

Blazars:

beams for astro &
astroparticle physics

Fabrizio Tavecchio

INAF-OAB

Active Galactic Nuclei 11

23-26 September 2014, Trieste



Where Black Holes and Galaxies Meet

Outline

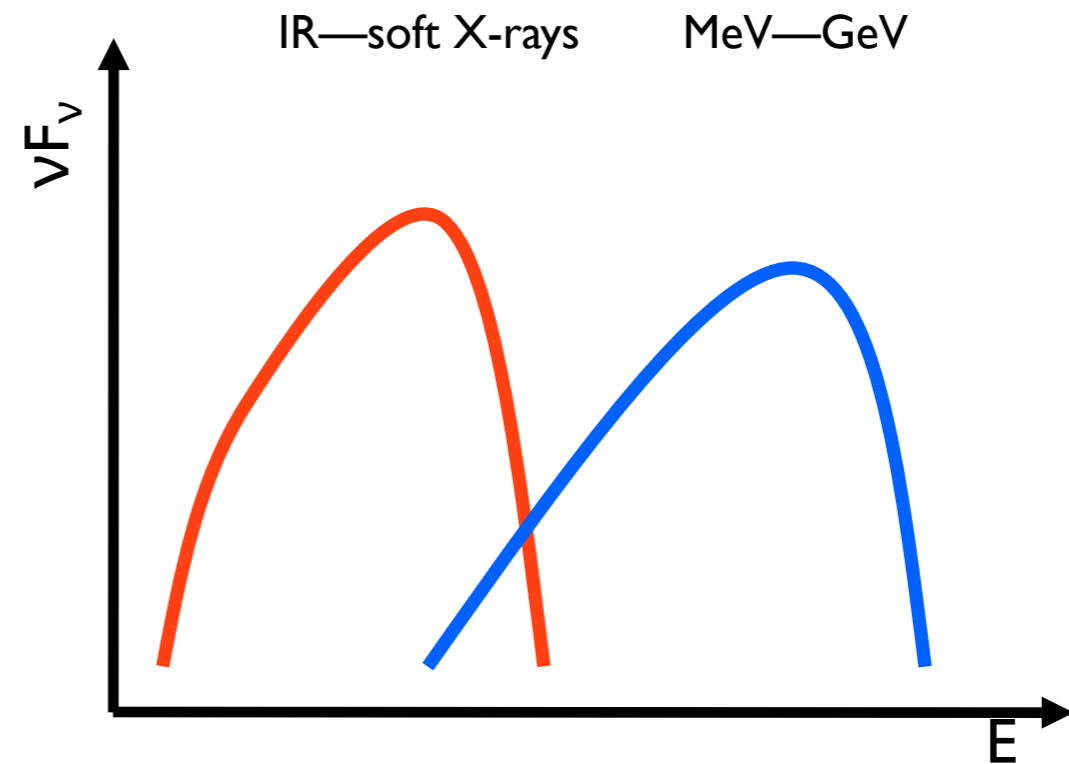
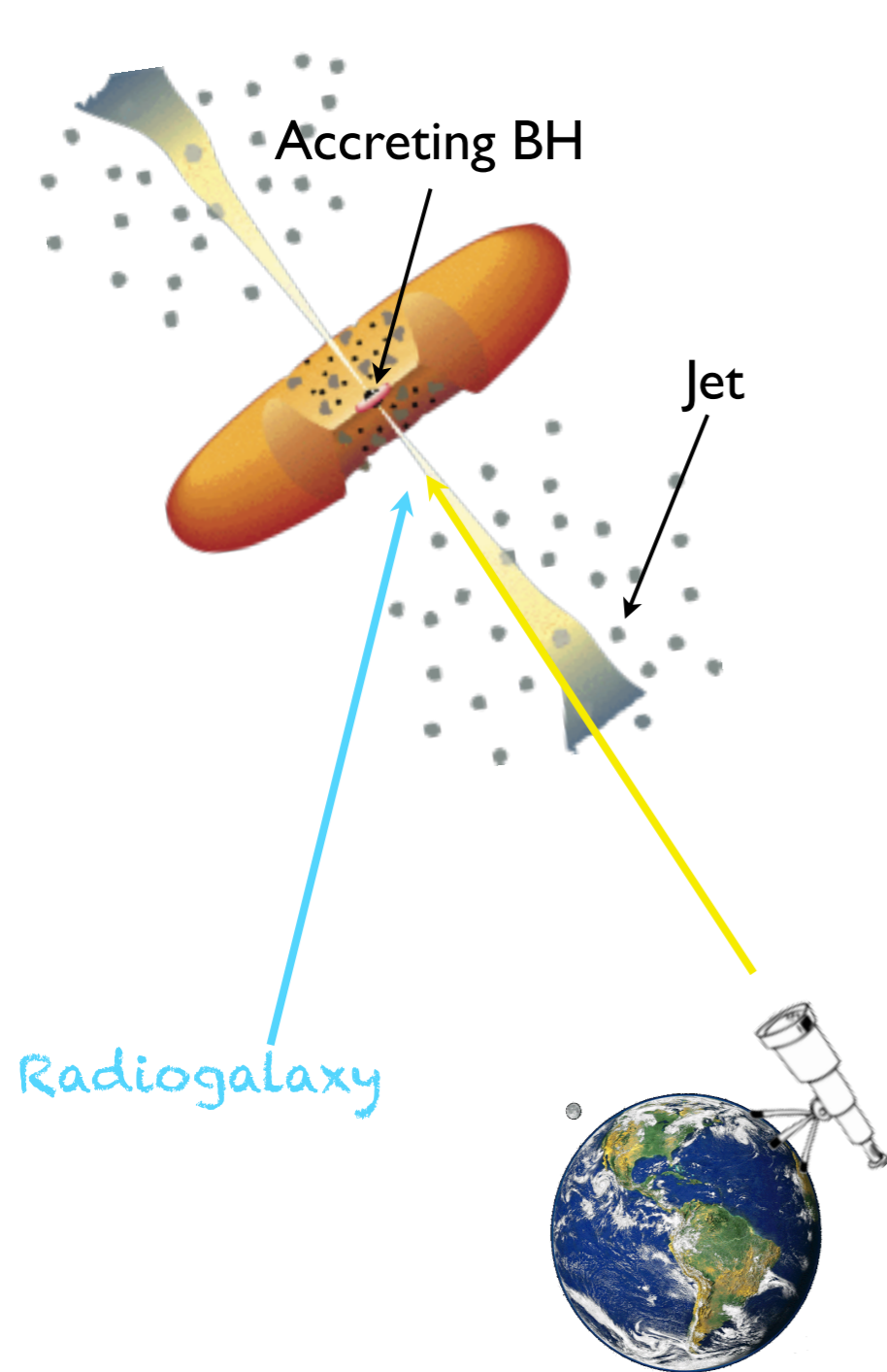
Introduction

Current issues: location and dissipation

Structured jets: radiogalaxies & neutrinos

Extreme BL Lacs: UHECR, ALP & LIV

Blazars in a nutshell



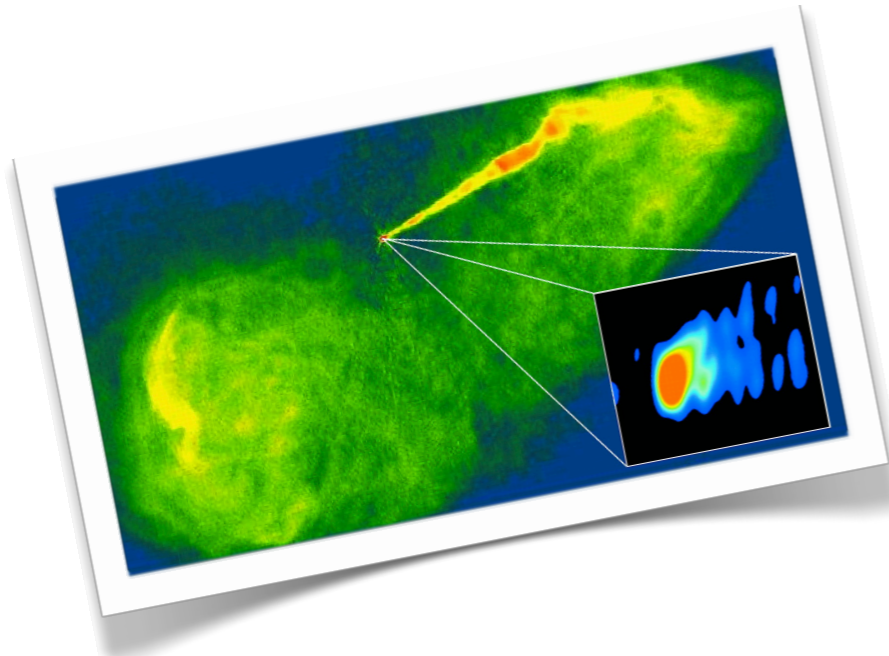
SED dominated by the relativistically boosted non-thermal continuum emission of the jet.

$$L_{\text{obs}} = L' \delta^4 \quad \delta = \frac{1}{\Gamma(1 - \beta \cos \theta_v)}$$

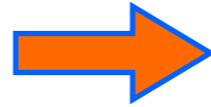
Synchrotron and **IC** in leptonic models.

Also hadronic scenarios (synchrotron or photo-meson emission)

Blazars: jet physics

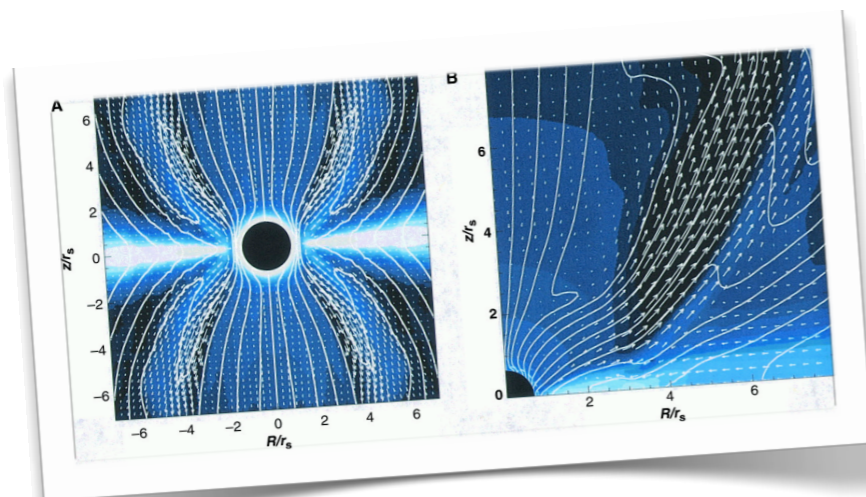
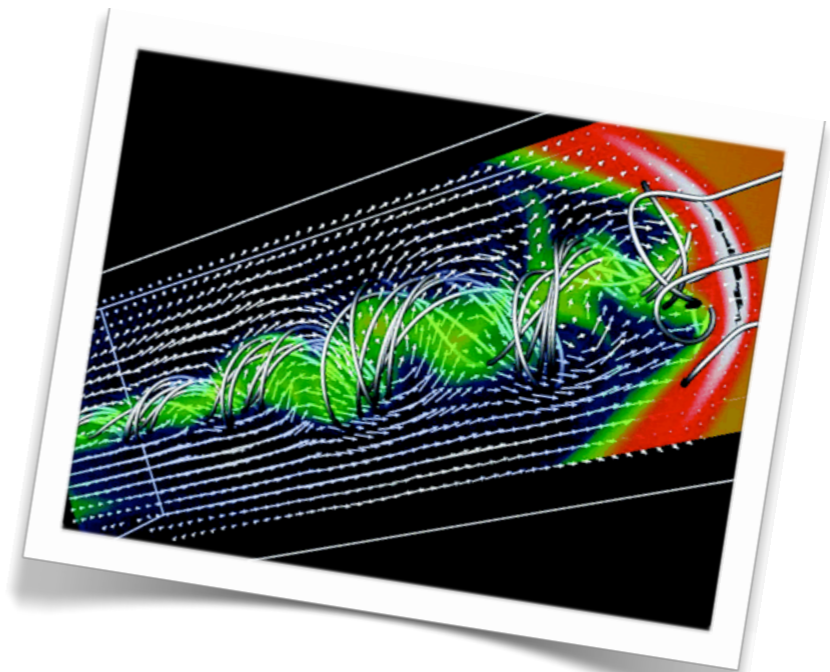


