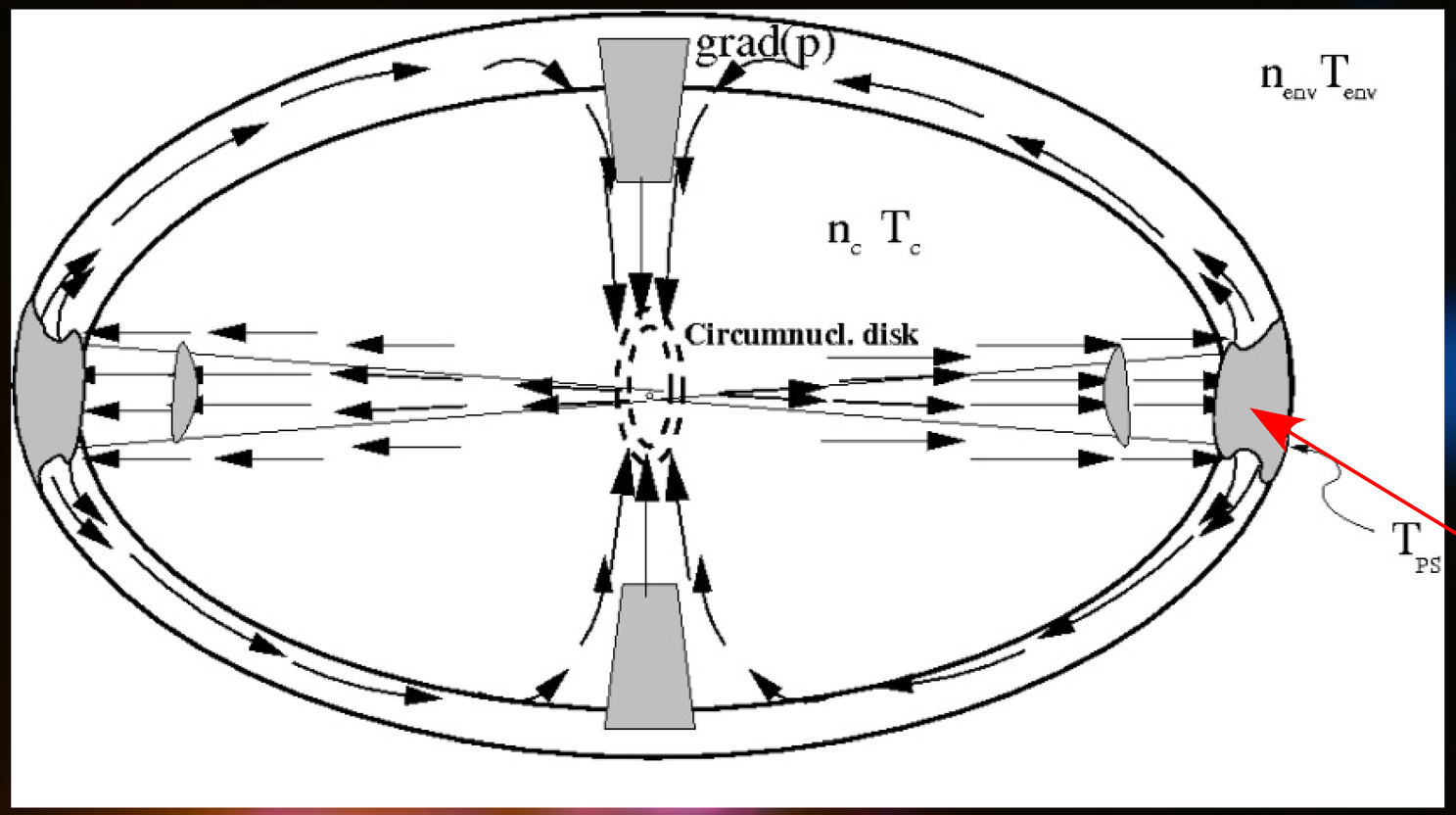
Backflow: gas streaming opposite to and along the boundary of the cocoon.

- Origin is *thermodynamical* : vorticity is produced from jet's gas crossing the shock before the Hot Spot

Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

- Backflow is *spatially coherent* and *intrinsically axisymmetric* all way down to the accretion region

Backflows within AGNs relativistic jets:
The link between large- and subparsec scale feedback



Backflow: gas streaming opposite to and along the boundary of the cocoon.

- Origin is *thermodynamical* : vorticity is produced from jet's gas crossing the shock before the Hot Spot

Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

- Backflow is *spatially coherent* and *intrinsically axisymmetric* all way down to the accretion region
- Backflow drives low-L_z gas into the SMBH accretion region

Observational evidence for backflows

Observational evidence for backflows

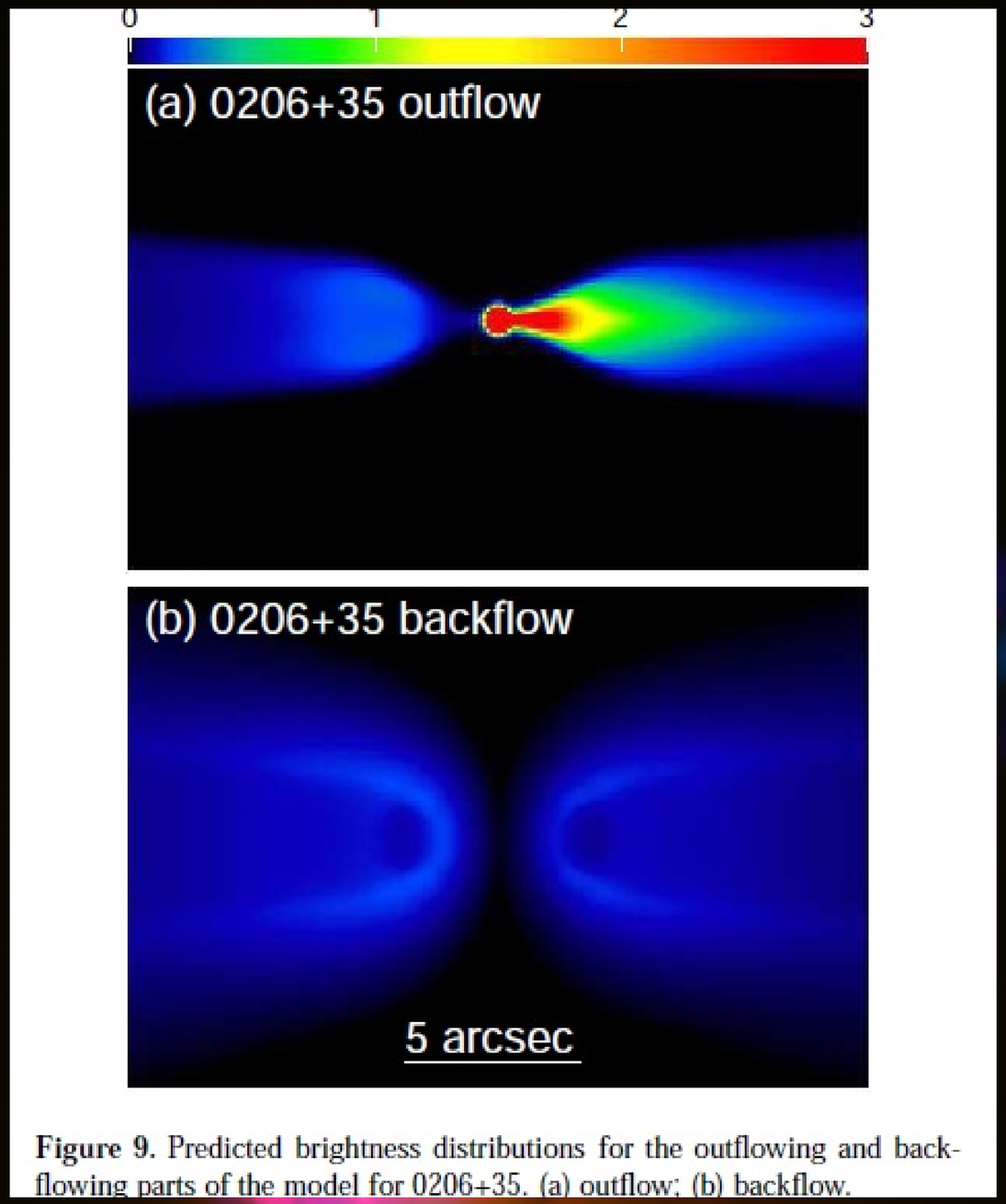


Figure 9. Predicted brightness distributions for the outflowing and backflowing parts of the model for 0206+35. (a) outflow; (b) backflow.