Jet speed,
composition,
power



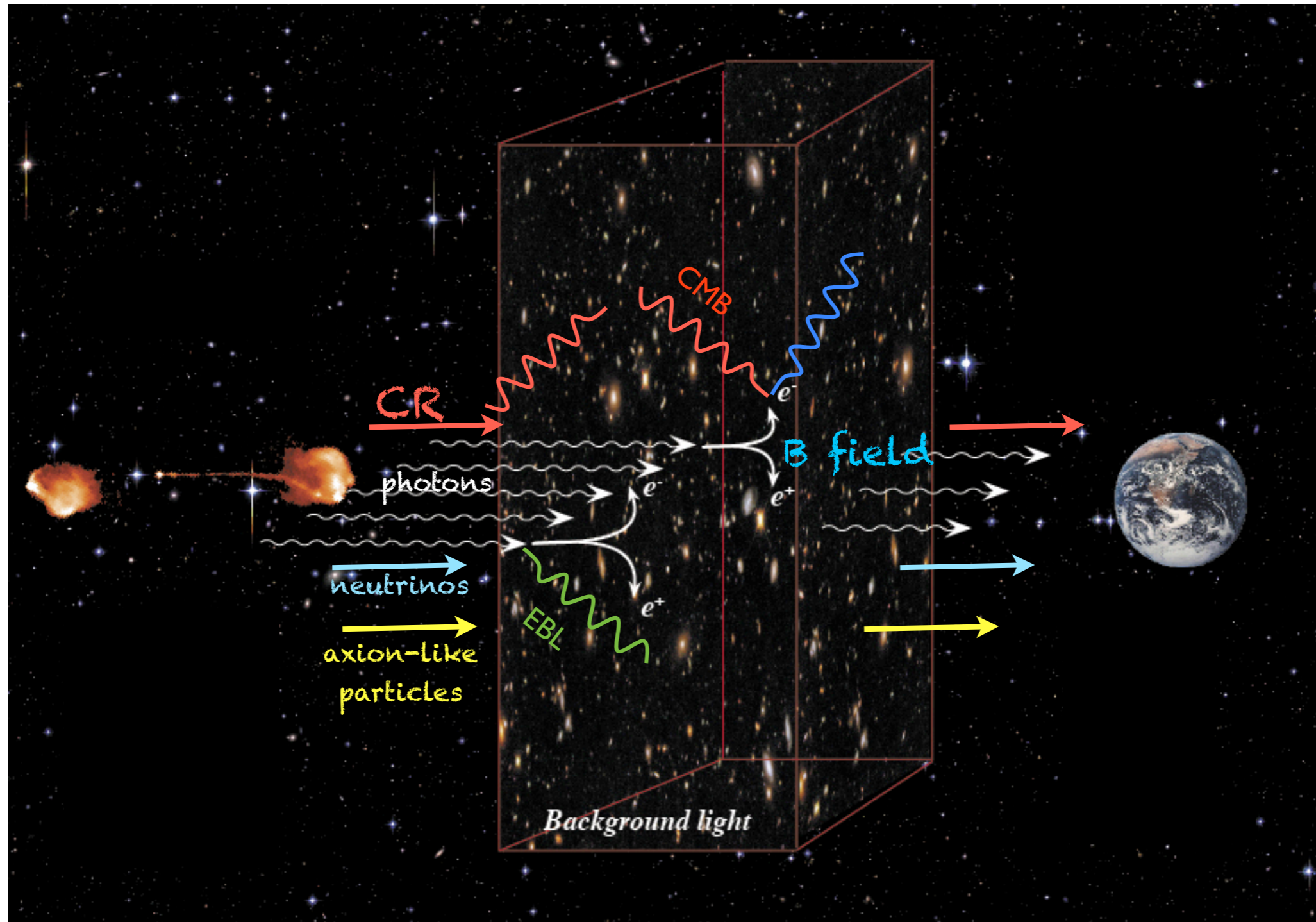
G. Ghisellini & T. Sbarrato
talks

Magnetic fields,
particle acceleration



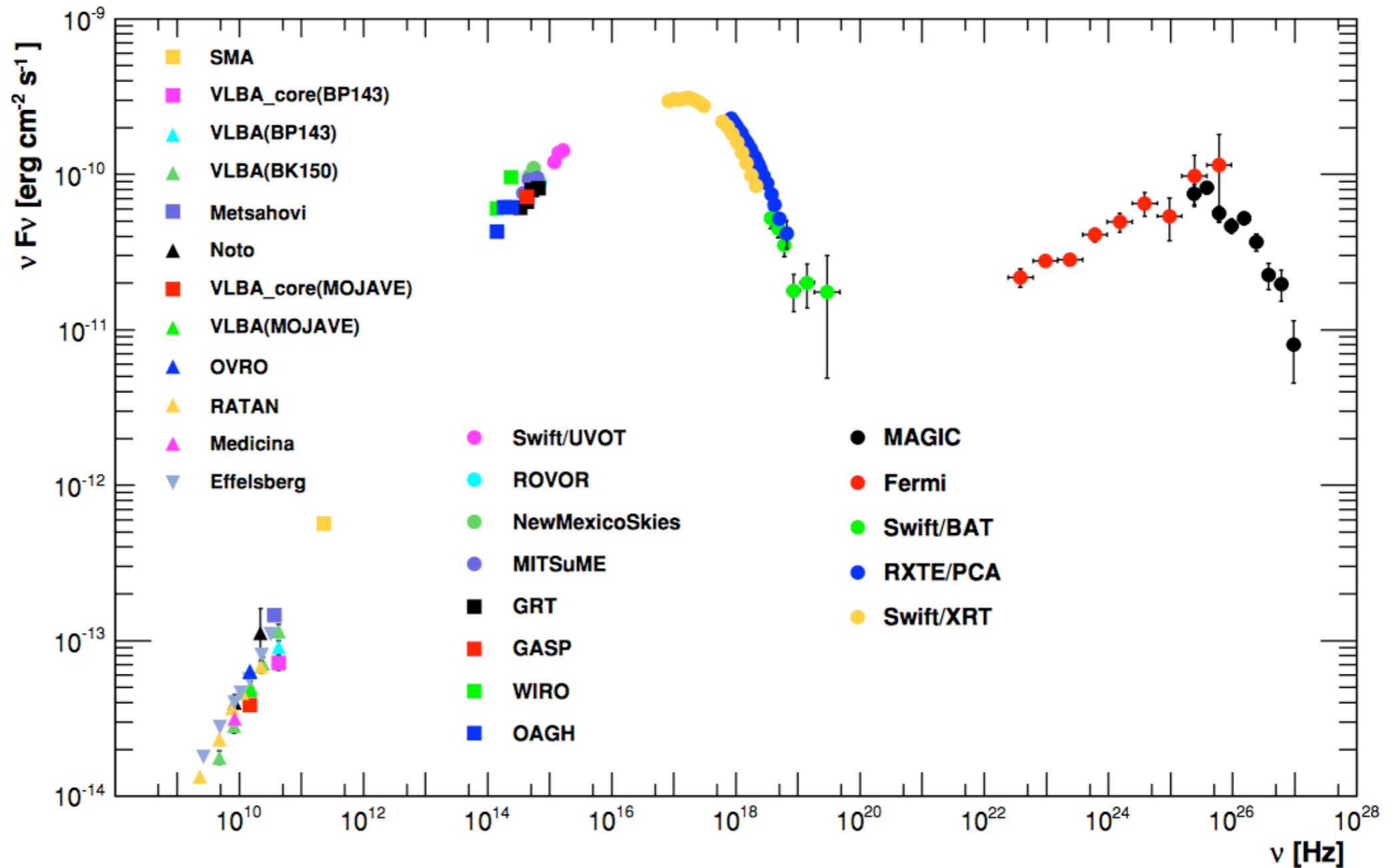
Formation, collimation,
acceleration

Blazars: high-energy particle beams

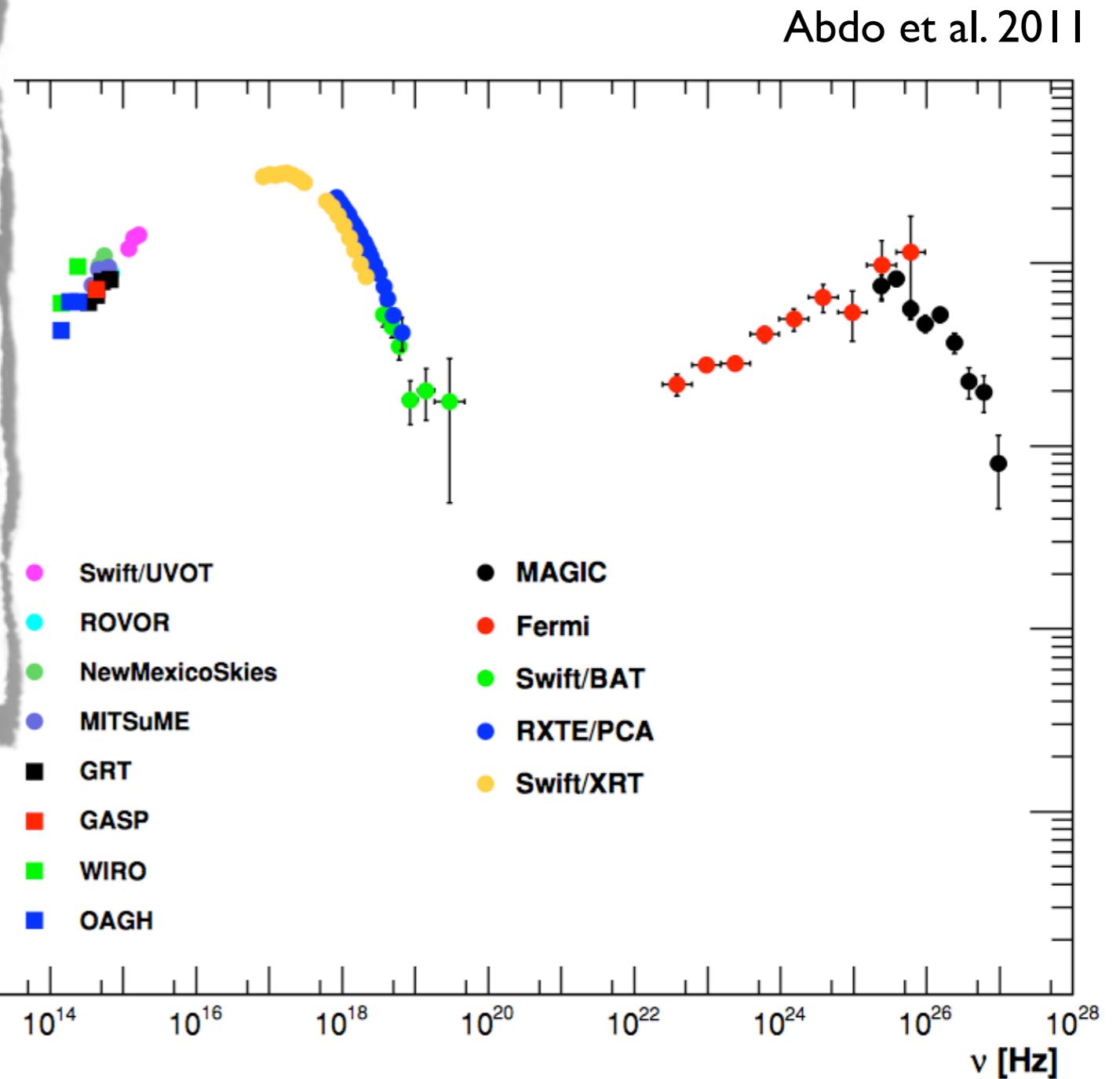
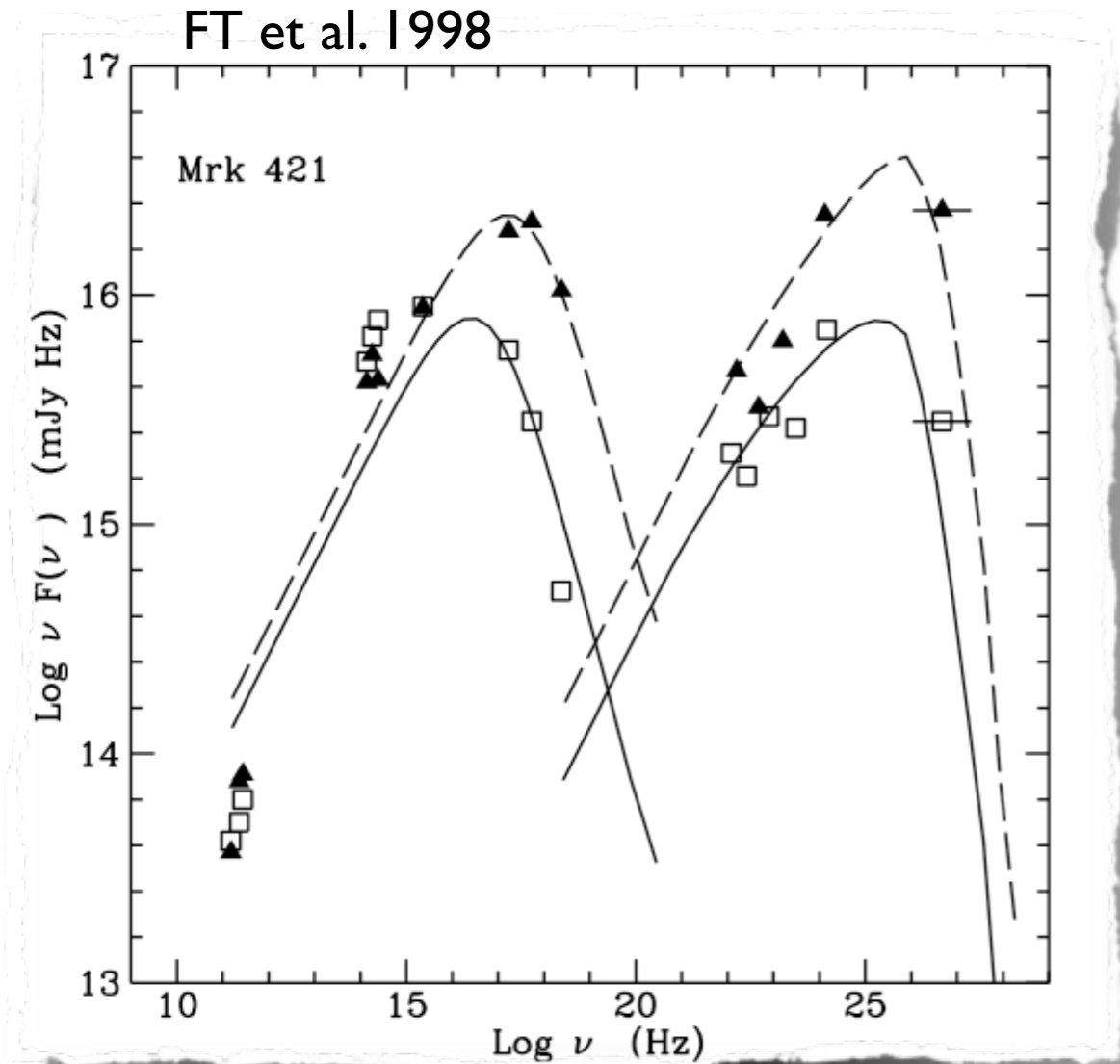


The golden era

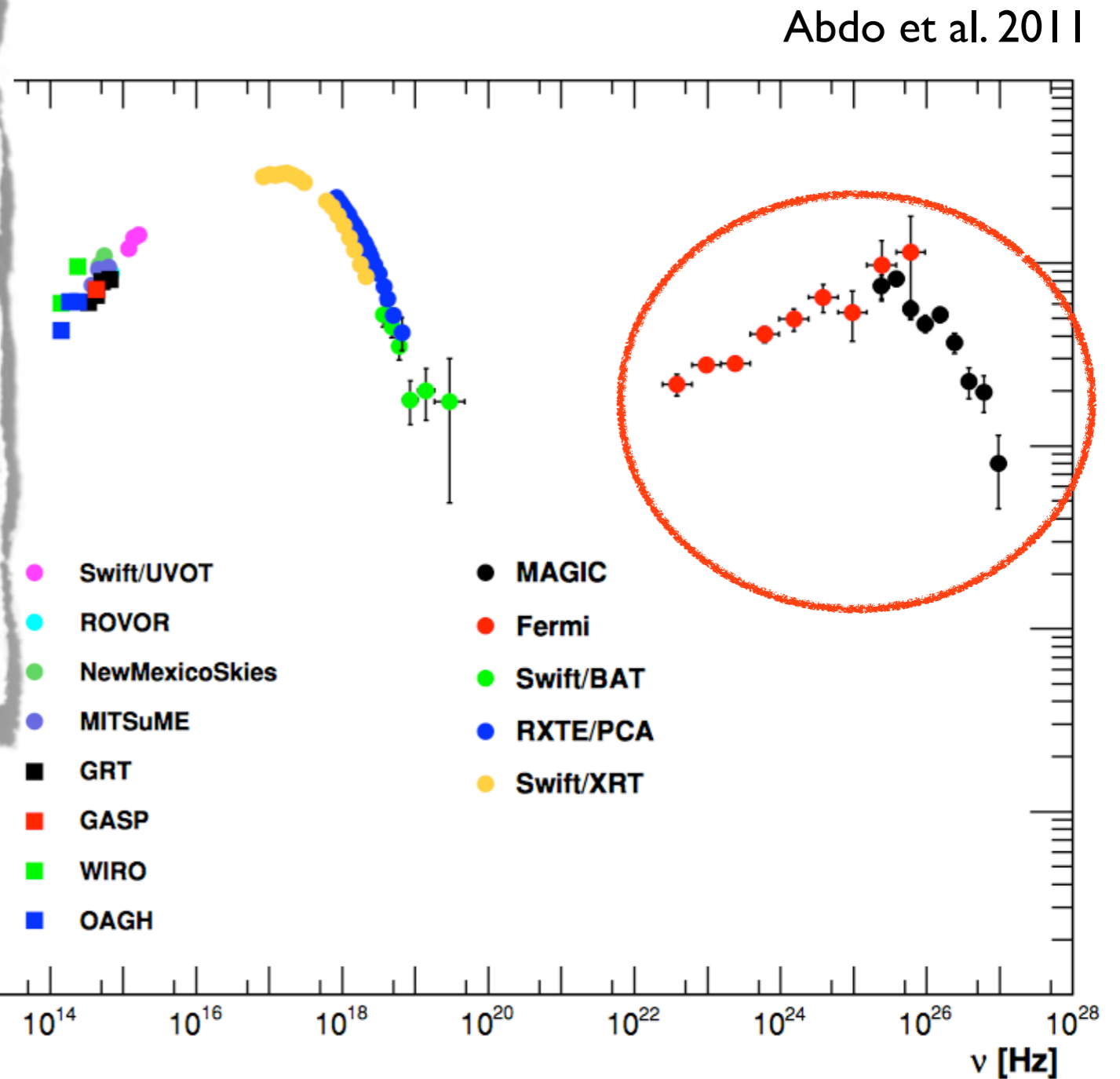
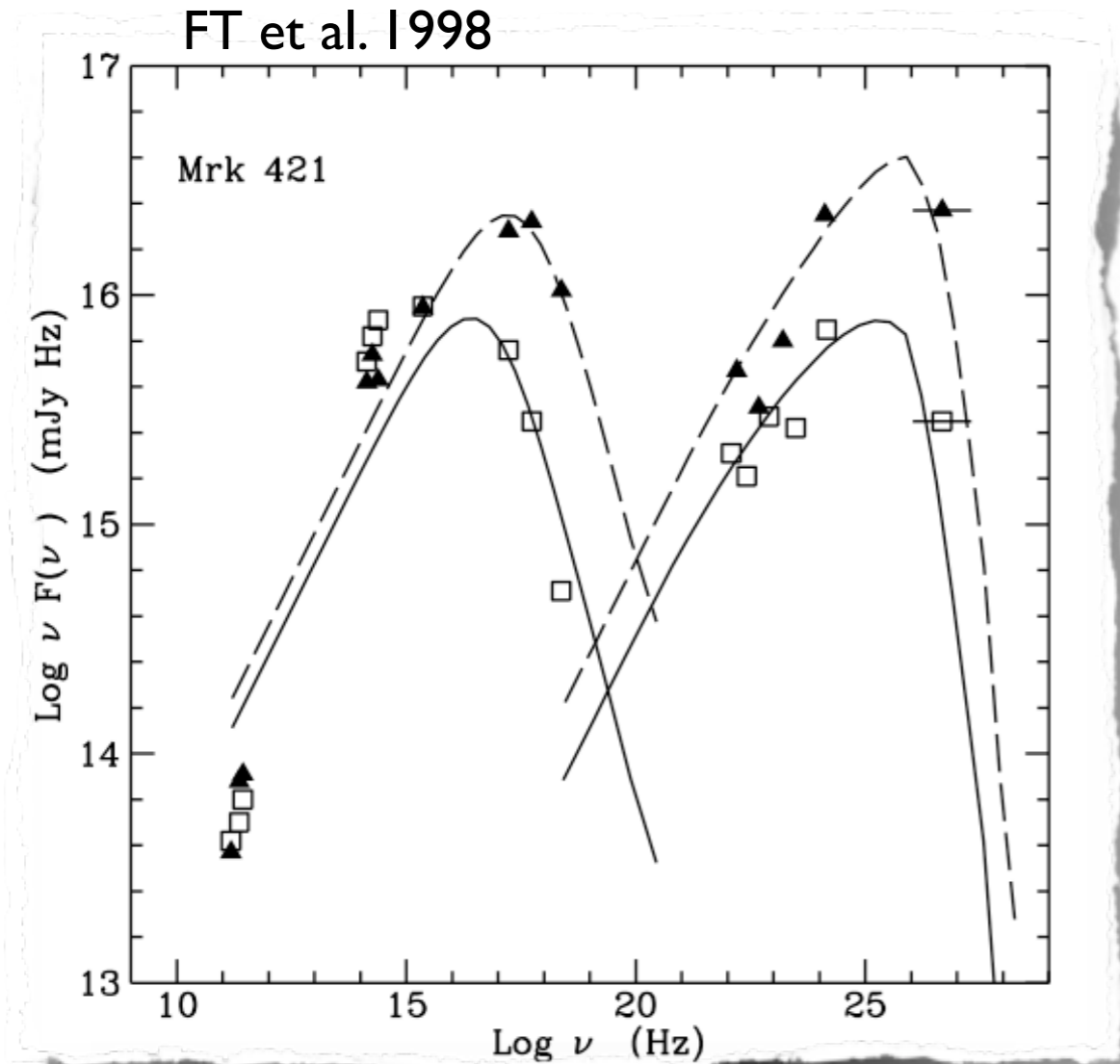
Abdo et al. 2011