Laing & Bridle, 2012: FRI, mildly relativistic velocities

Observational evidence for backflows

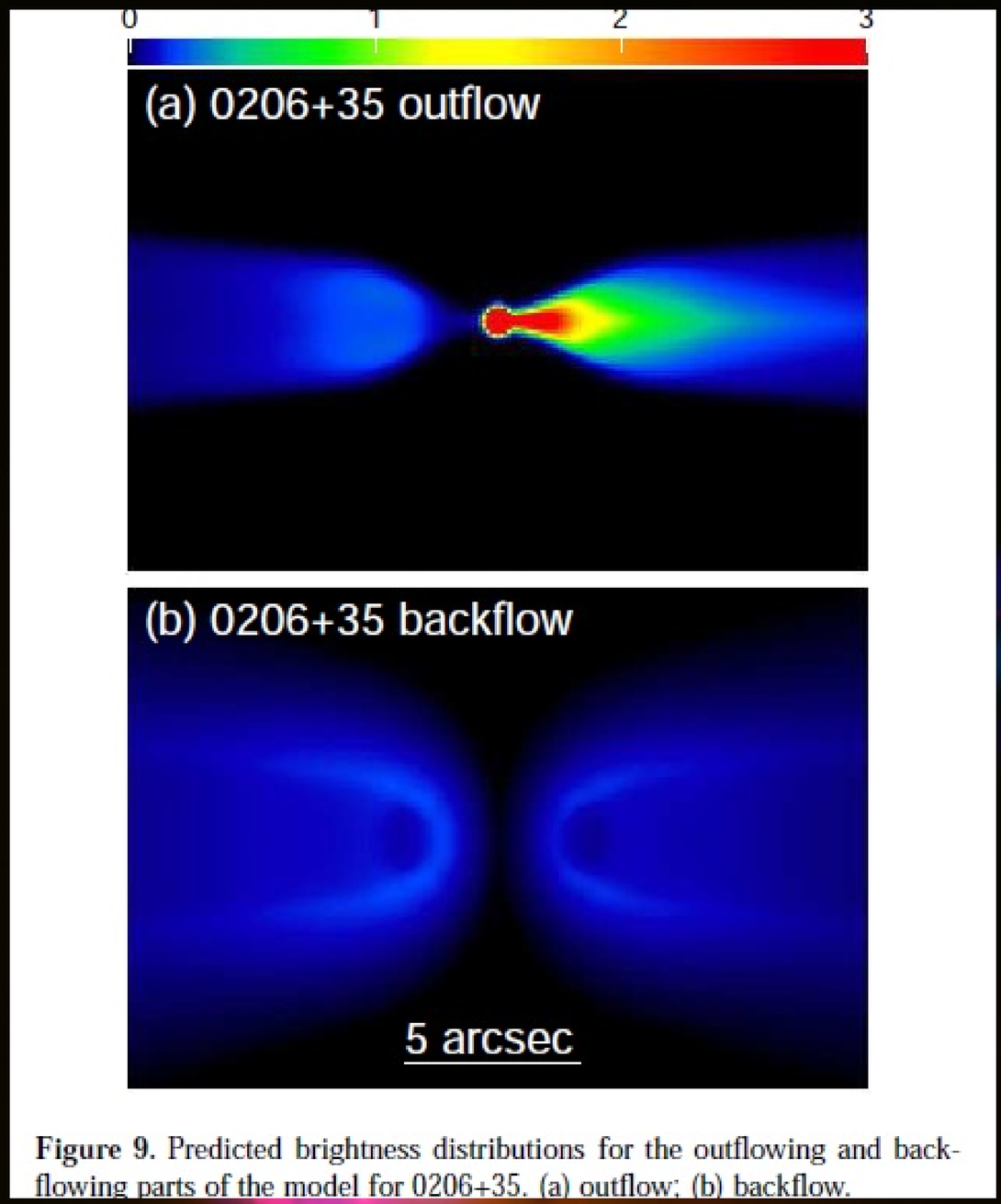
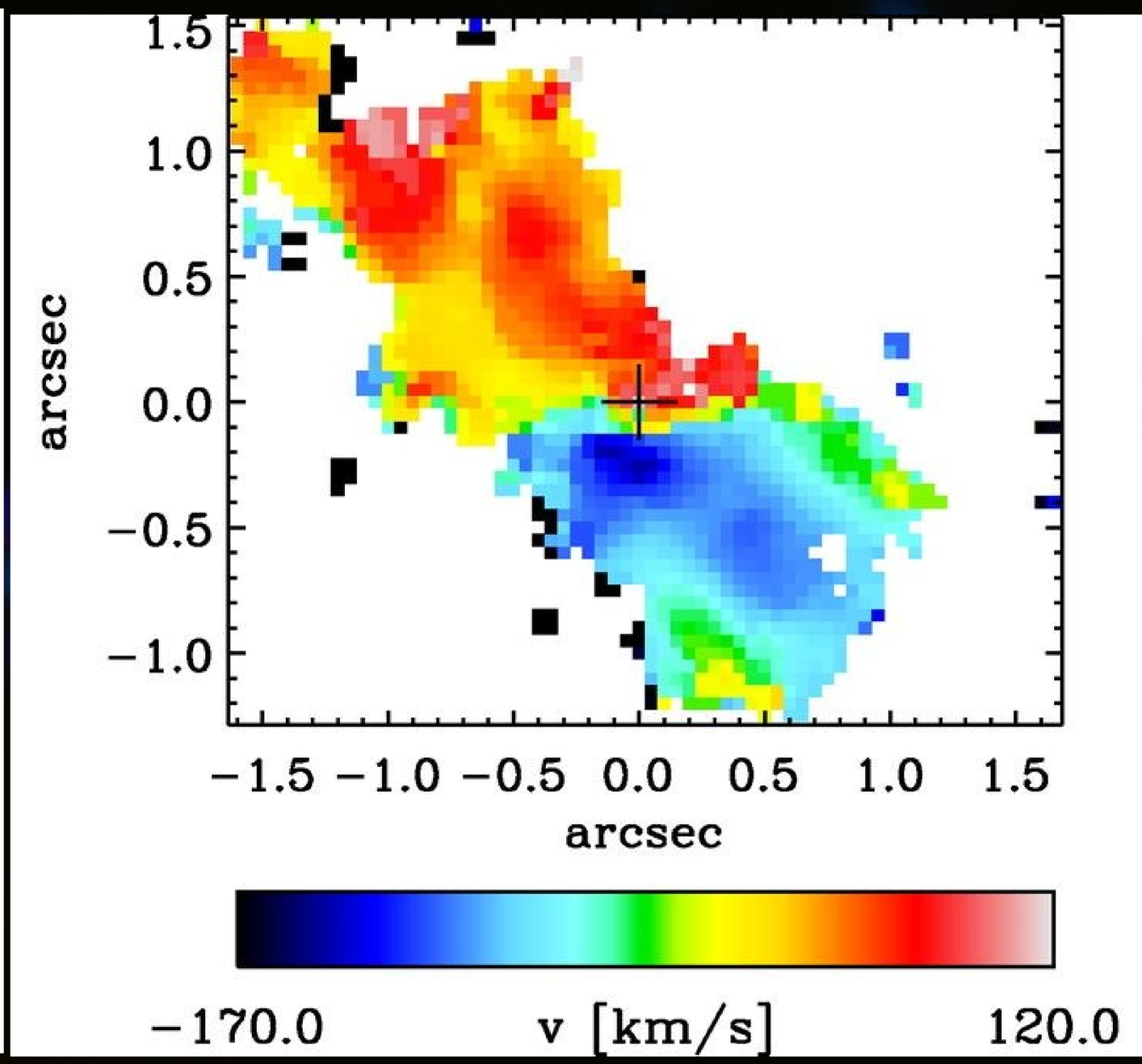


Figure 9. Predicted brightness distributions for the outflowing and backflowing parts of the model for 0206+35. (a) outflow; (b) backflow.

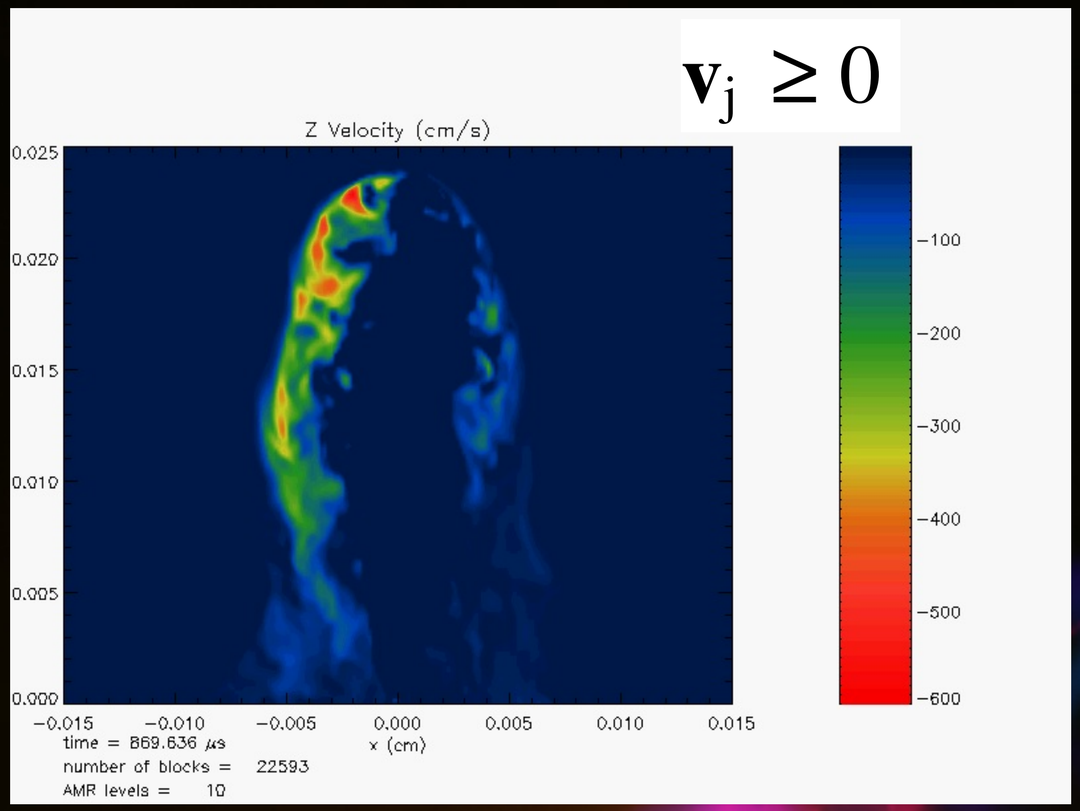
Laing & Bridle, 2012: FRI, mildly relativistic velocities