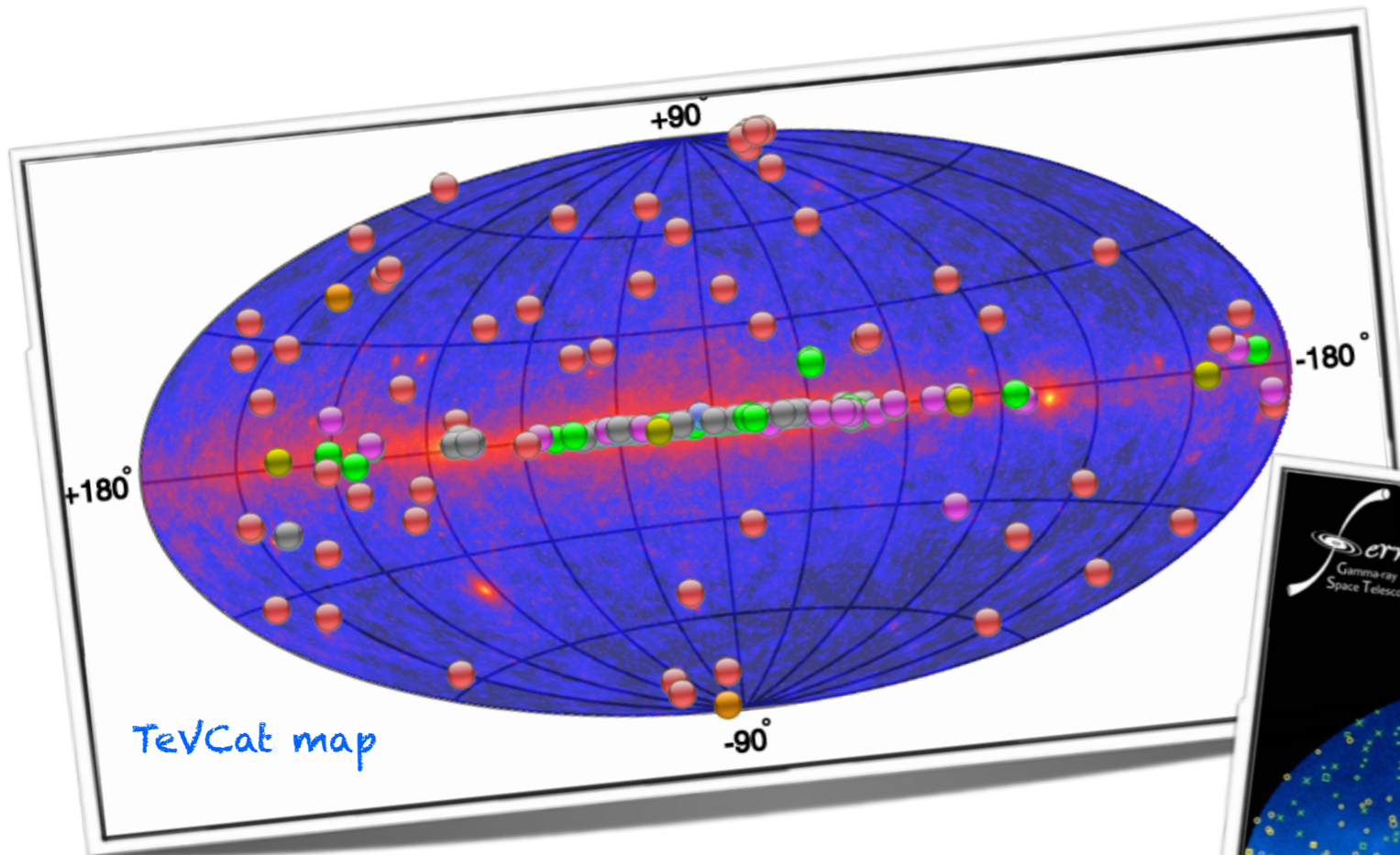
The golden era



The golden era



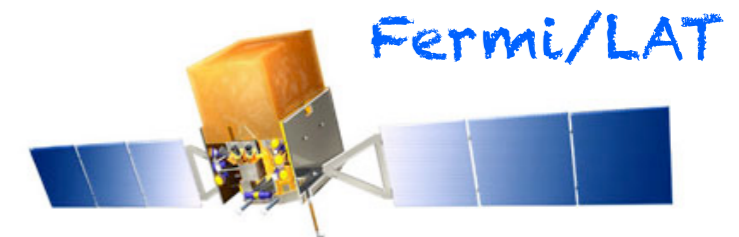
The golden era



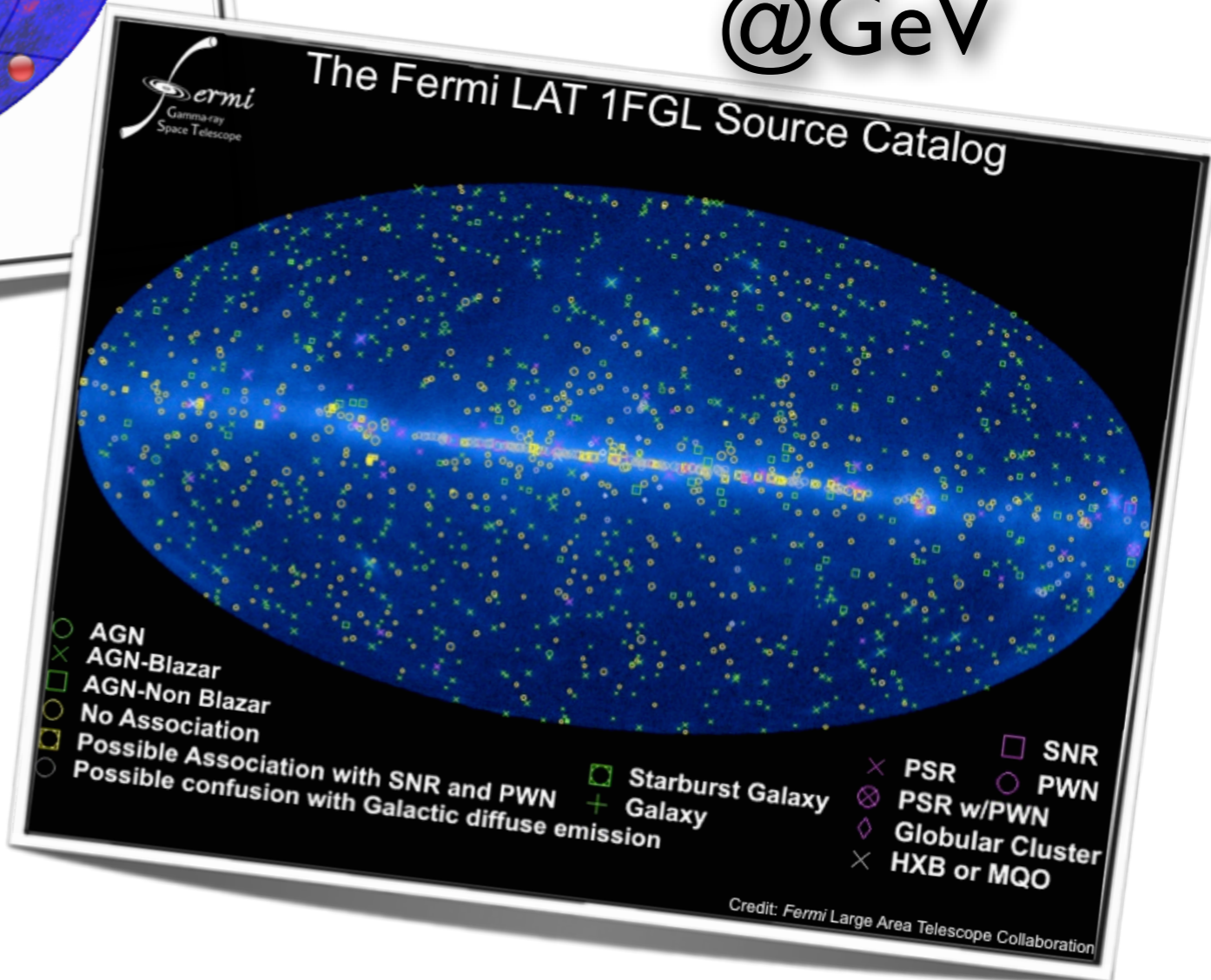
@TeV



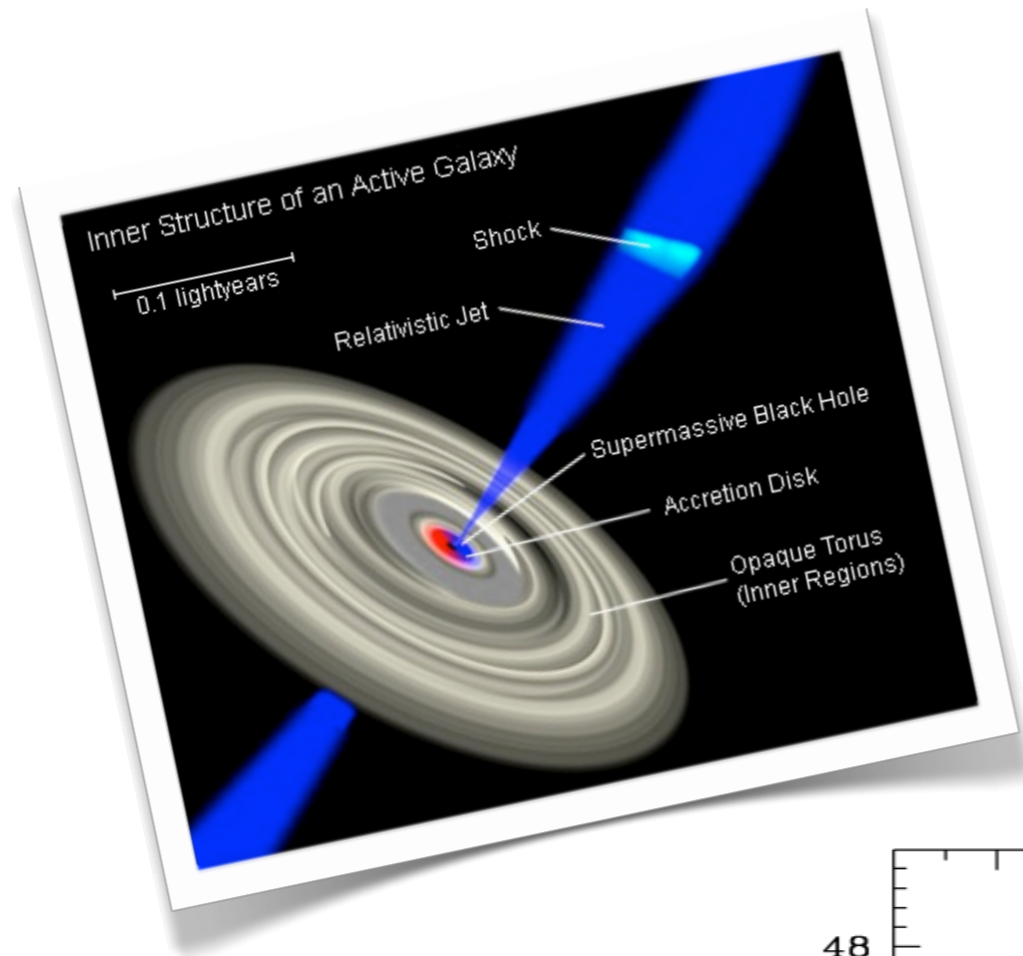
Cherenkov Arrays



@GeV



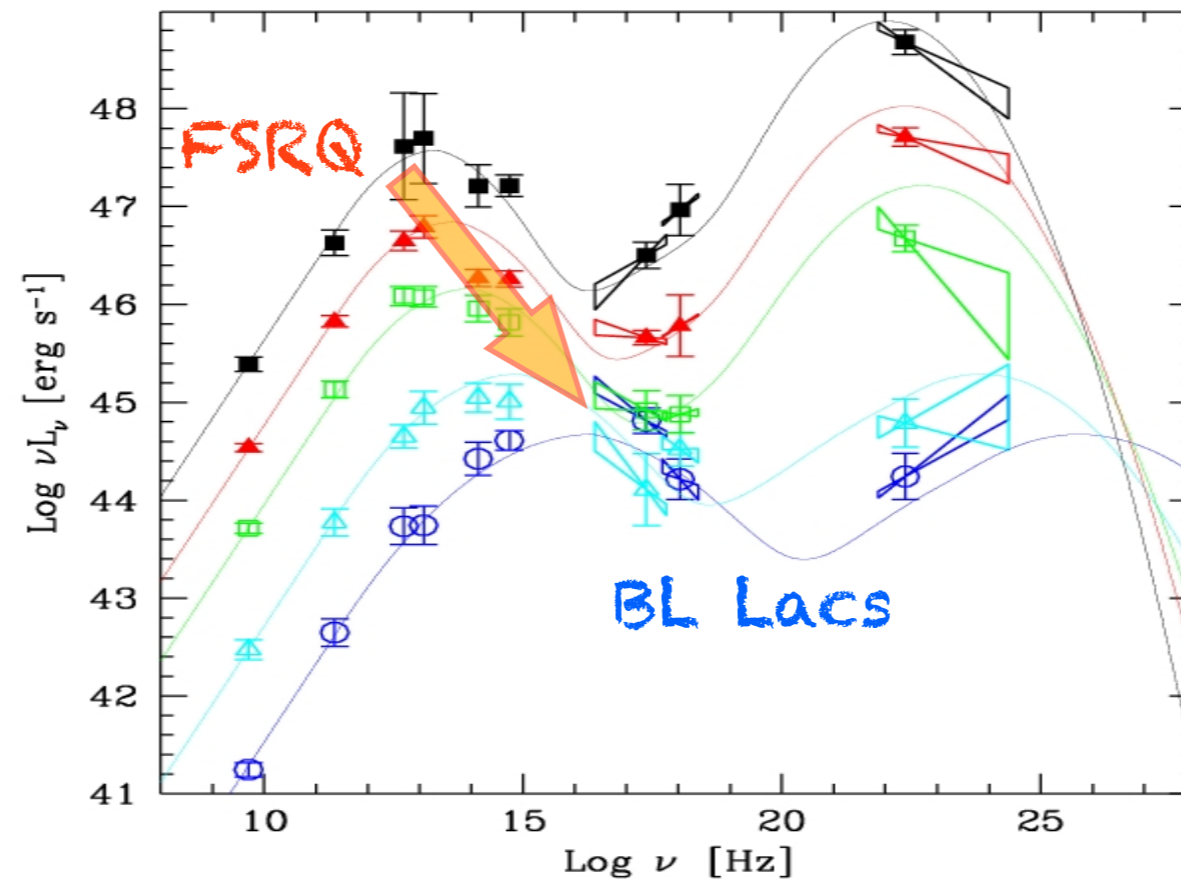
Blazars: basic phenomenology



Blazars occur in two flavors:

FSRQ: high power, thermal optical components

BL Lacs: low power, lack of important thermal comp.

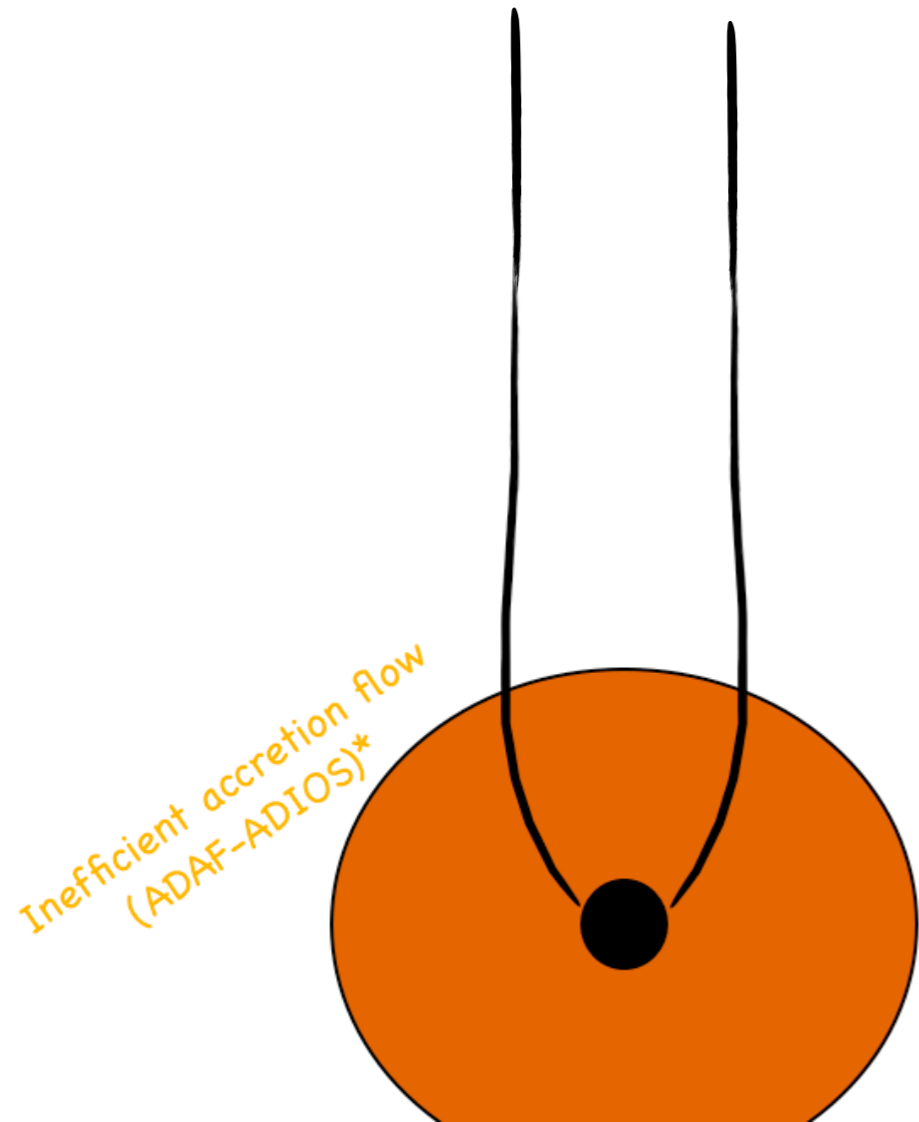


The "blazar sequence"