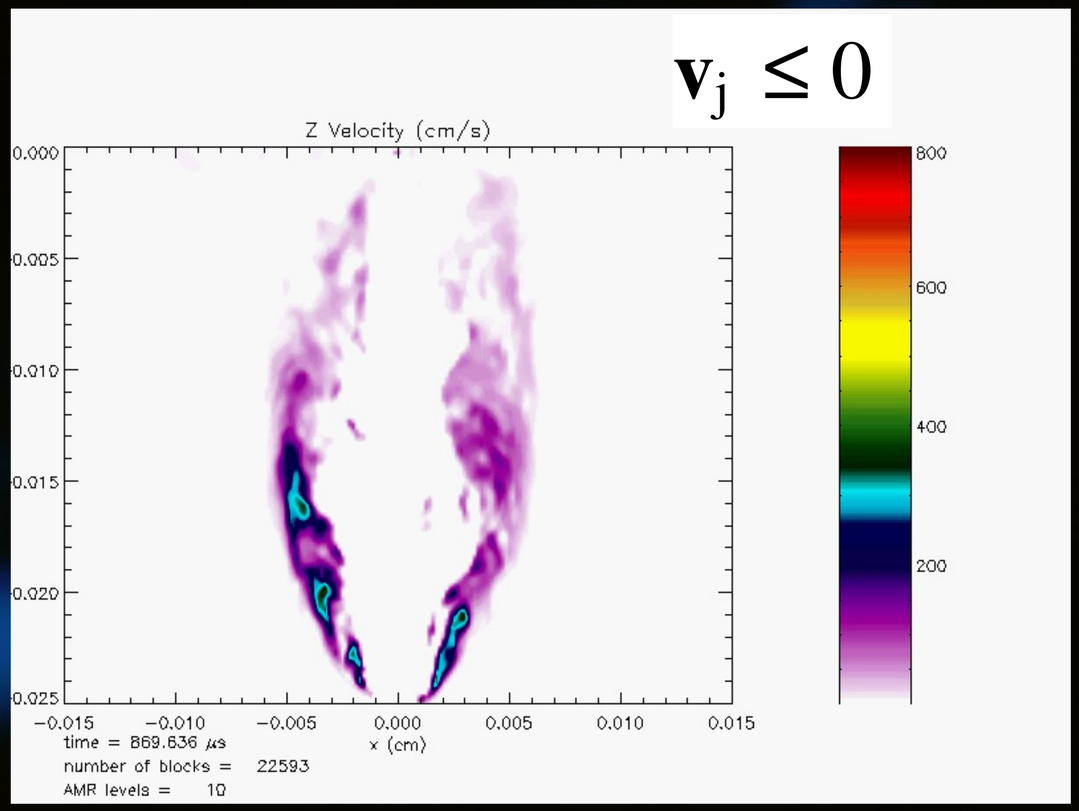


Neumayer et al, 2007: CEN A
Deeper analysis with MUSE: Hamer et al., 2015

Numerical experiments

Numerical experiments

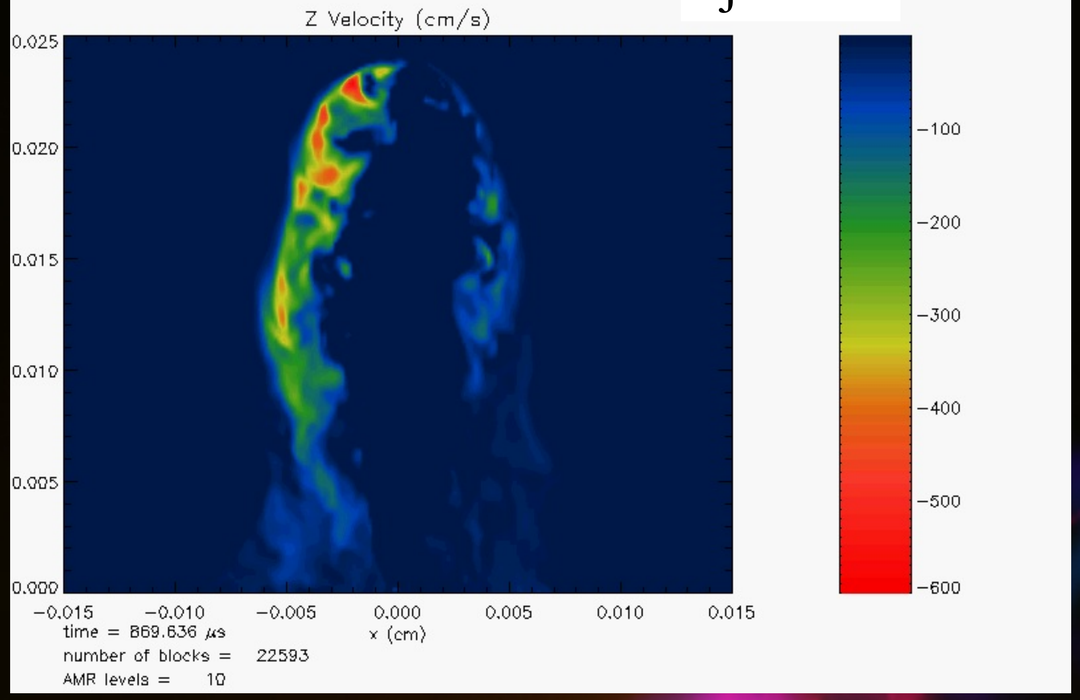




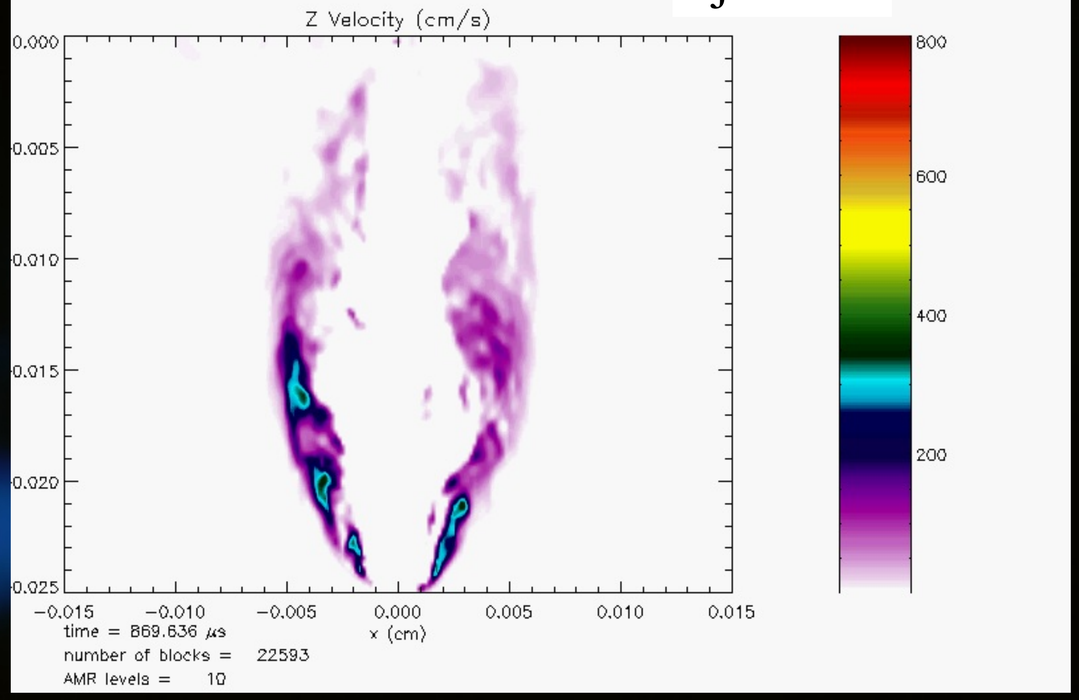
- Plotting only counterstreaming gas ($\mathbf{v}_z \cdot \mathbf{v}_j \leq 0$)

Numerical experiments

$\mathbf{v}_j \geq 0$



$\mathbf{v}_j \leq 0$

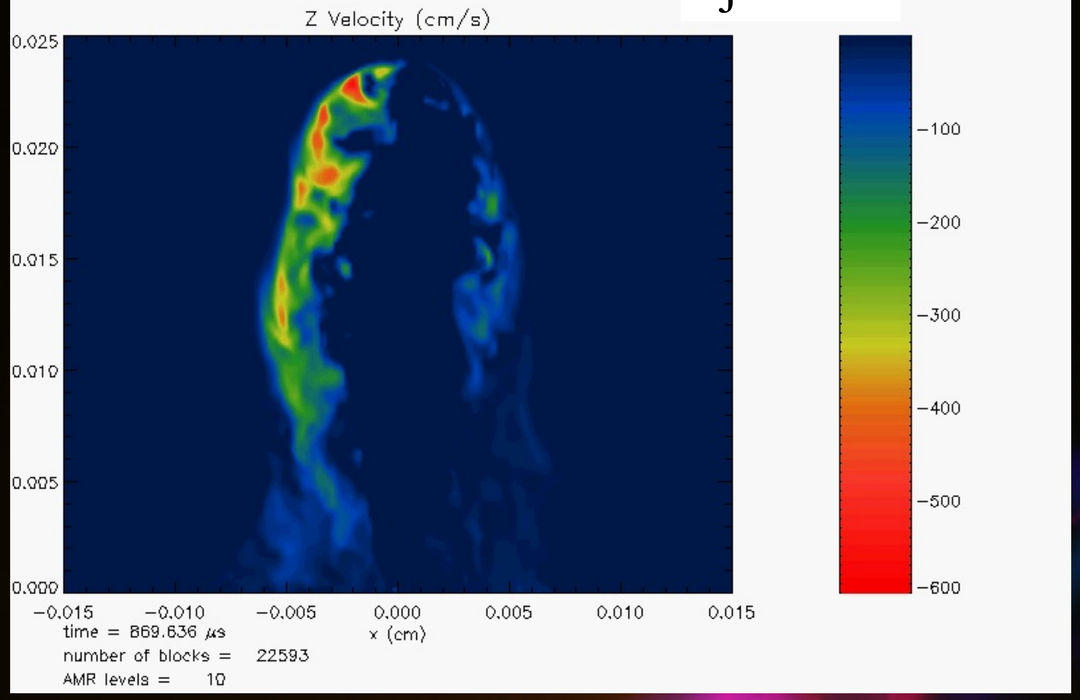


- Plotting only counterstreaming gas ($\mathbf{v}_z \cdot \mathbf{v}_j \leq 0$)
- Lessons:

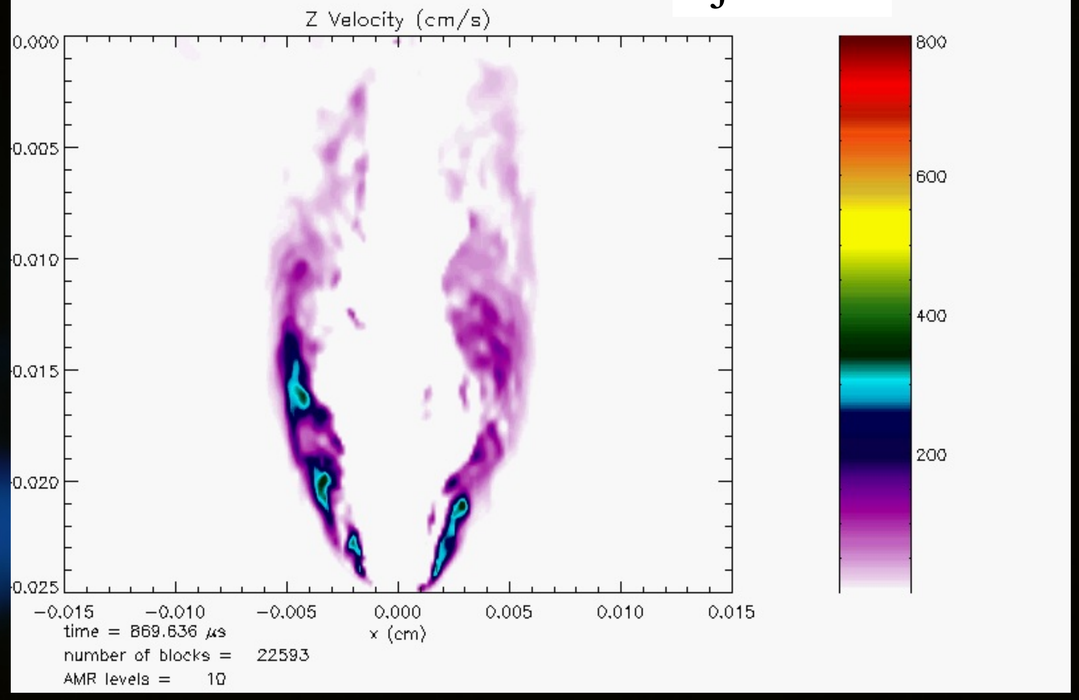
  1) Backflow develops a *large-scale pattern* (HS → meridional plane)