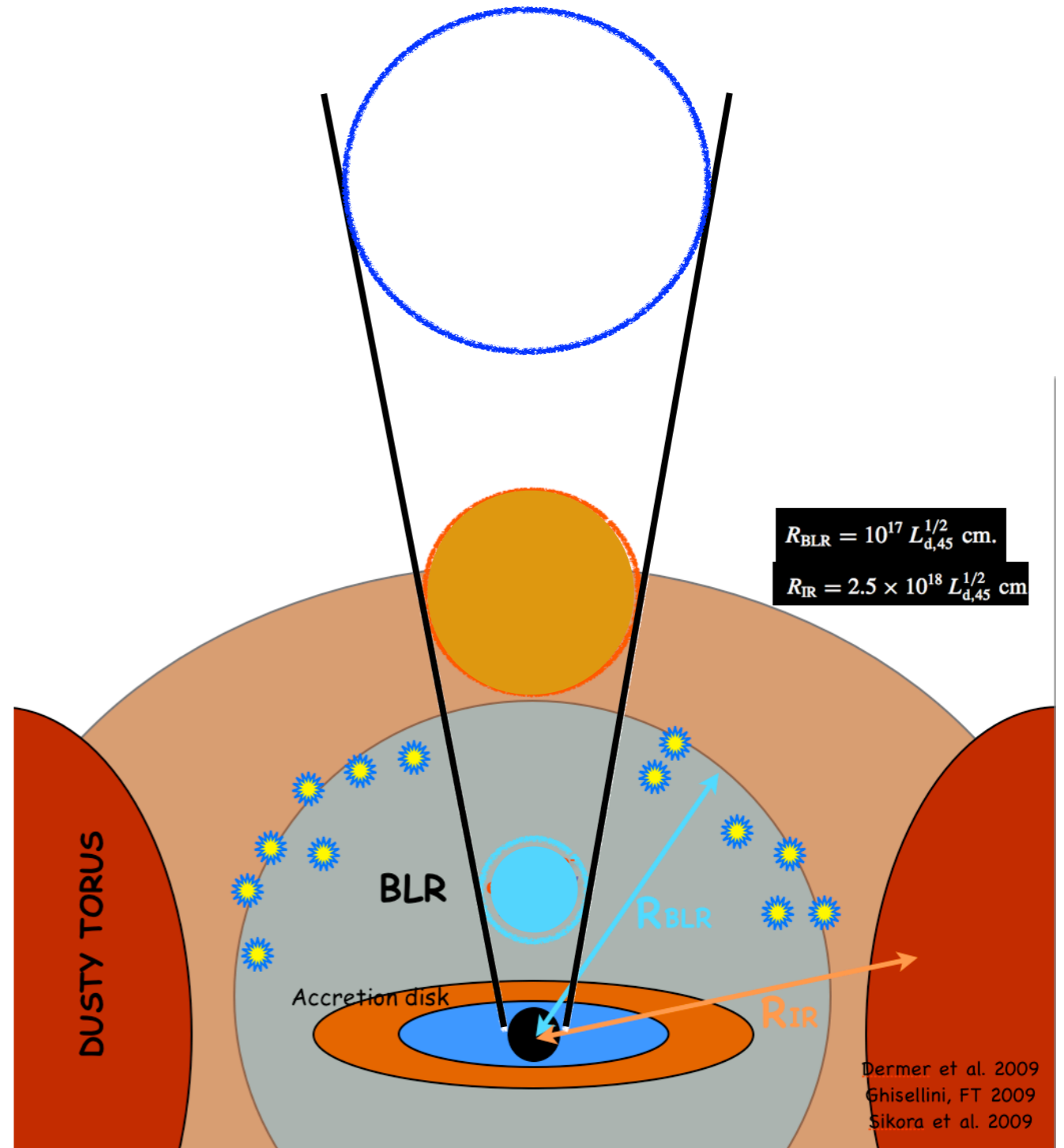
Fossati et al. 1998

Blazars: current discussions

BL Lacs: “naked” jets

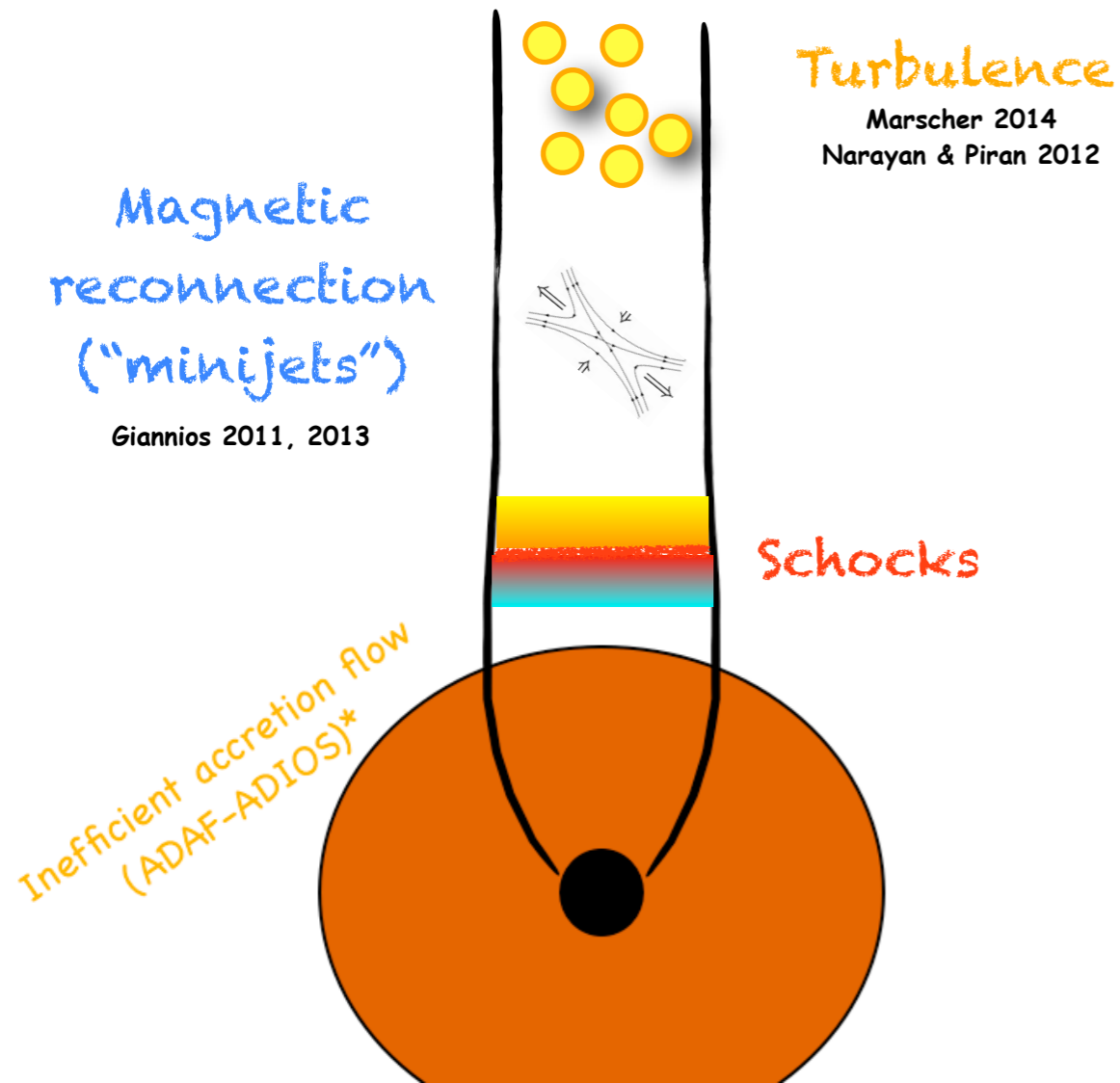


FSRQ: “dressed” jets

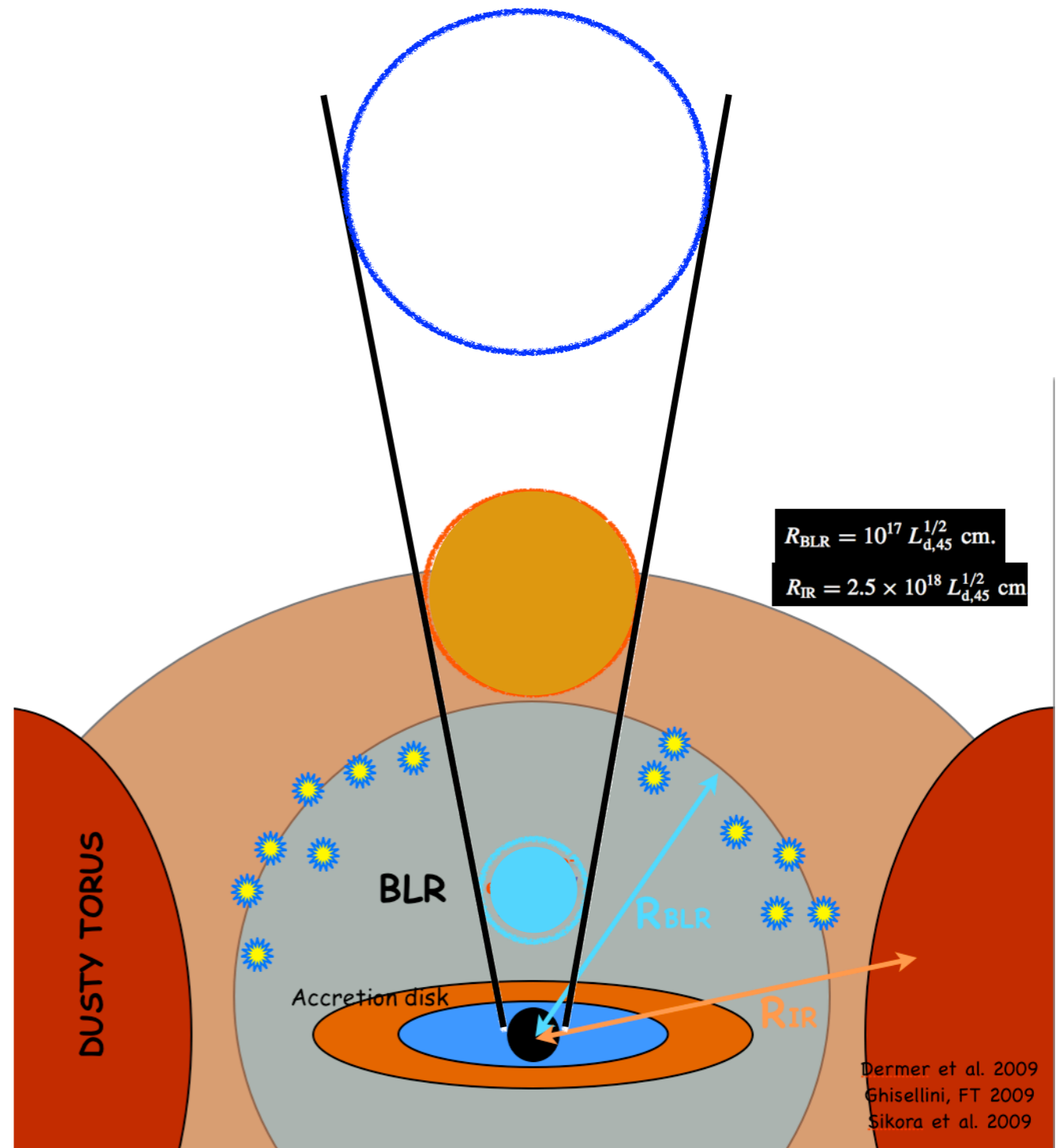


Blazars: current discussions

BL Lacs: "naked" jets

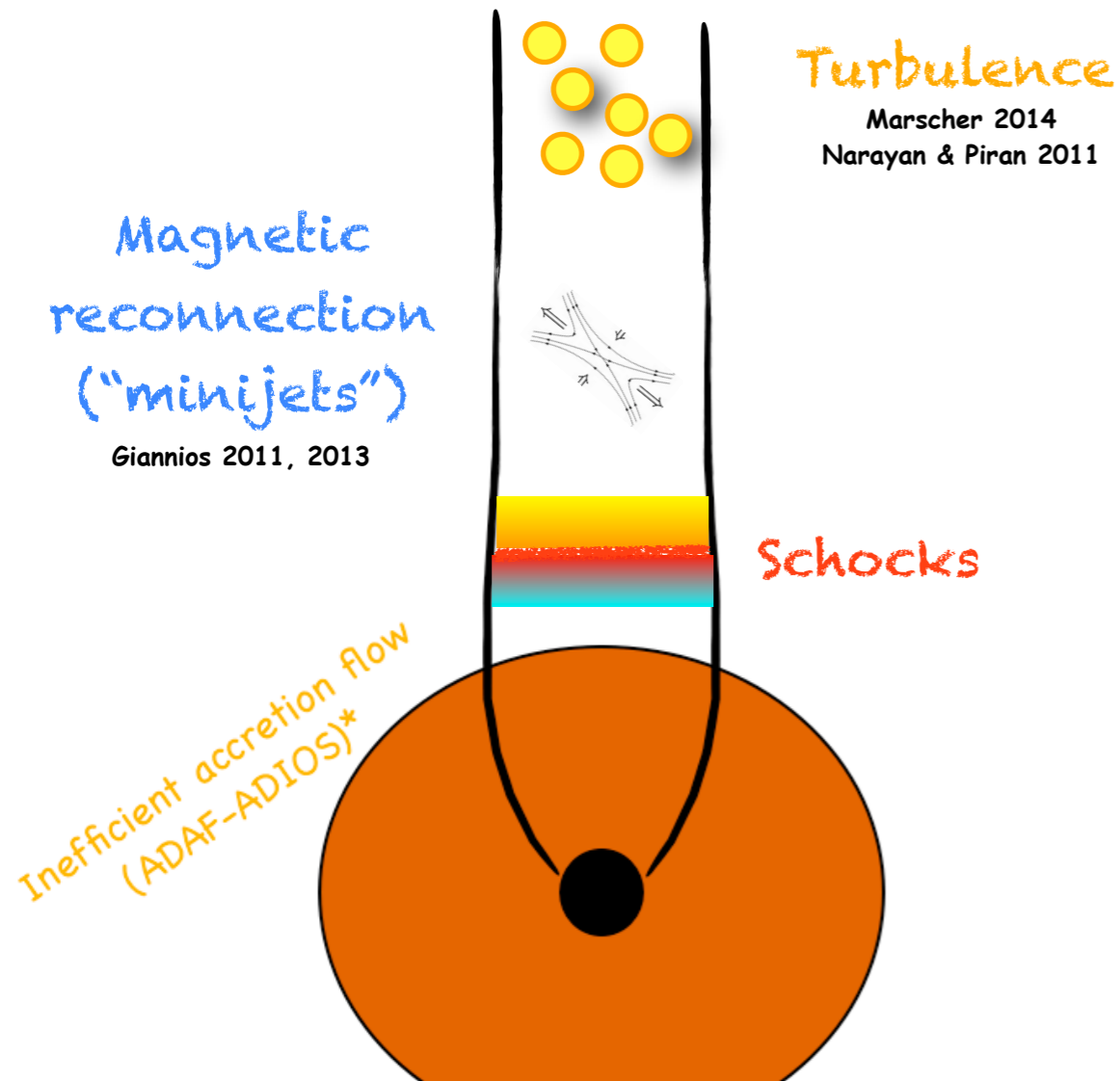


FSRQ: "dressed" jets

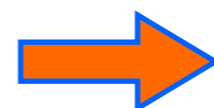
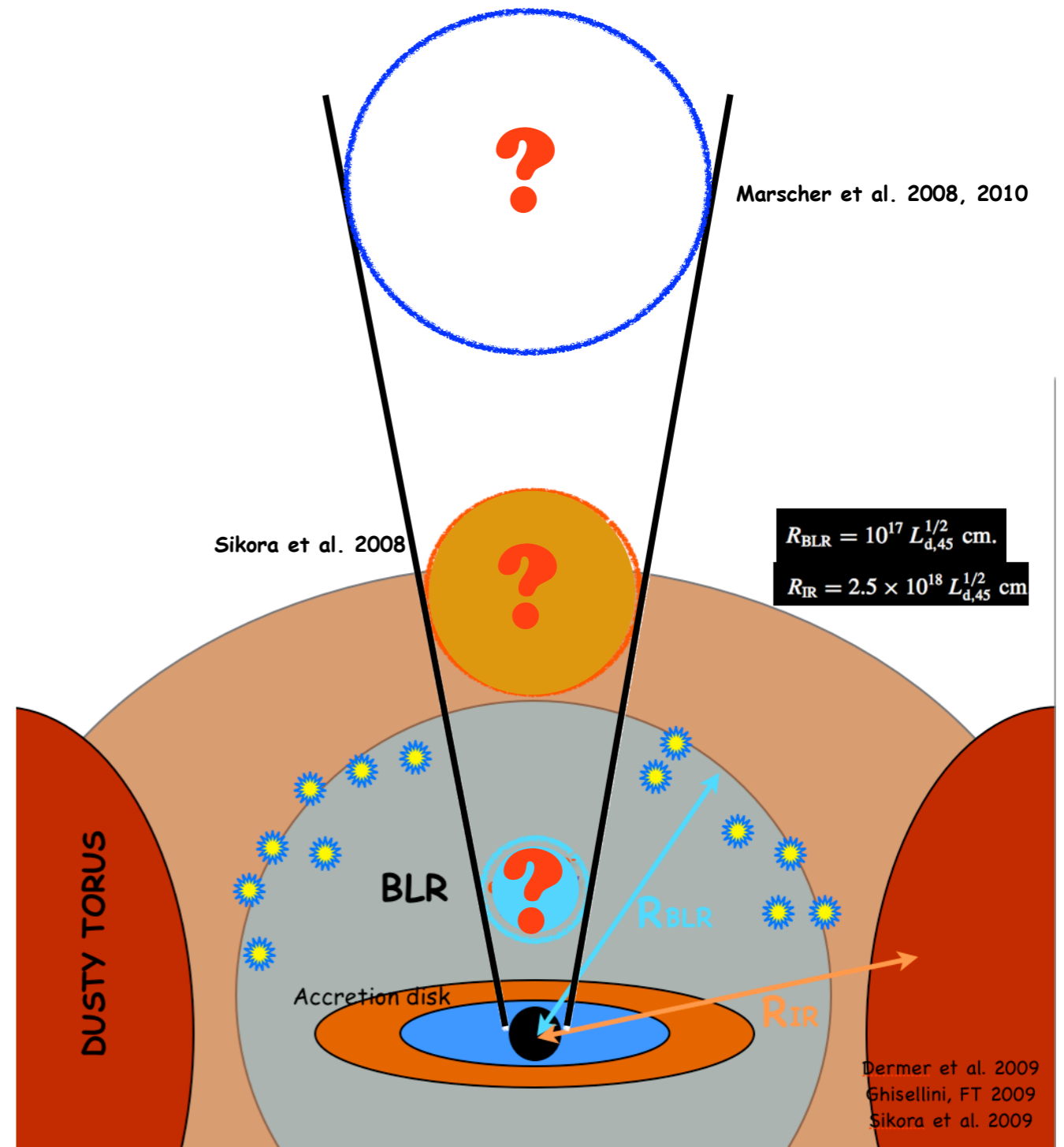


Blazars: current discussions

BL Lacs: "naked" jets

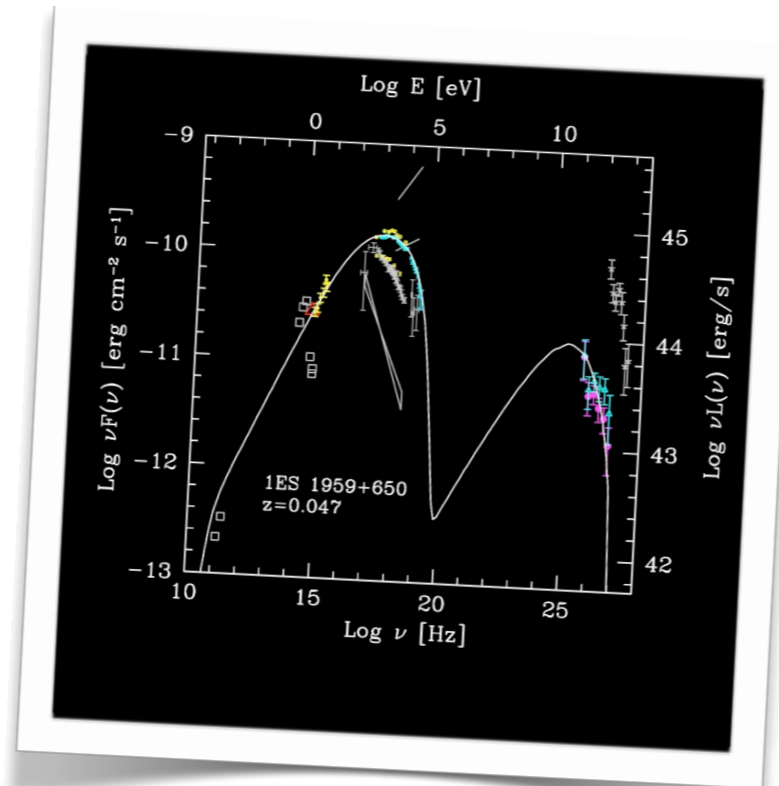
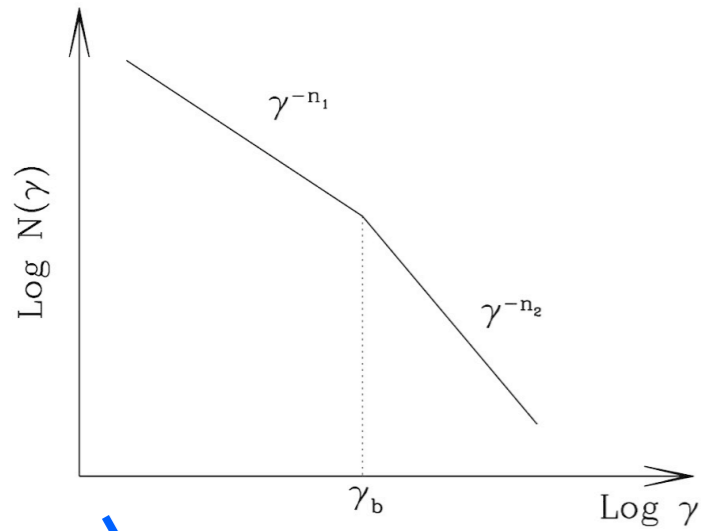