Numerical experiments

$\mathbf{v}_j \geq 0$



$\mathbf{v}_j \leq 0$



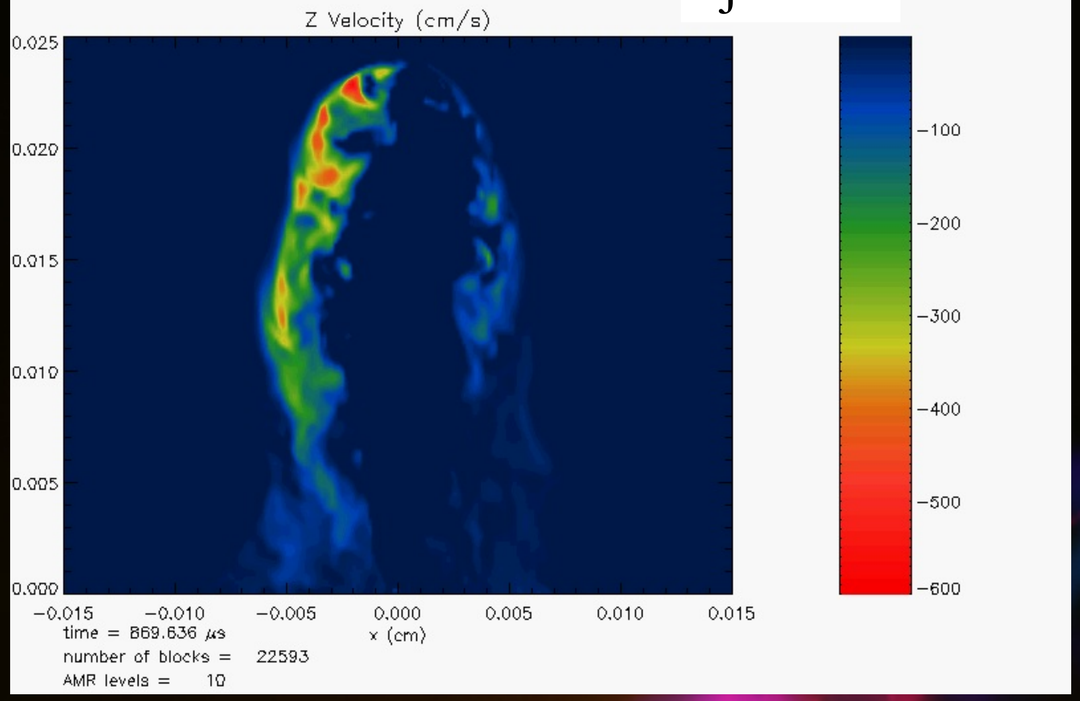
- Plotting only counterstreaming gas ($\mathbf{v}_z \cdot \mathbf{v}_j \leq 0$)
- Lessons:

  1) Backflow develops a *large-scale pattern* (HS → meridional plane)

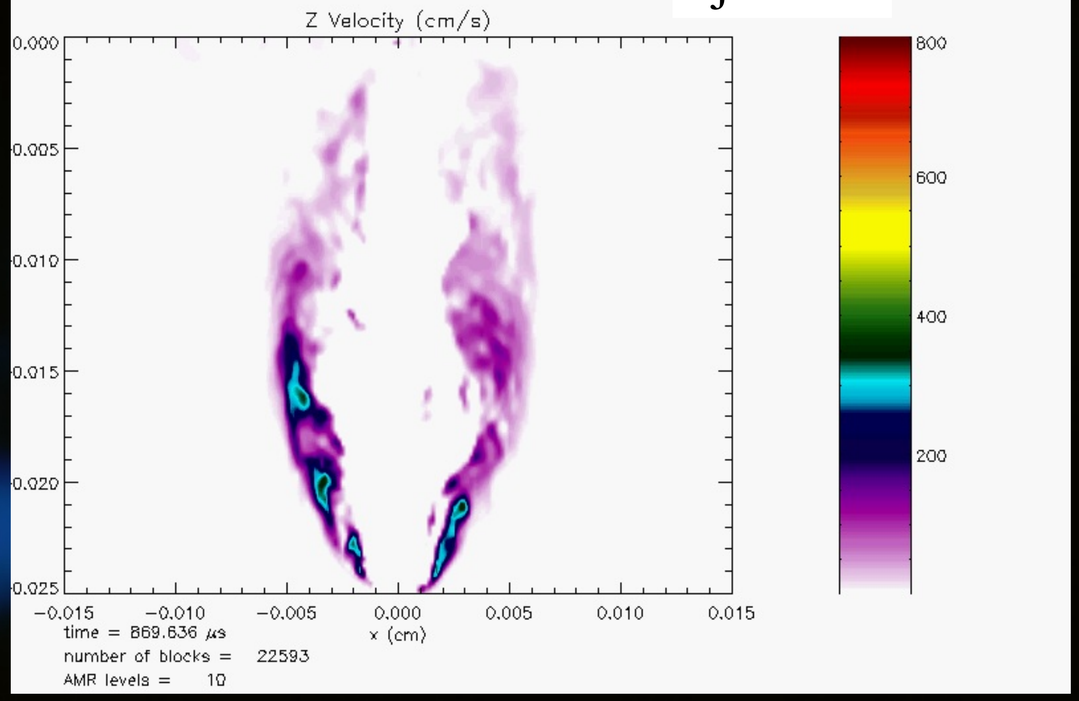
  2) Dynamics is *stochastic*

Numerical experiments

$\mathbf{v}_j \geq 0$



$\mathbf{v}_j \leq 0$



- Plotting only counterstreaming gas ($\mathbf{v}_z \cdot \mathbf{v}_j \leq 0$)
- Lessons:

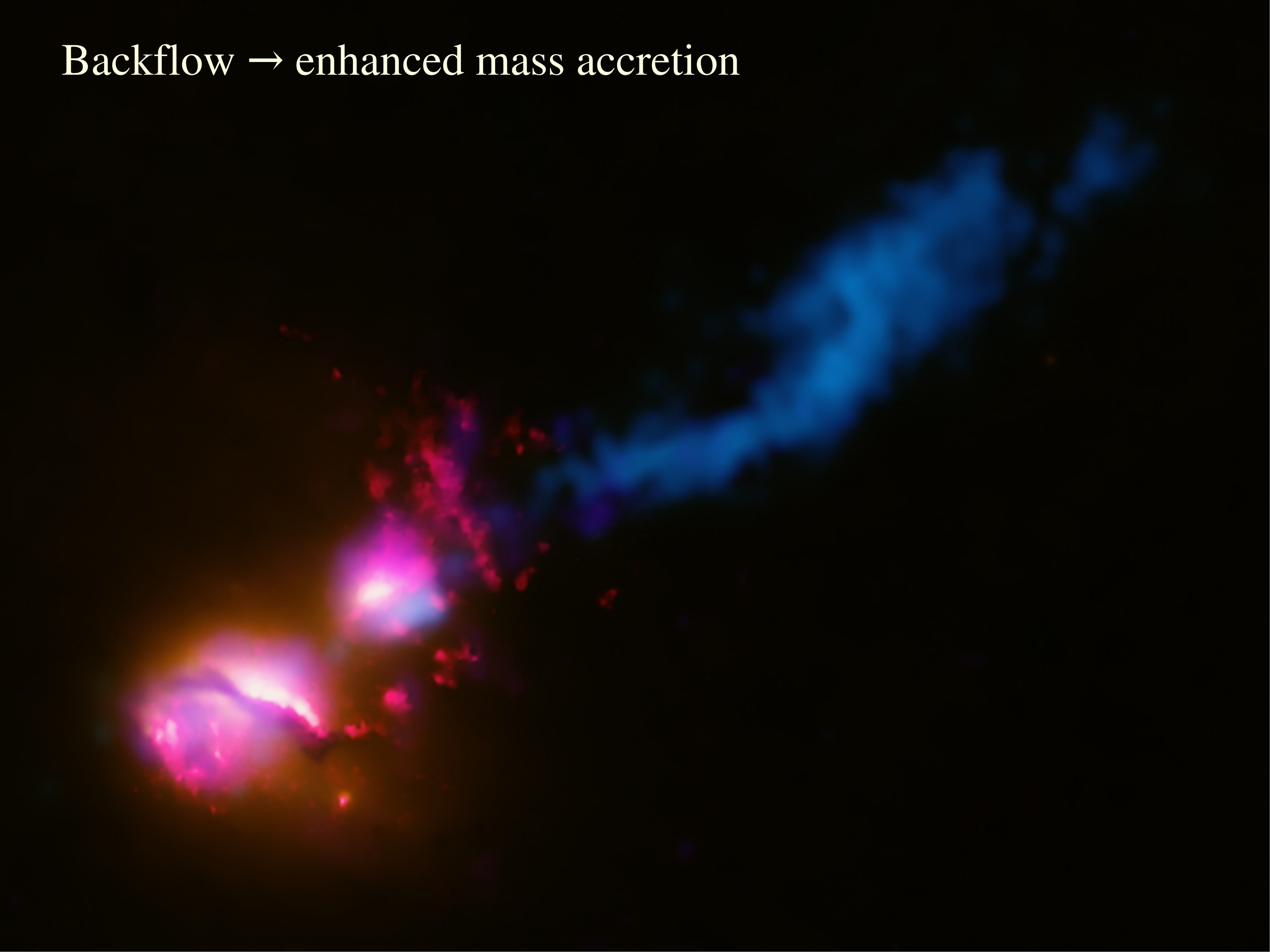
  1) Backflow develops a *large-scale pattern* (HS → meridional plane)