FSRQ: "dressed" jets

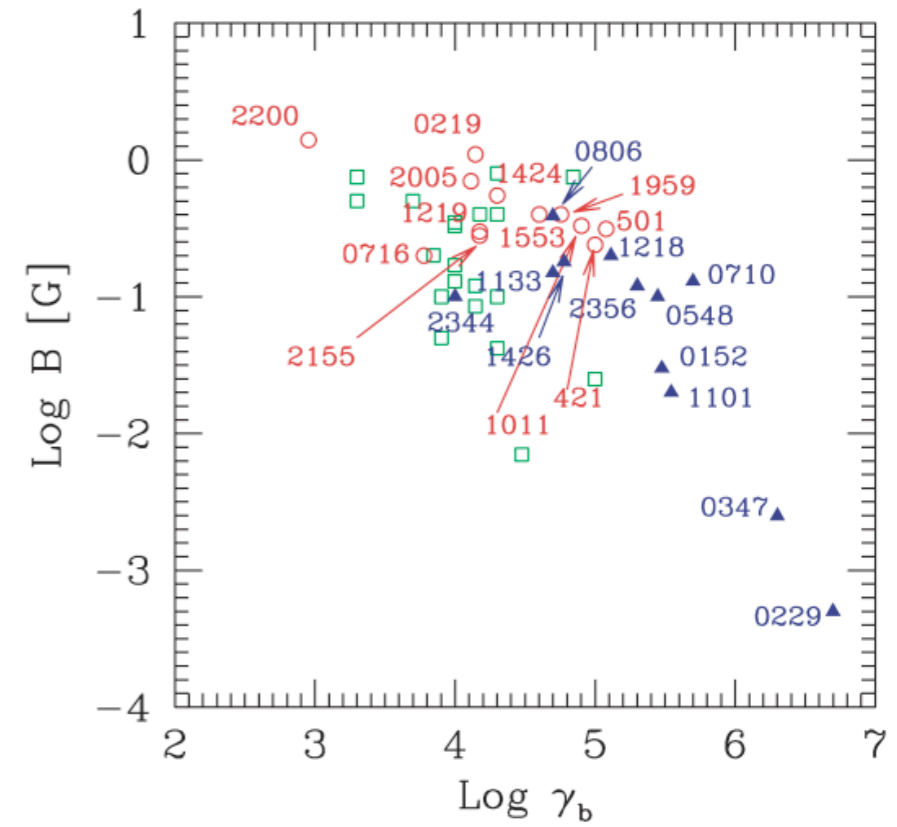


L. Pacciani talk

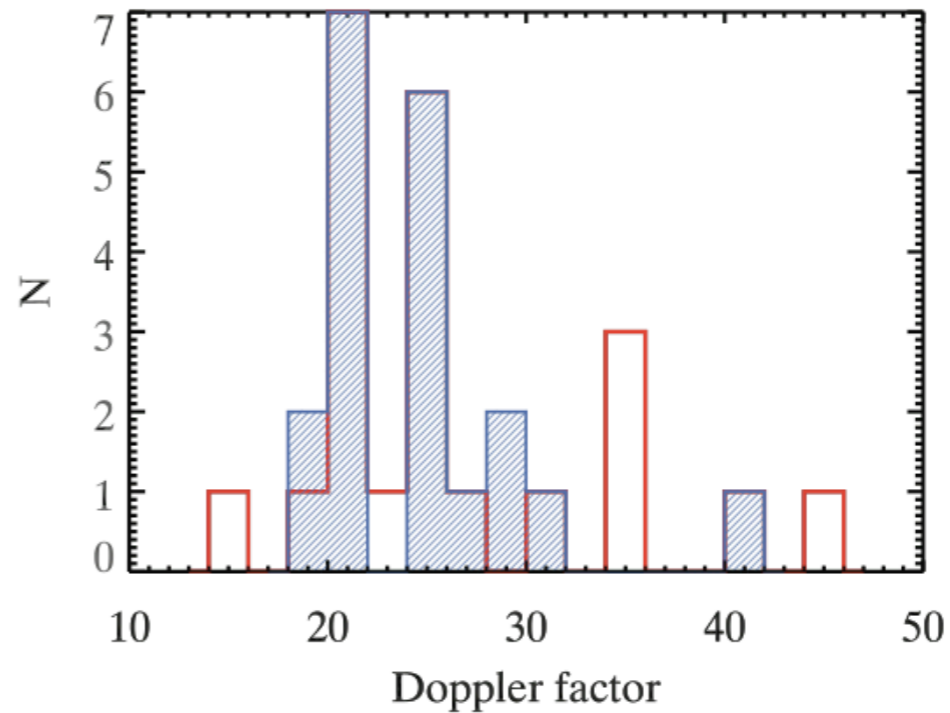
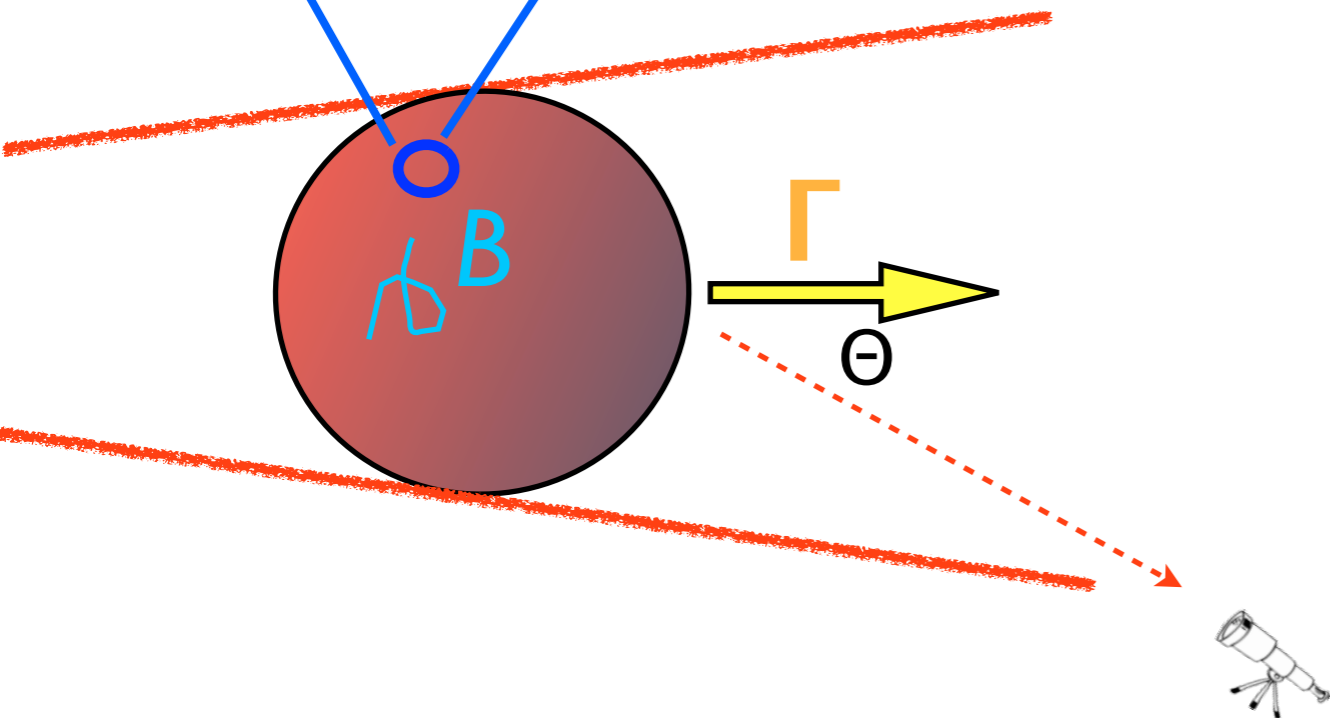
BL Lacs: the one zone model



Tagliaferri et al. 2009

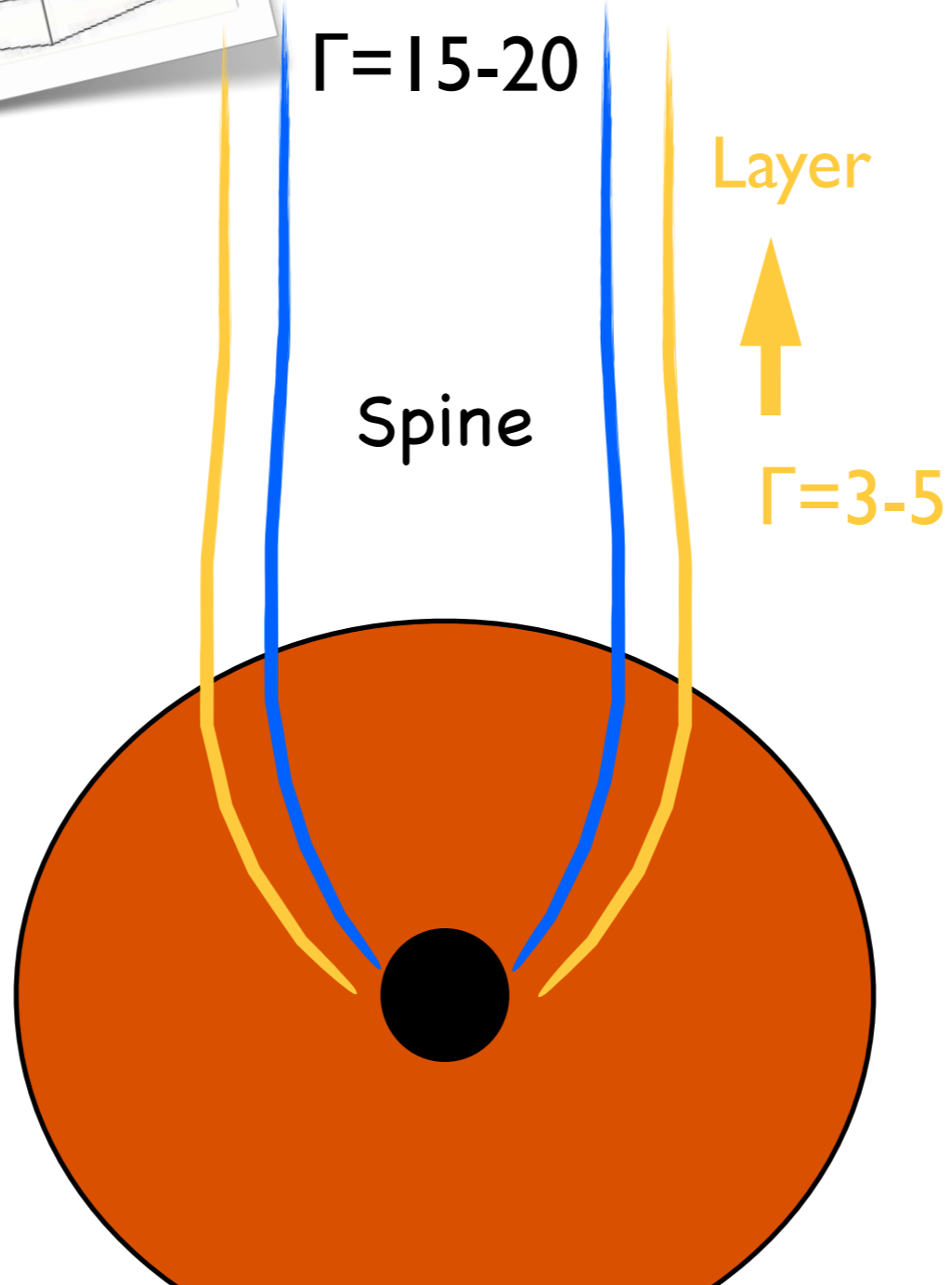
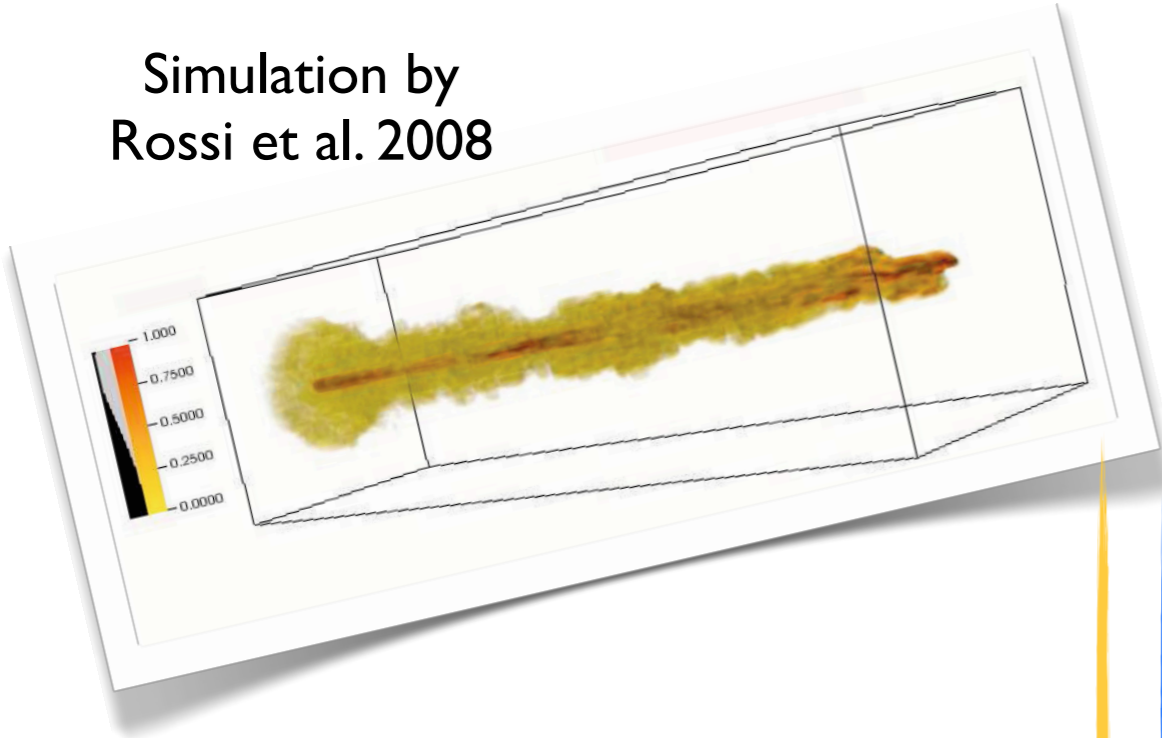