  2) Dynamics is *stochastic*

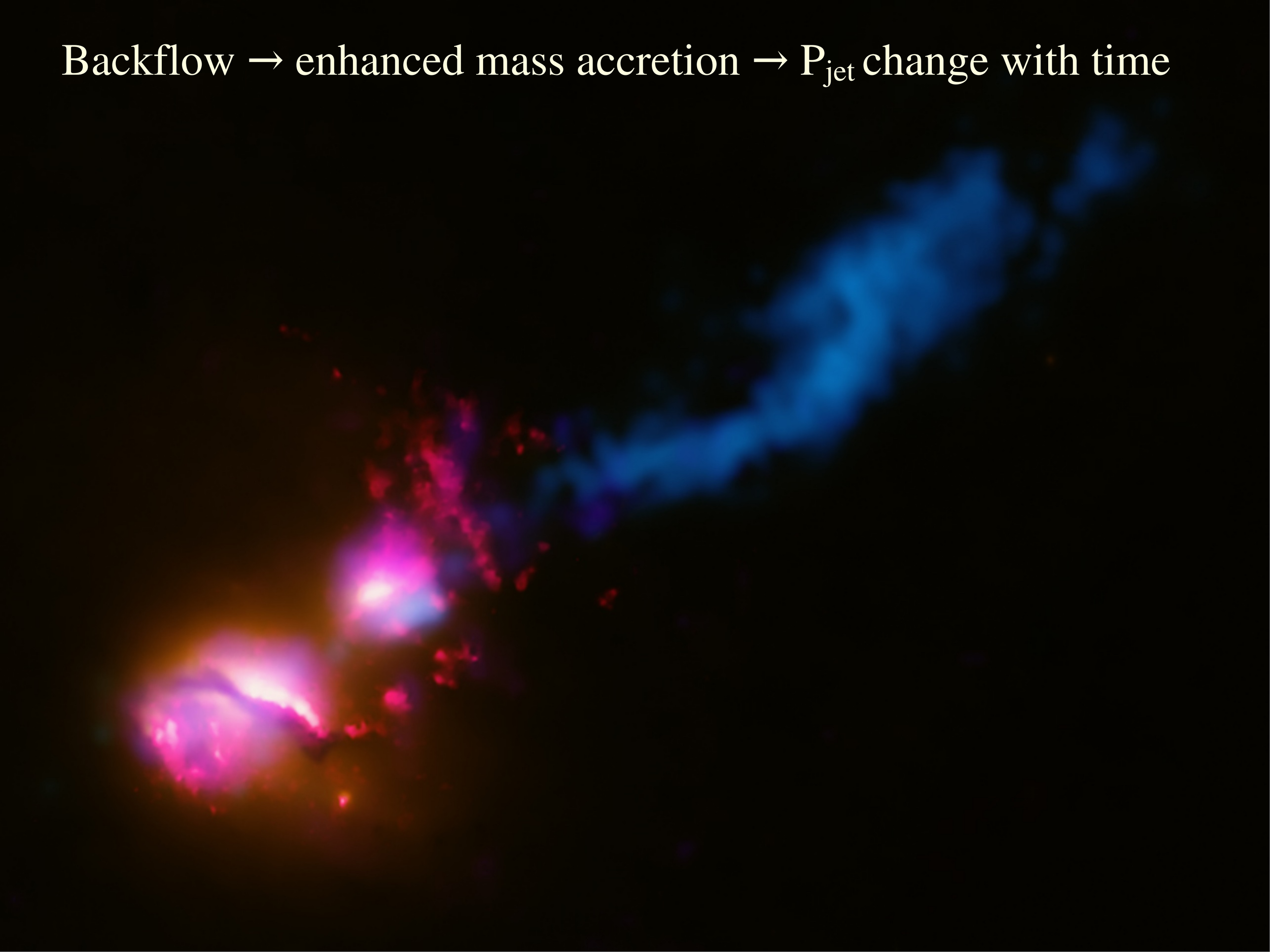
  3) No meridional plane bending of the flow is detected → perpendicolar backflow

Backflow

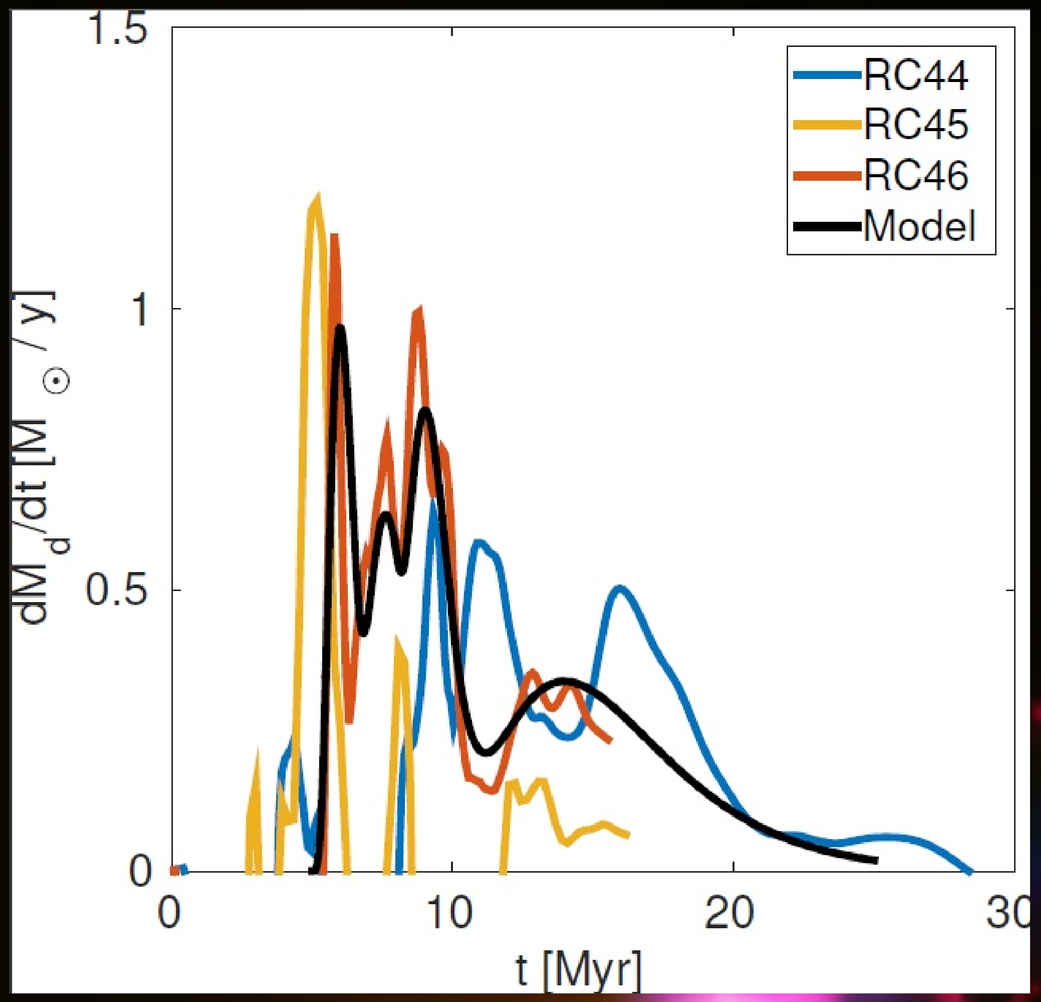
Backflow → enhanced mass accretion



Backflow → enhanced mass accretion → P_{jet} change with time



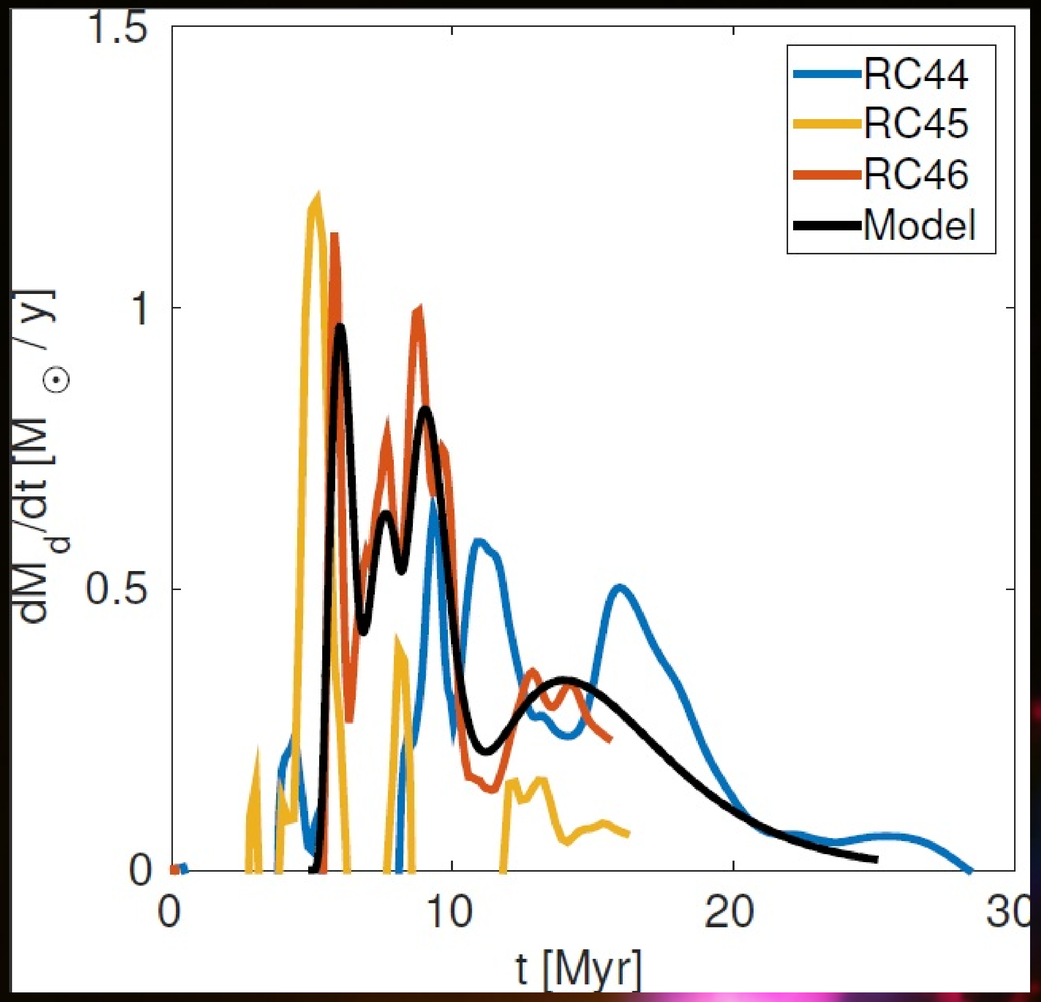
Backflow → enhanced mass accretion → P_{jet} change with time



Black curve: template mass flow model

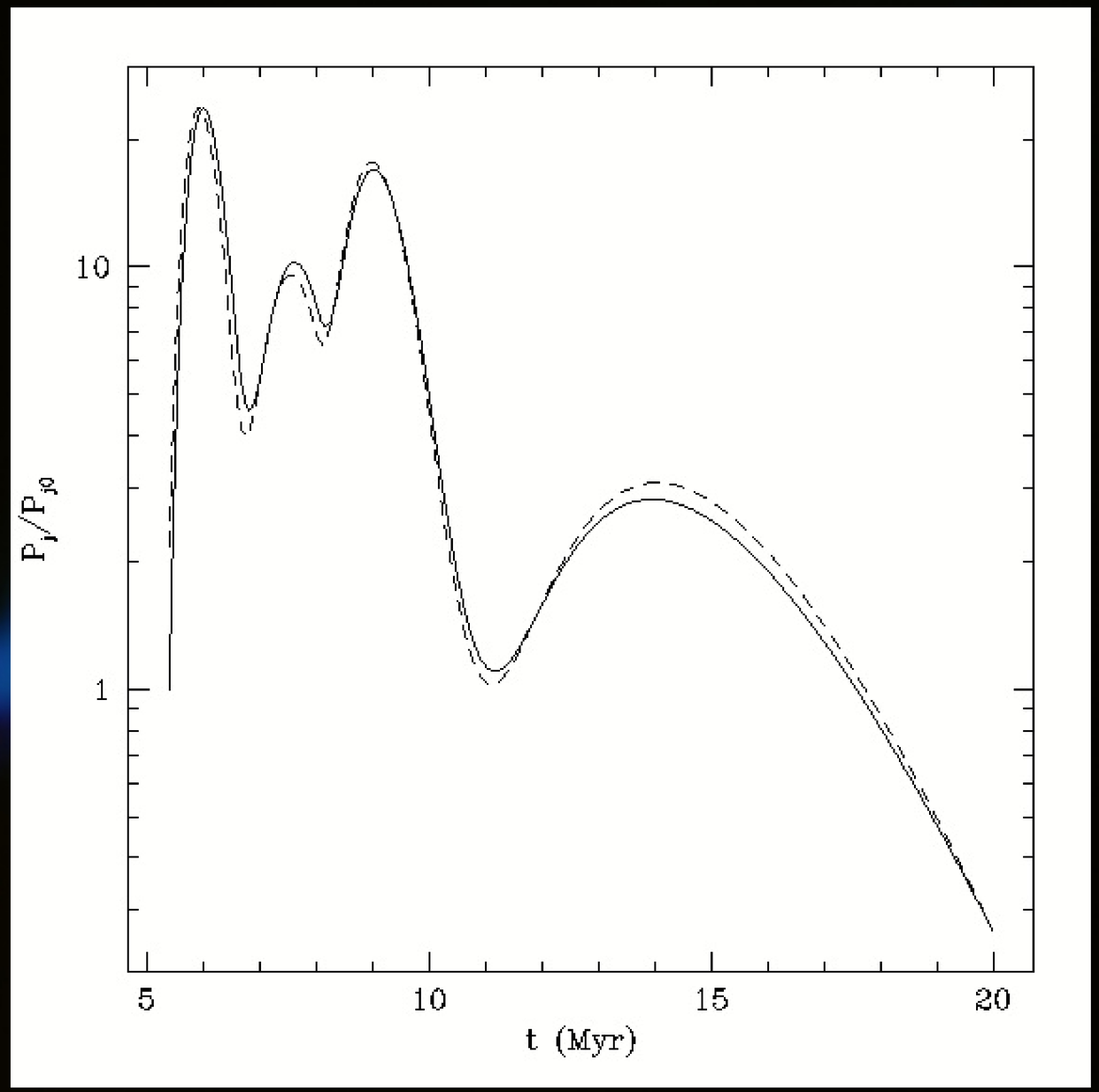
On the right plot the predicted change in $P_{jet} \propto \Sigma^{3/2}(\rho v_z)^{1/2}$

Backflow → enhanced mass accretion → P_{jet} change with time



Black curve: template mass flow model

On the right plot the predicted change in $P_{jet} \propto \Sigma^{3/2}(\rho v_z)^{1/2}$



- P_{jet} enhanced by a factor ~ 10 on 15-20 Myrs.
- EUV ($\lambda \lesssim 1100$ Å) correlates with GHz synchr for RLQ

And now? So what??

Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org

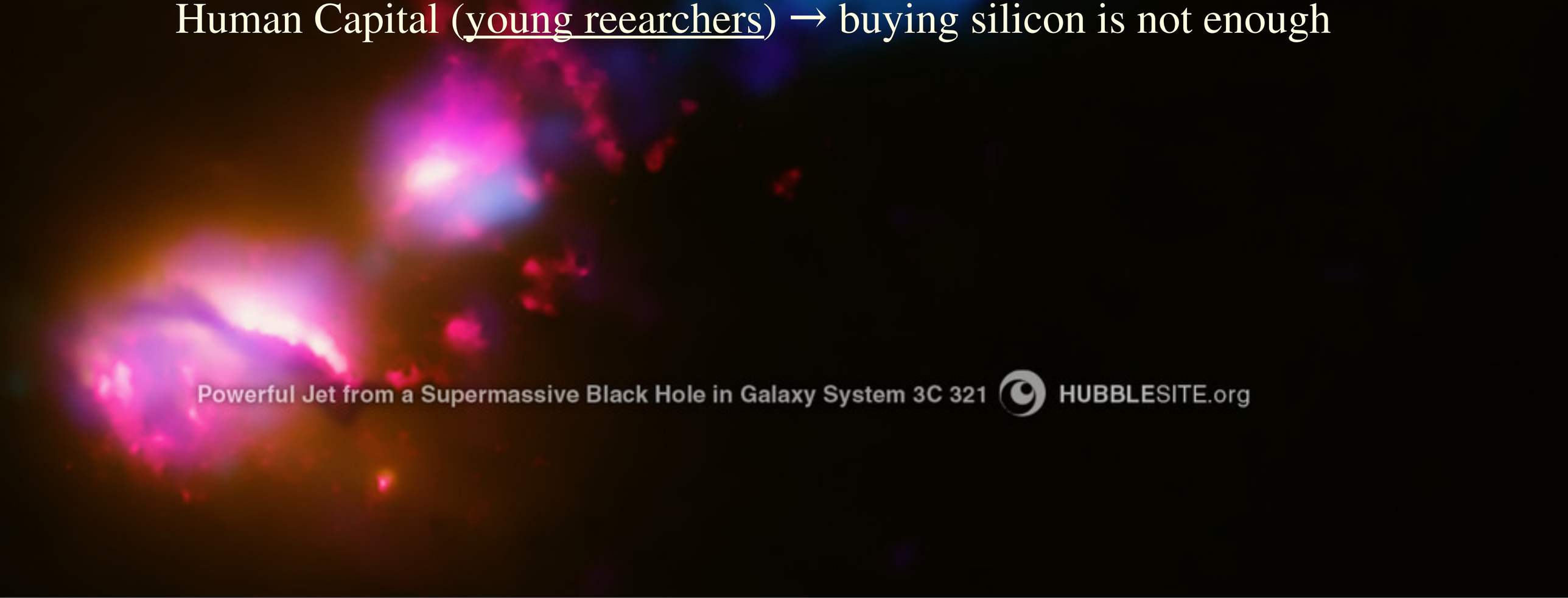
And now? So what??

- Star formation is classical physics → yet SMBH-galaxy co-evolution bridges with GWs and GR

And now? So what??

- Star formation is classical physics → yet SMBH-galaxy co-evolution bridges with GWs and GR

- The investment to fill the gap between subparsec (star formation) and kpc scales is also in new algorithms and HPC computing → Human Capital (young reearchers) → buying silicon is not enough

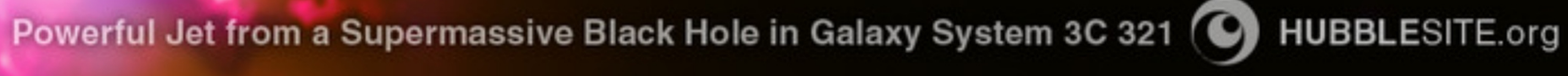


And now? So what??

- Star formation is classical physics → yet SMBH-galaxy co-evolution bridges with GWs and GR

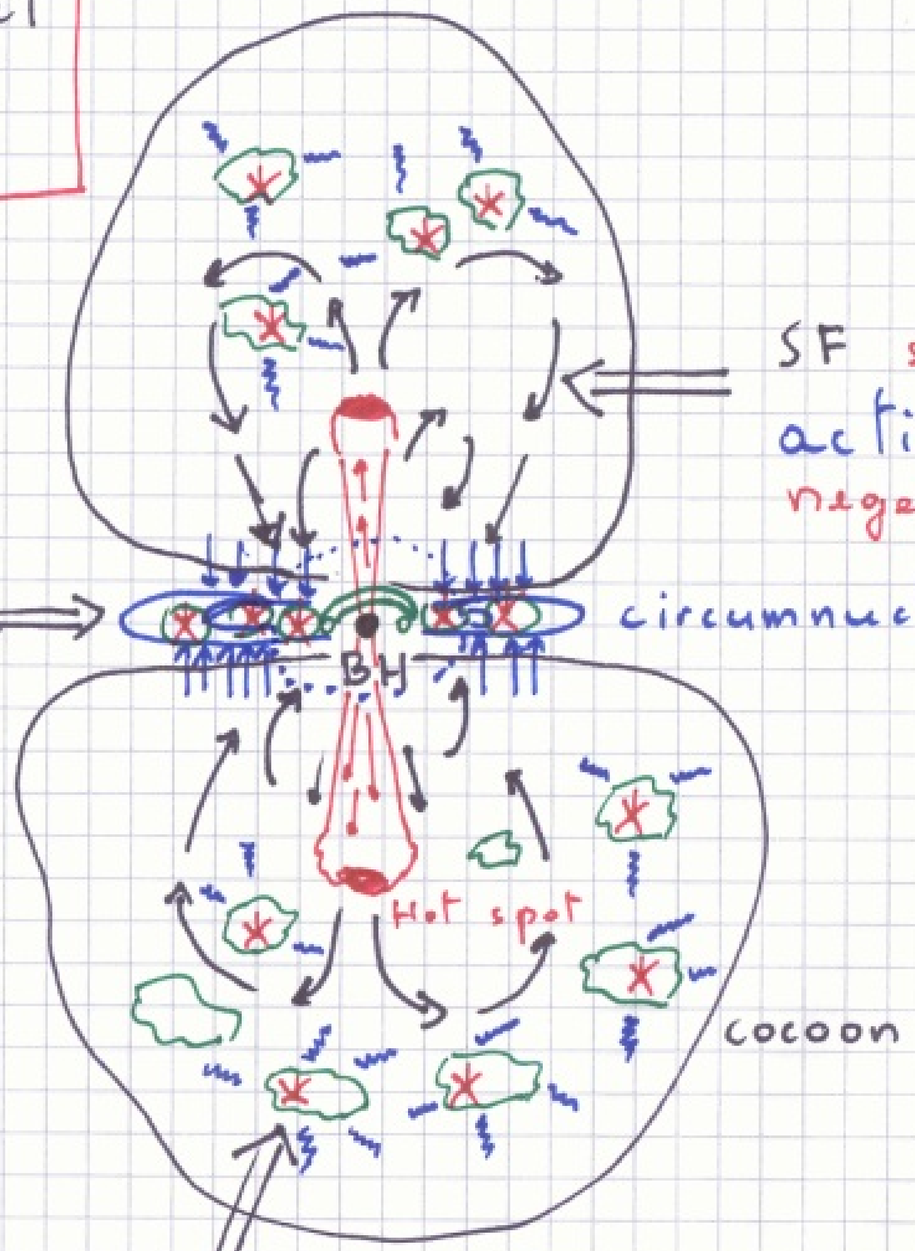
- The investment to fill the gap between subparsec (star formation) and kpc scales is also in new algorithms and HPC computing → Human Capital (young reearchers) → buying silicon is not enough

- Tighter interaction with the Numerical Analysis and Computational Physics communities → credible physical targets and language



My "unified" mechanical AGN feedback model

A "unified" model of AGN feedback



SF suppressed during active cocoon exp. negative FB

circumnuclear disc (~ 20 pc)

SF induced by backflow's compr. in the circumnucl. disc positive FB

cocoon

: cooling ISM/IGM clouds

SF induced by therm. inst. during passive cocoon exp. positive FB

V. Emonts Oct. '02 Delhi

Black Holes-galaxies co-evolution
Physical Models

- Jet's feedback - Jet's production, SMBH accretion: highly decoupled in spatial scales, <u>but</u> strongly coupled energetically (GRMHD - Tchekhovskoy, Sądowski, McKinney..., theory: Paczinsky, Abramowicz, Narayan, Lasota)
- Positive feedback: outflows at $z \gtrsim 1$ (Maiolino, Lehnert) - is it <u>cosmologically</u> significant? Can it boost sognificantly the SFR?
- SMBH physics: directly accessible through GWs→ BHs' growth history and SF history

- Backflow: low L_z accreting gas from large (SF) to AU (BH) accretion scales, connection between

Powerful Jet from a Supermassive Black Hole in Galaxy System 3C 321 HUBBLESITE.org