FT et al. 2010

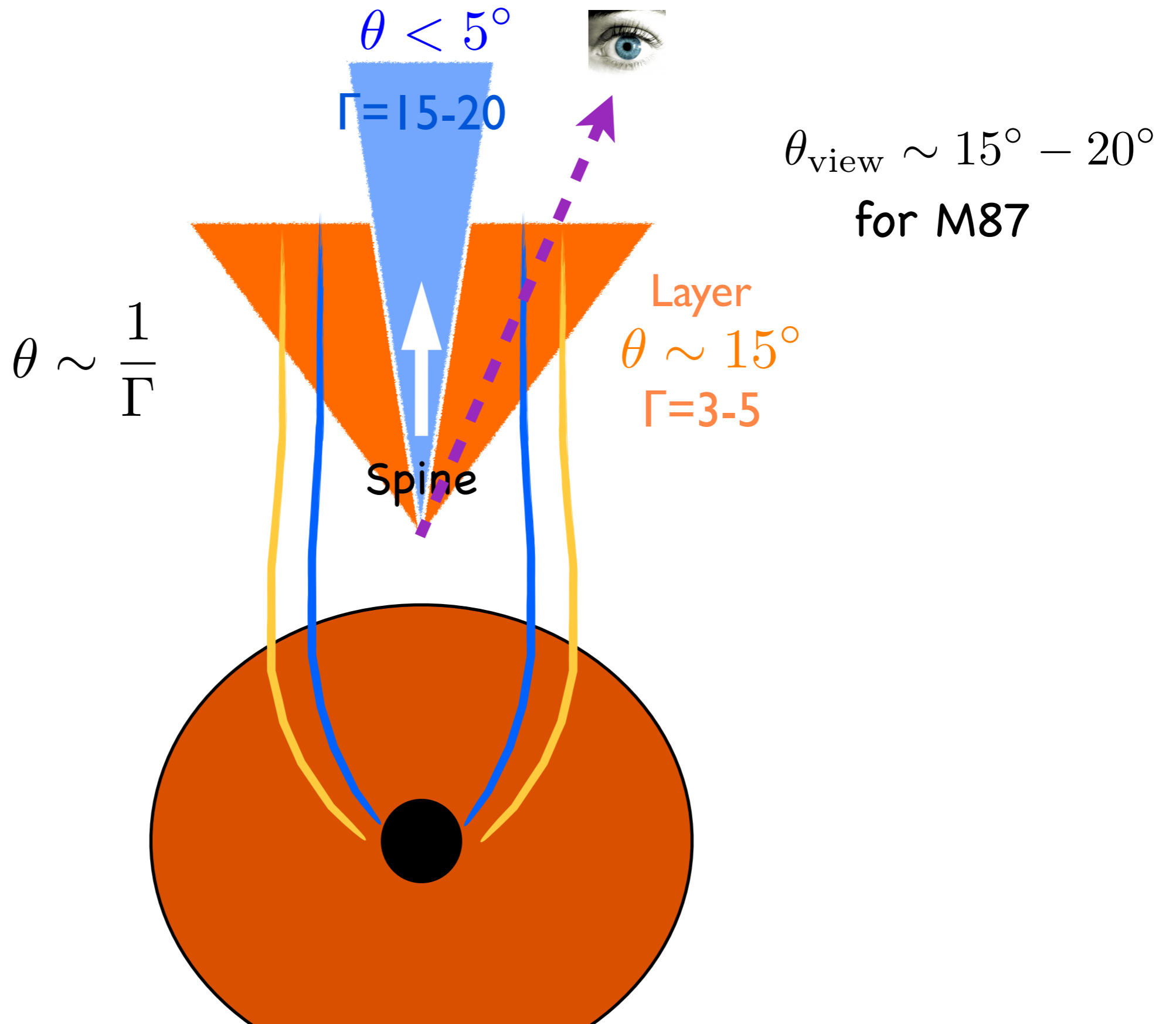


BL Lacs: the structured jet model

Simulation by
Rossi et al. 2008



Structured jets

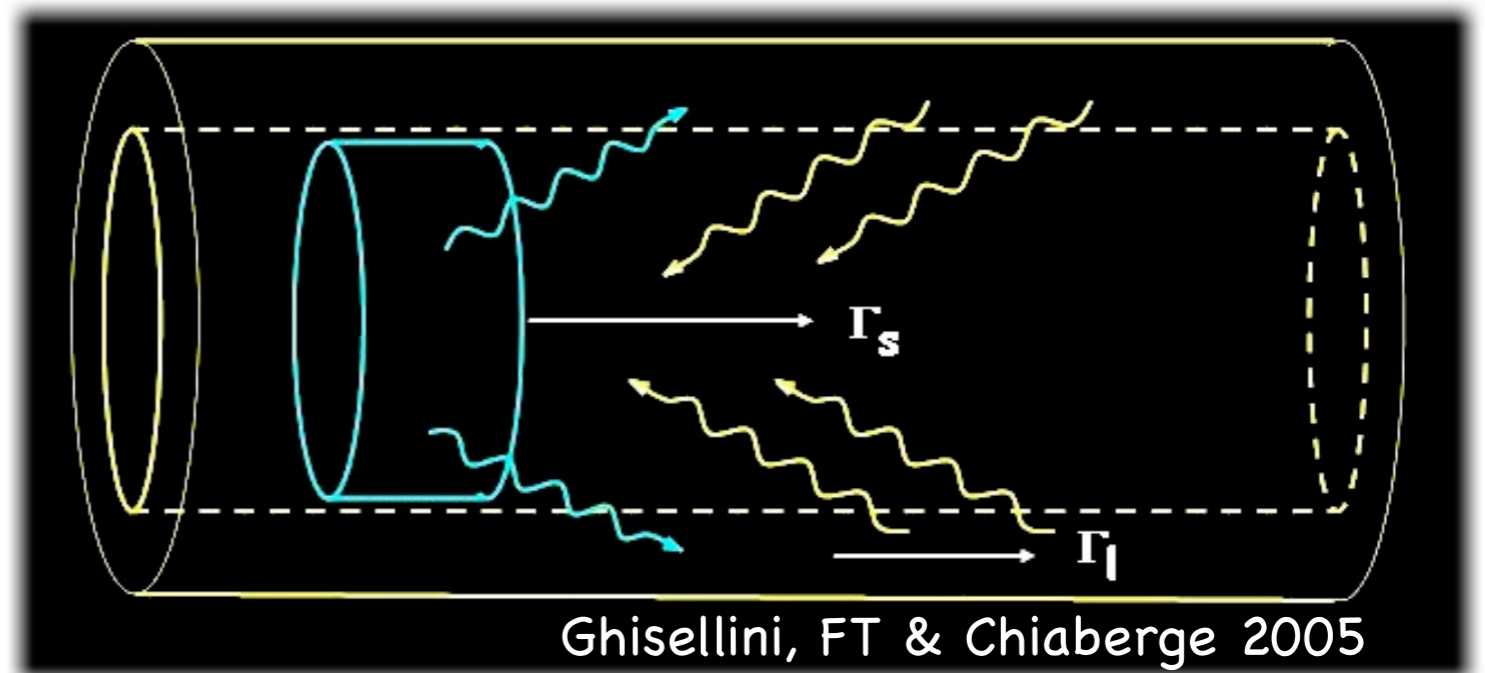


Structured jets

$$\Gamma_{\text{rel}} = \Gamma_s \Gamma_l (1 - \beta_s \beta_l)$$

$$U' \simeq U \Gamma_{\text{rel}}^2$$

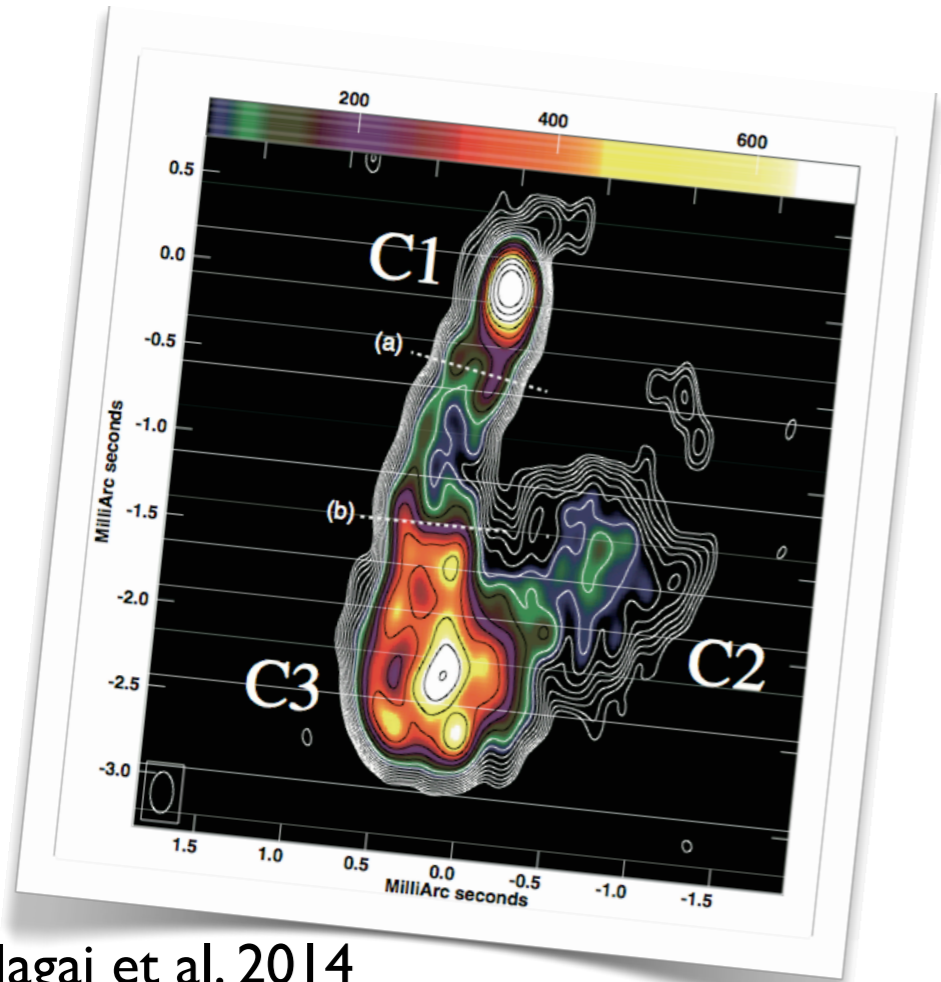
$$L_{\text{IC}} \propto U'$$



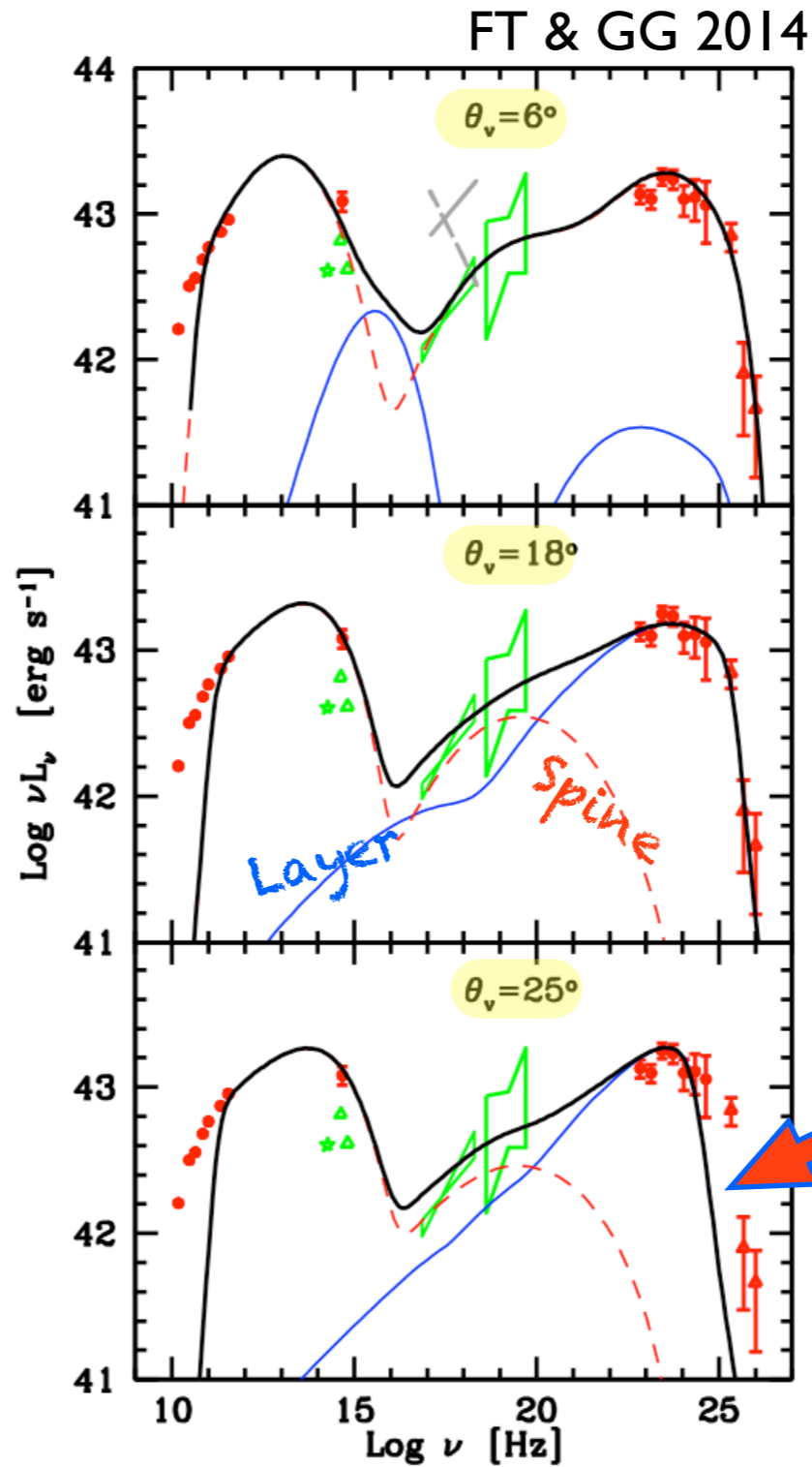
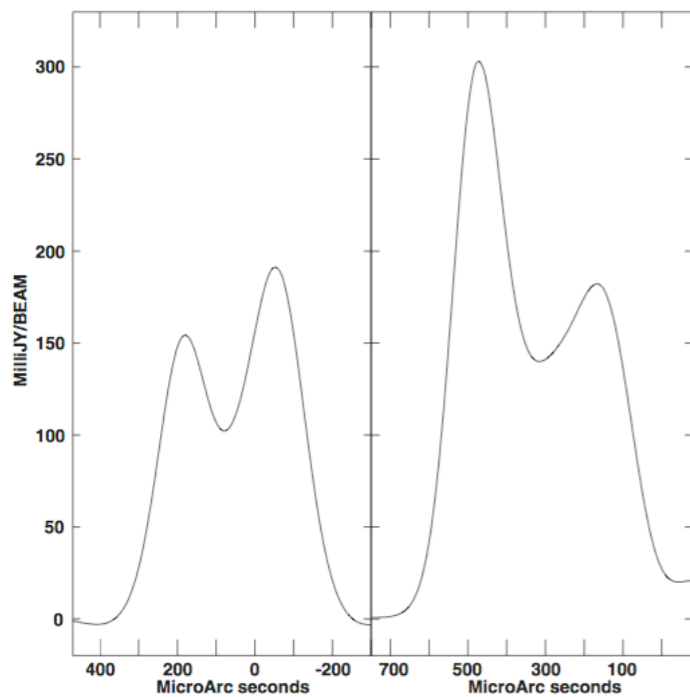
- ★ The **spine** sees an enhanced U_{rad} coming from the **layer**
- ★ Also the **layer** sees an enhanced U_{rad} coming from the **spine**

The IC emission is enhanced w.r.t. to the one-zone model

Structured jets: radiogalaxies



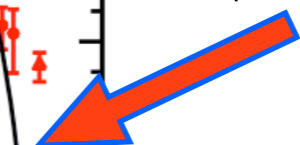
Nagai et al. 2014



NGC 1275

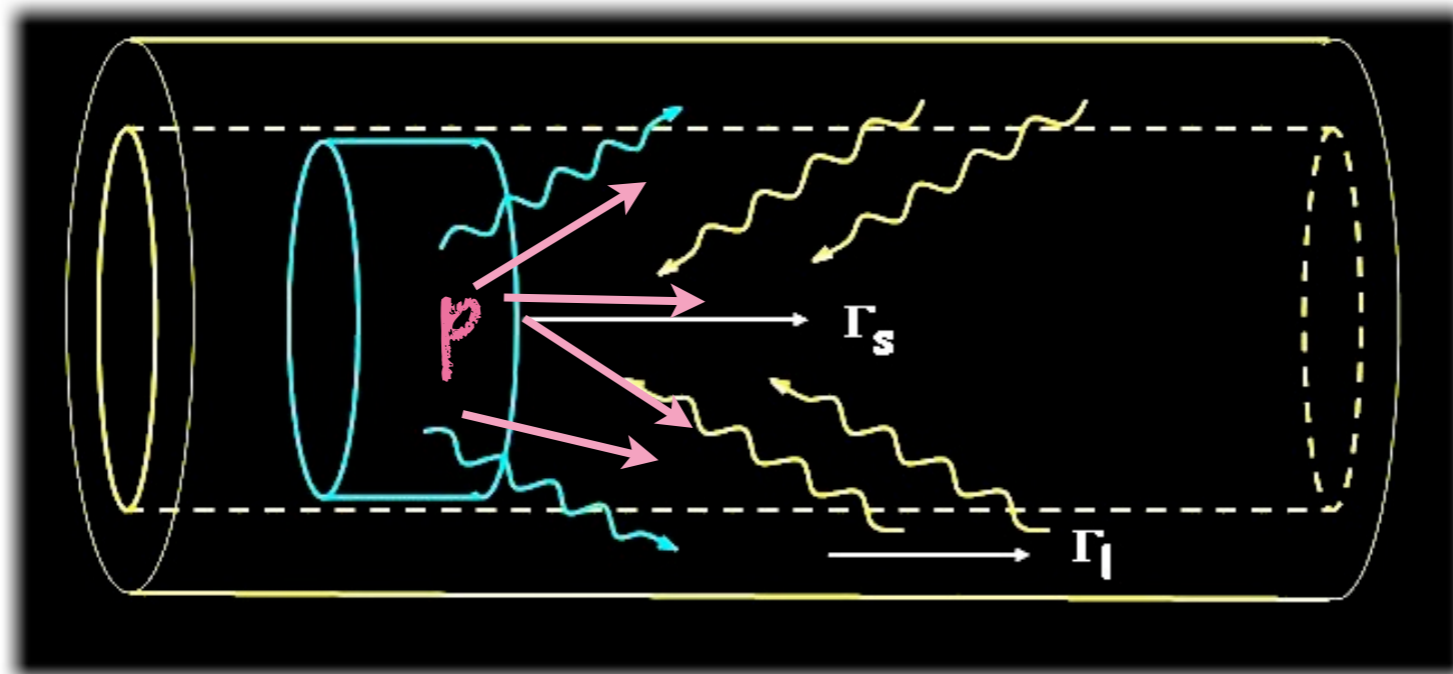
Absorbed $E > 50$ GeV

$\gamma\gamma \rightarrow e^\pm$



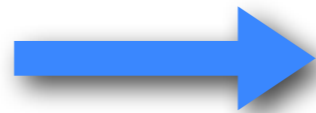
Structured jets: neutrinos

FT et al. 2014



$$p + \gamma \rightarrow n + \pi^+$$

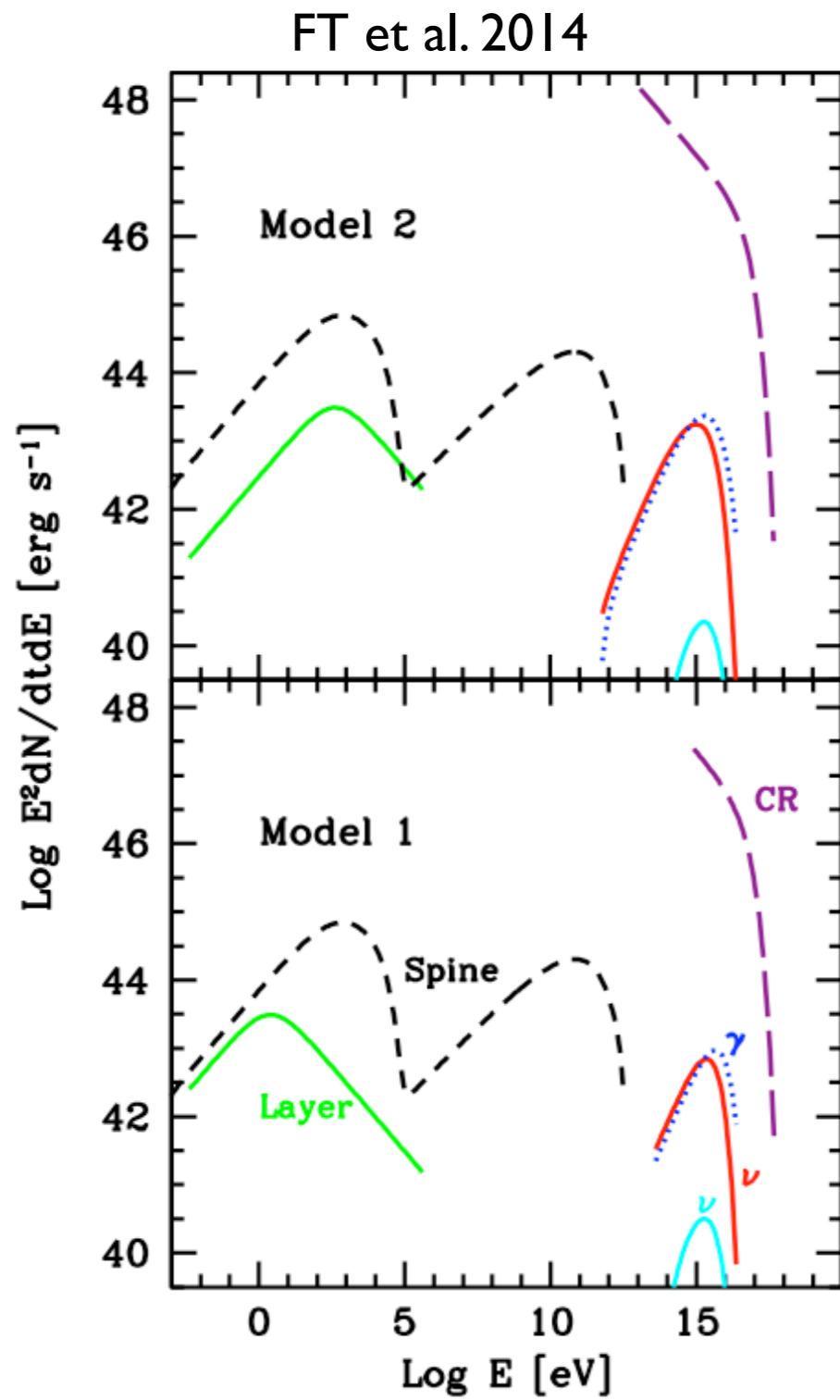
$$p + \gamma \rightarrow p + \pi^0$$



$$\pi^+ \rightarrow \mu^+ + \nu_\mu \rightarrow e^+ + \bar{\nu}_e + \bar{\nu}_\mu + \nu_\mu$$

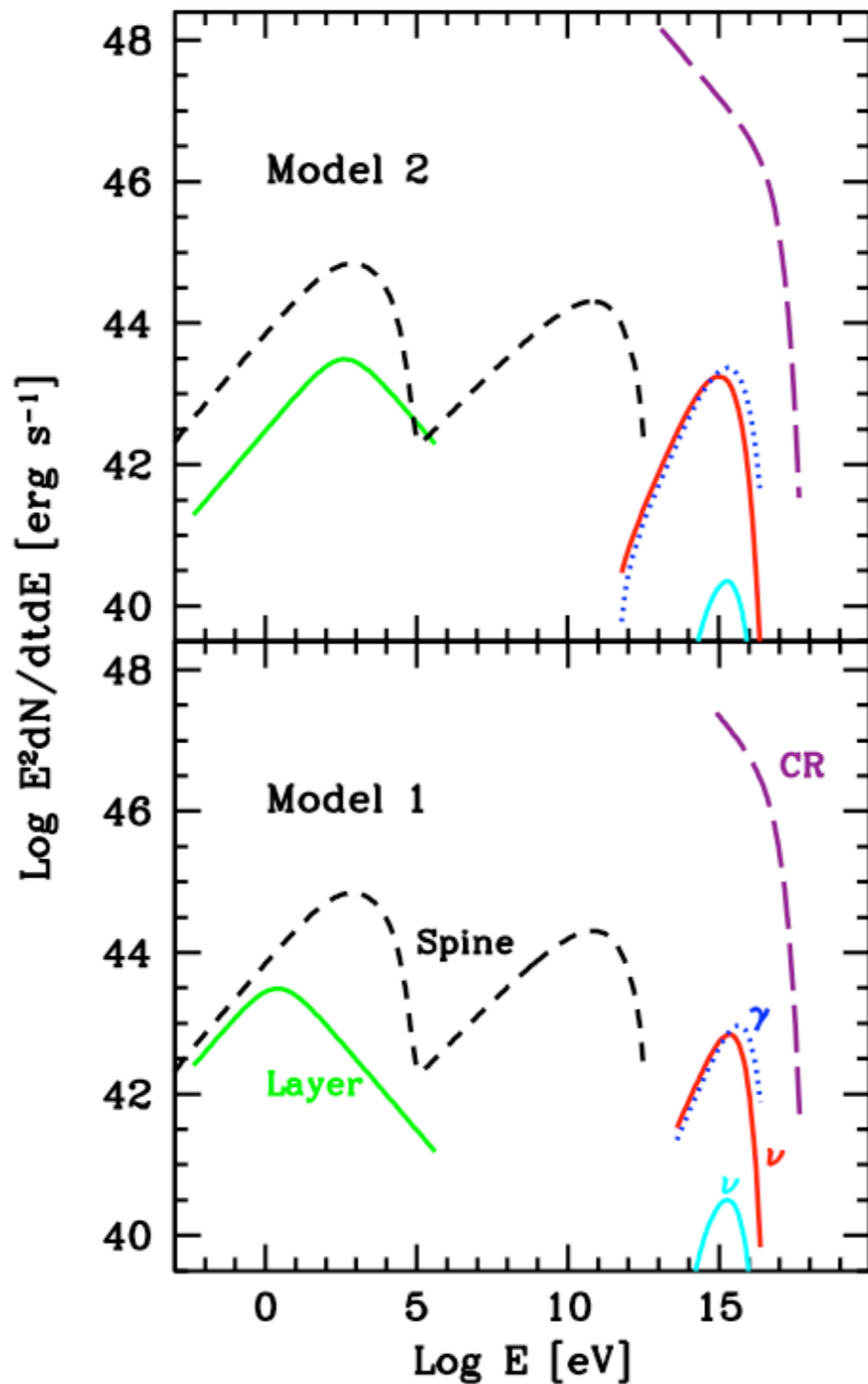
$$\pi^0 \rightarrow 2\gamma$$

Structured jets: neutrinos



Structured jets: neutrinos

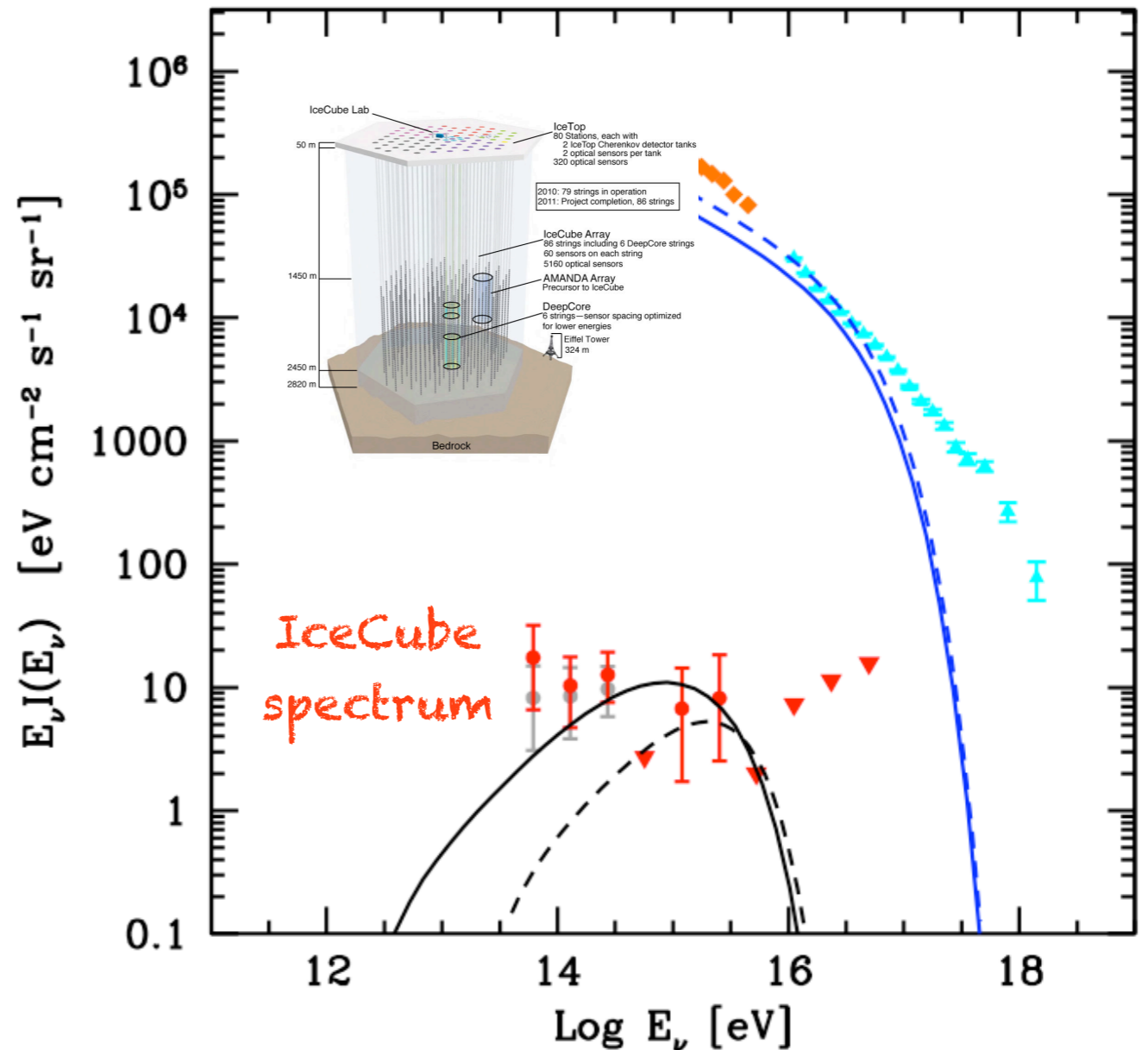
FT et al. 2014



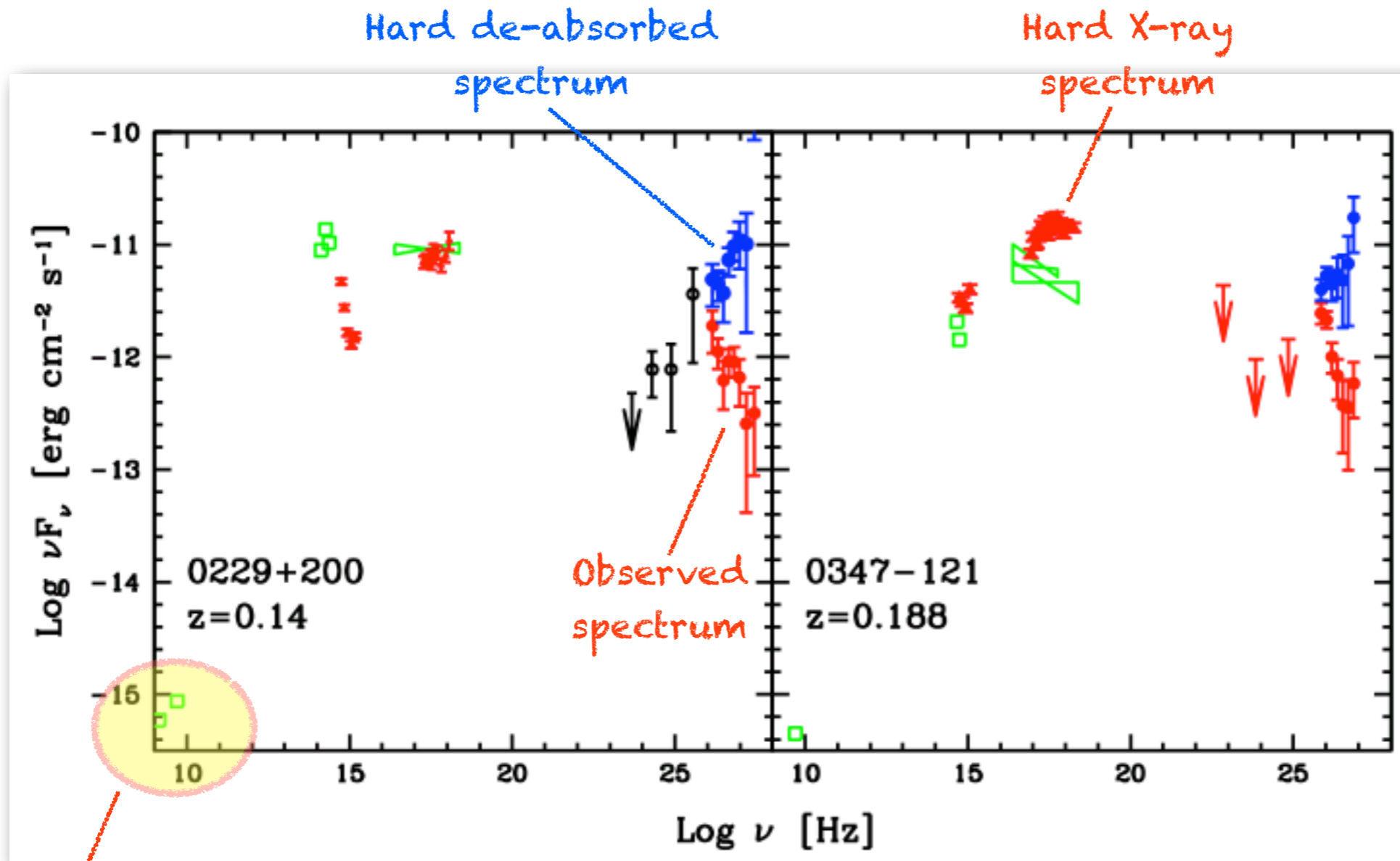
Cumulative emission
assuming a fast
evolution of HBL

$$N(z) = N_o(1 + z)^{-6}$$

Ajello et al. 2014



Extreme BL Lacs

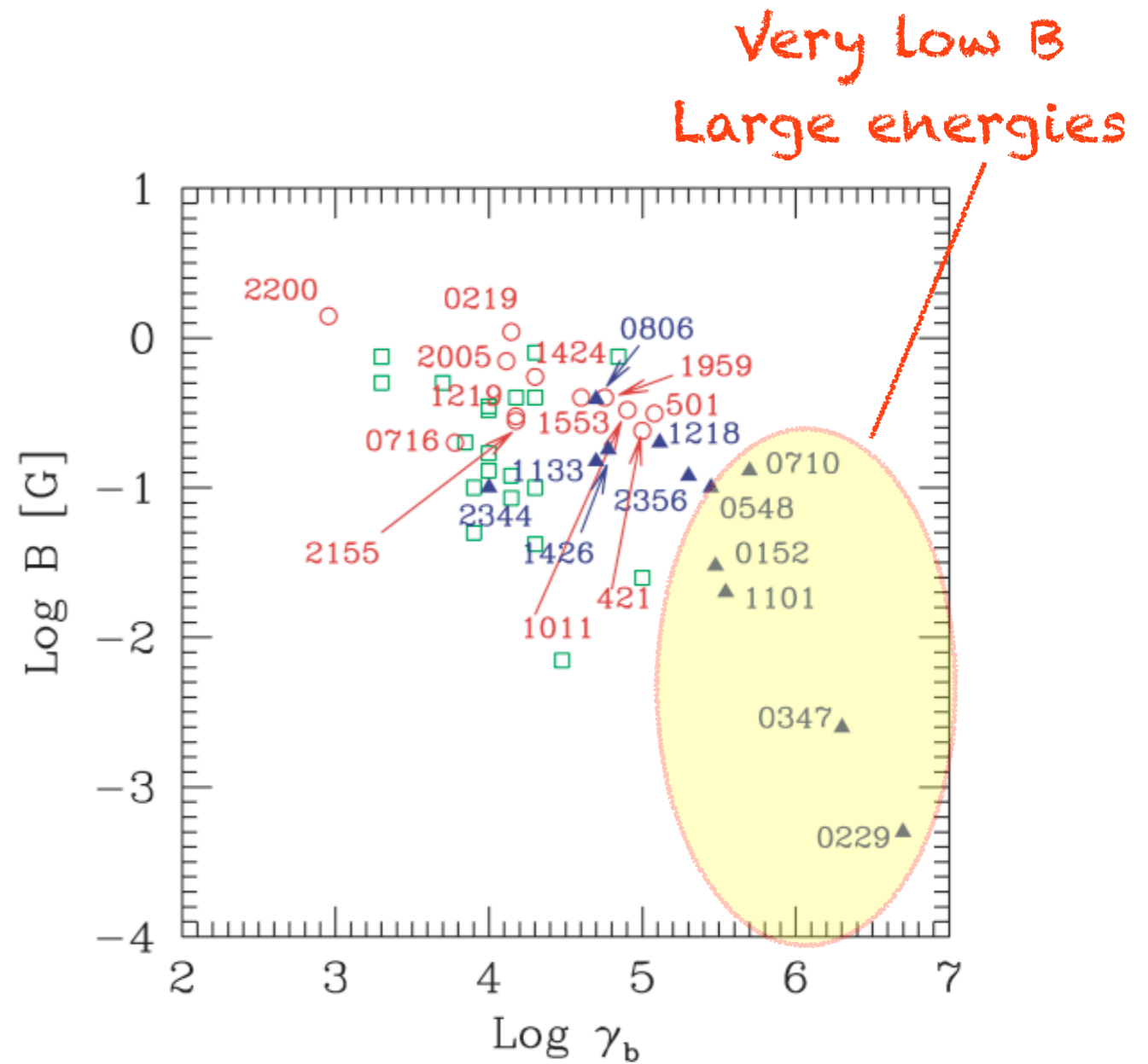
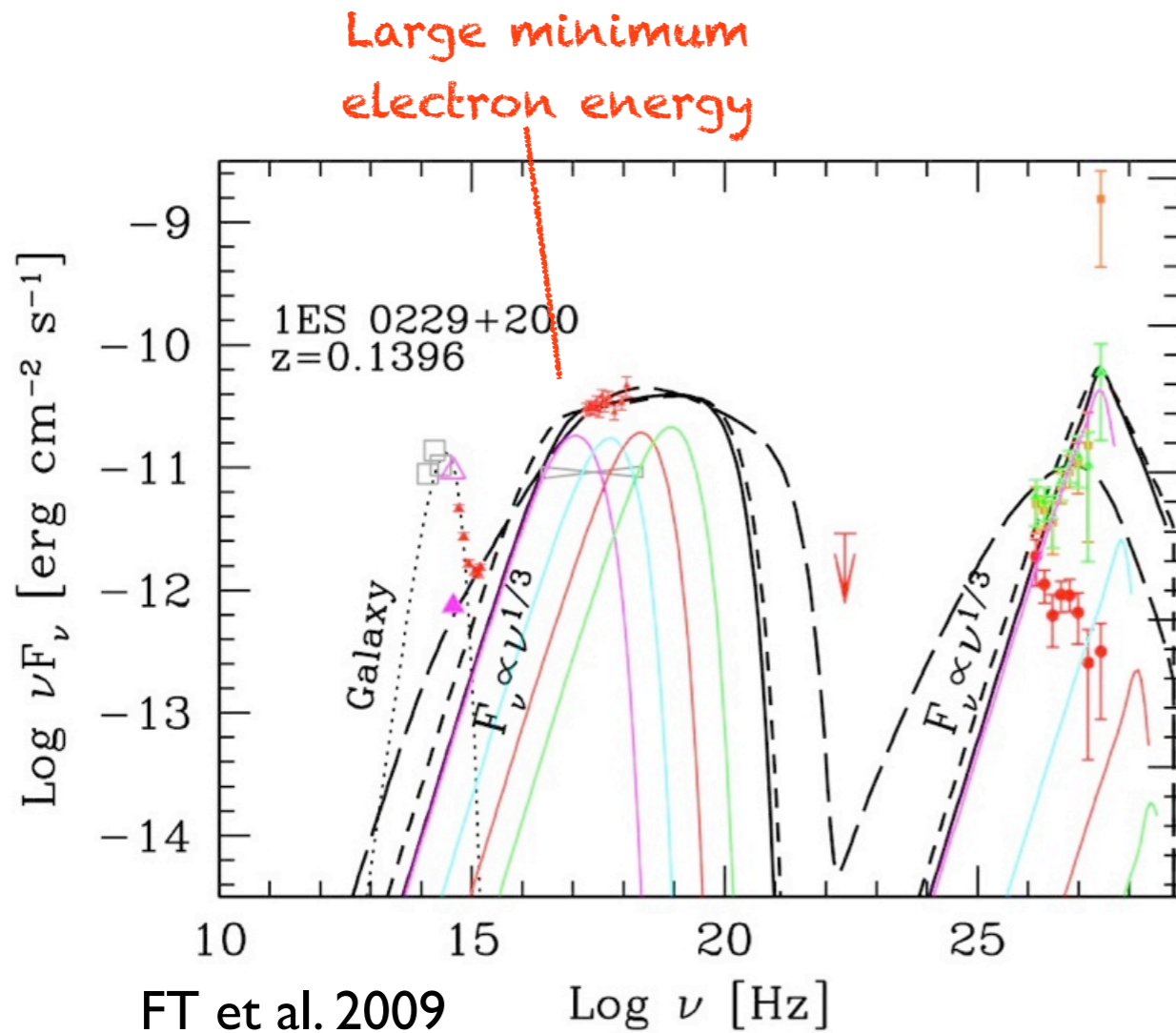


Very low radio power
(very compact)

Very hard X-ray and gamma-ray
(deabsorbed) spectra
Rather modest variability

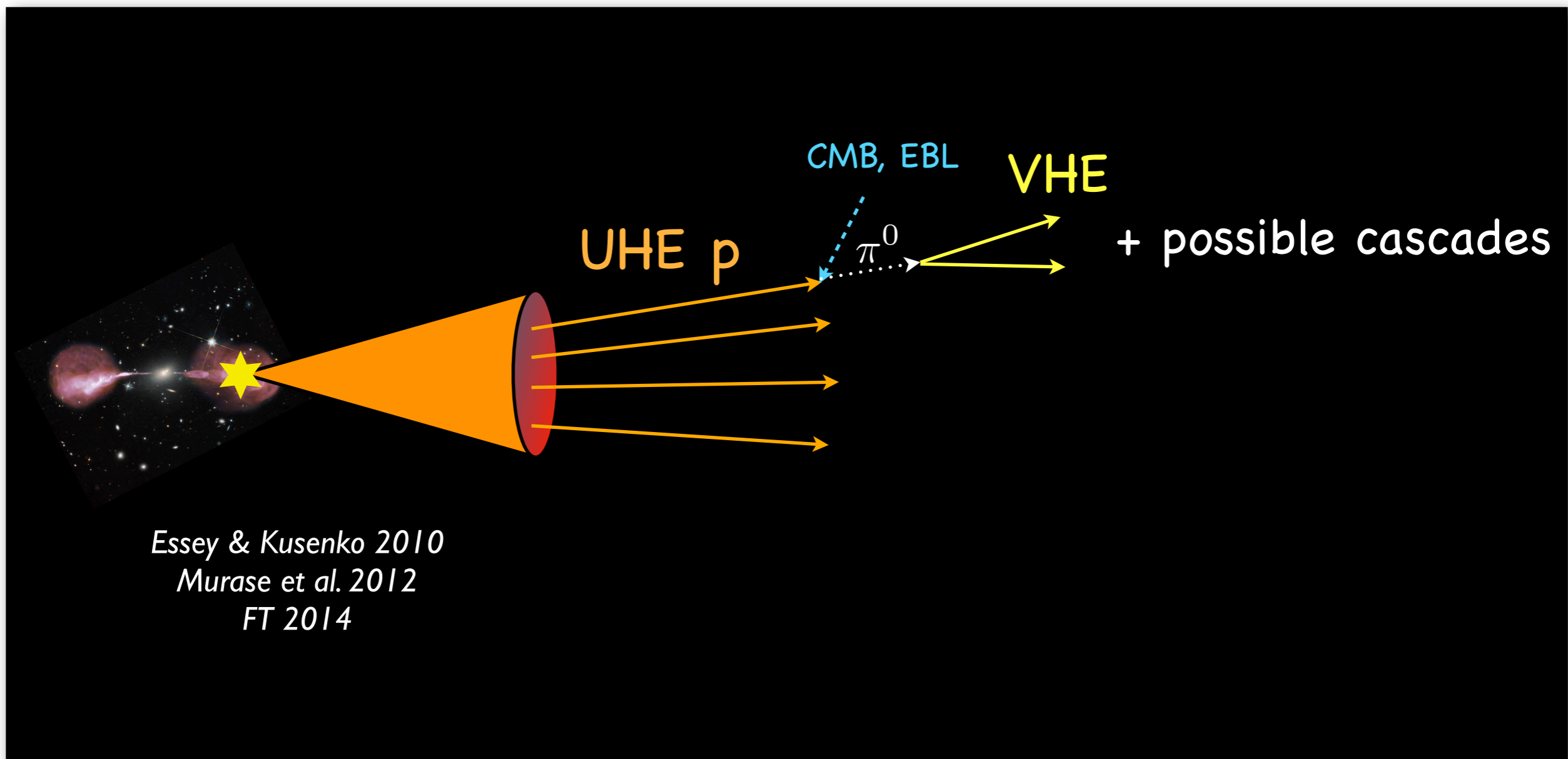
Extreme BL Lacs

after Costamante et al. 2001

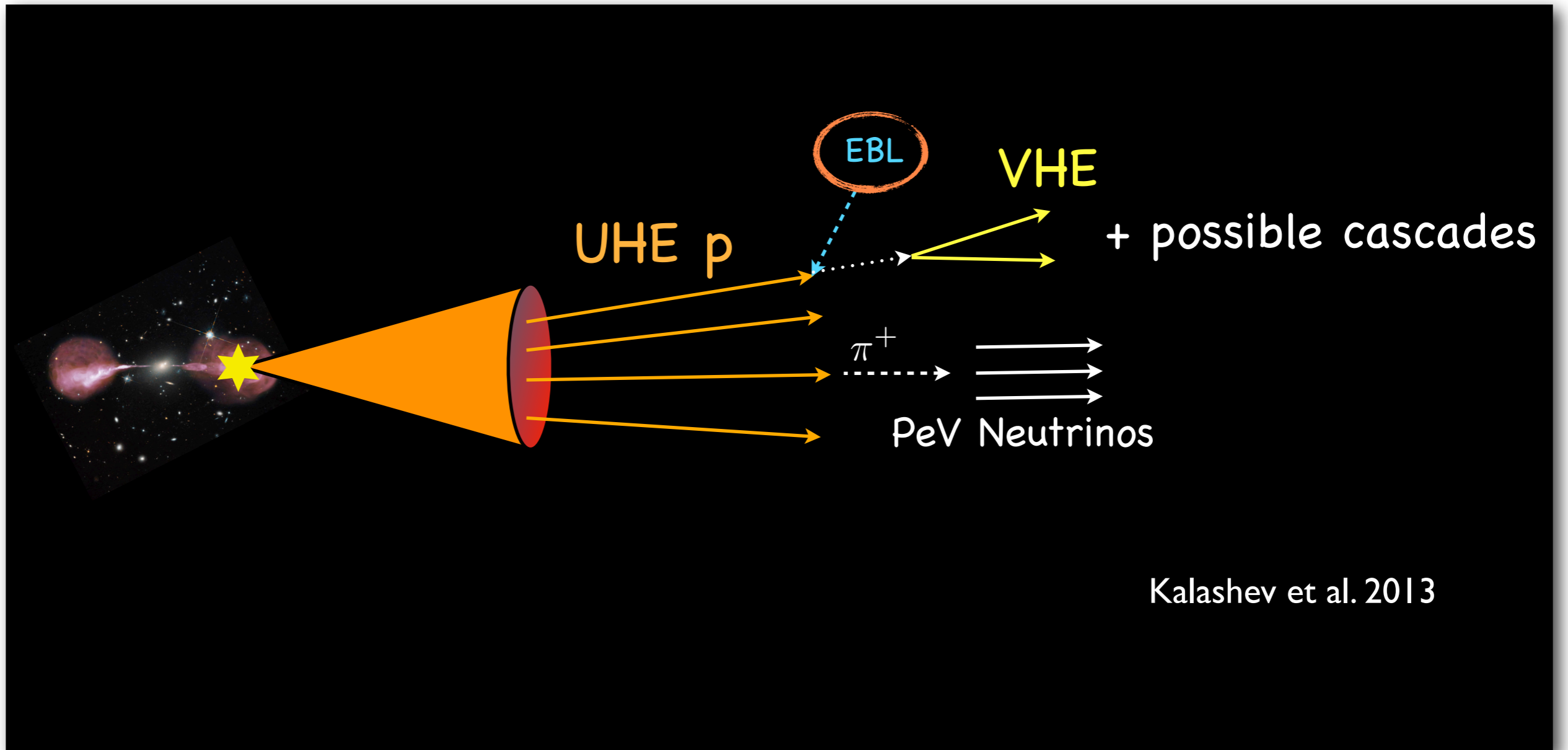


Acceleration process?
Why slowly variable?

Proton beams?



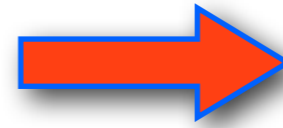
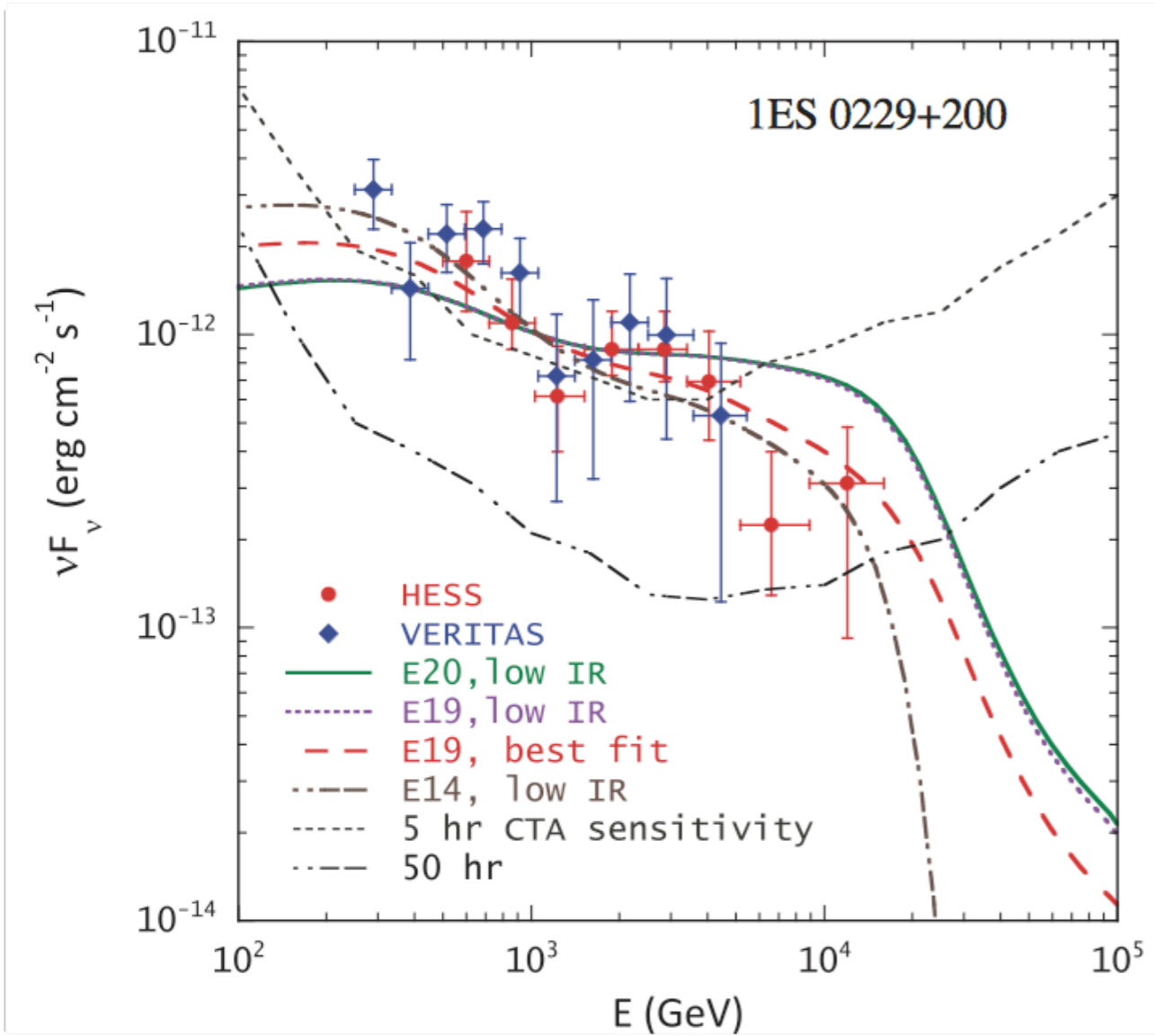
Proton beams?



Kalashov et al. 2013

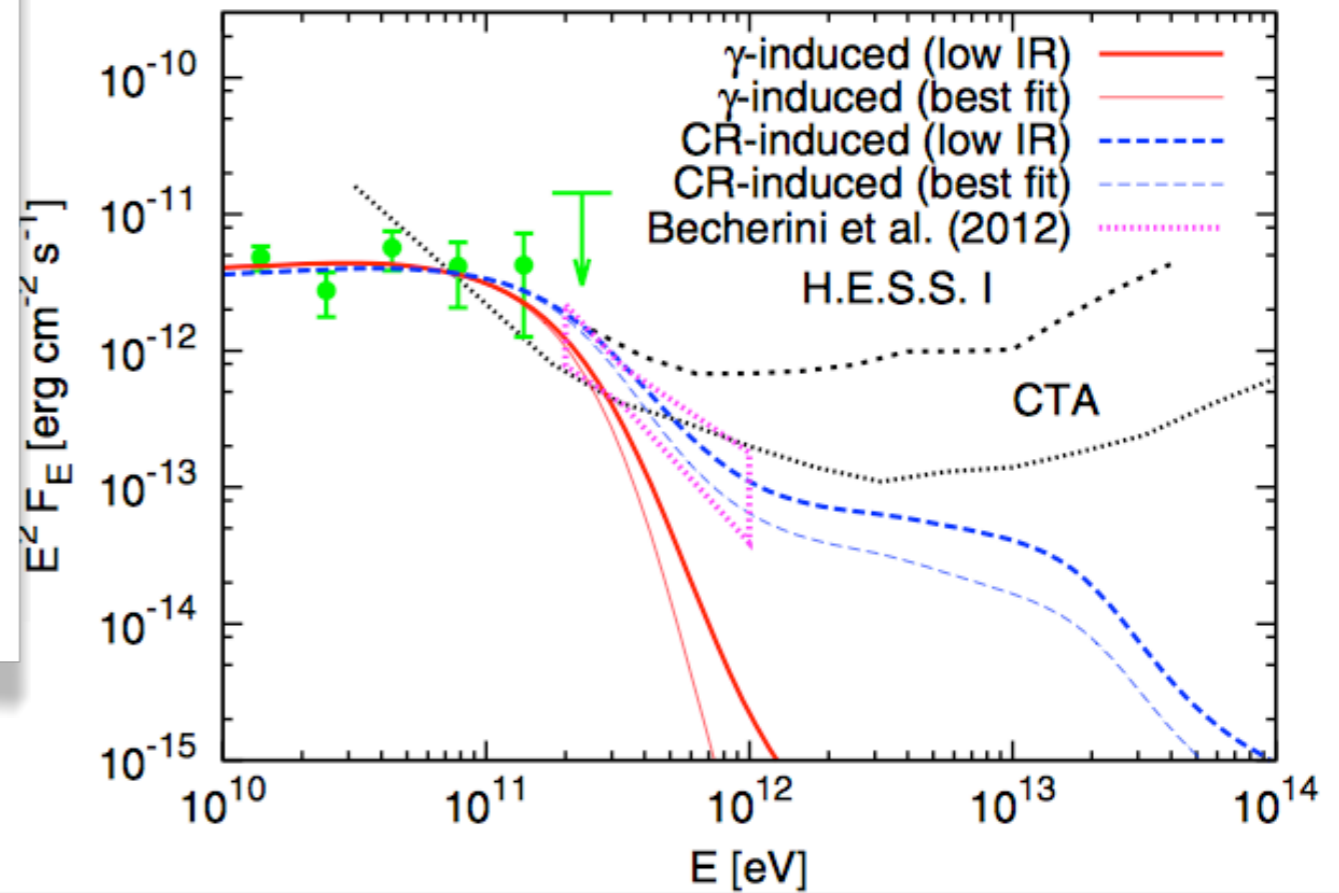
Proton beams?

Murase et al. 2012



To be observed with CTA

Takami et al. 2013



Misaligned EHBL: UHECR sources?

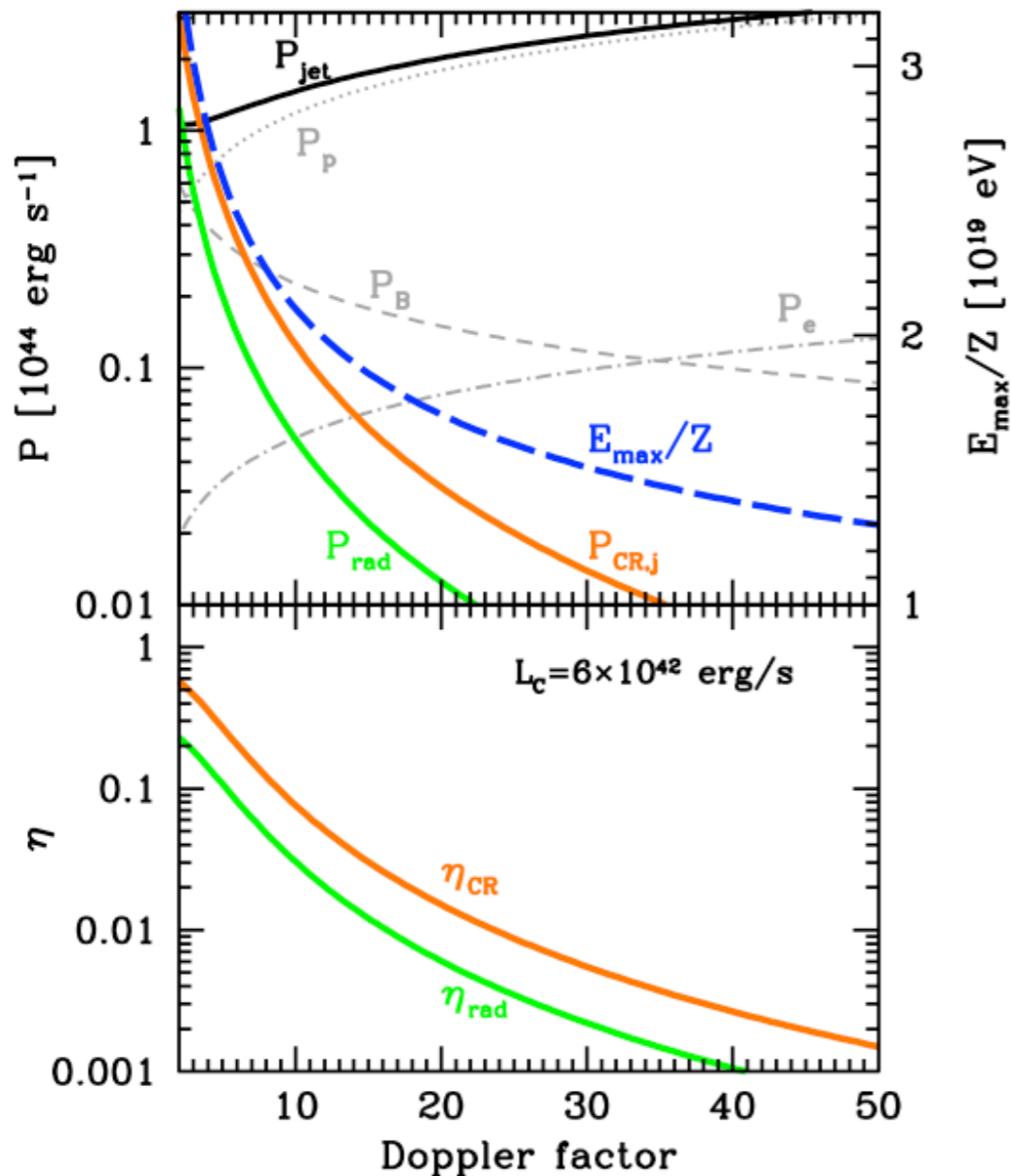
$$E_{UHECR} > 10^{19} \text{ eV}$$

NO EHBL within the
GZK radius (~ 100 Mpc)

$$E_{\text{max}} = ZeBR\Gamma \quad \leftarrow \text{Emission model}$$

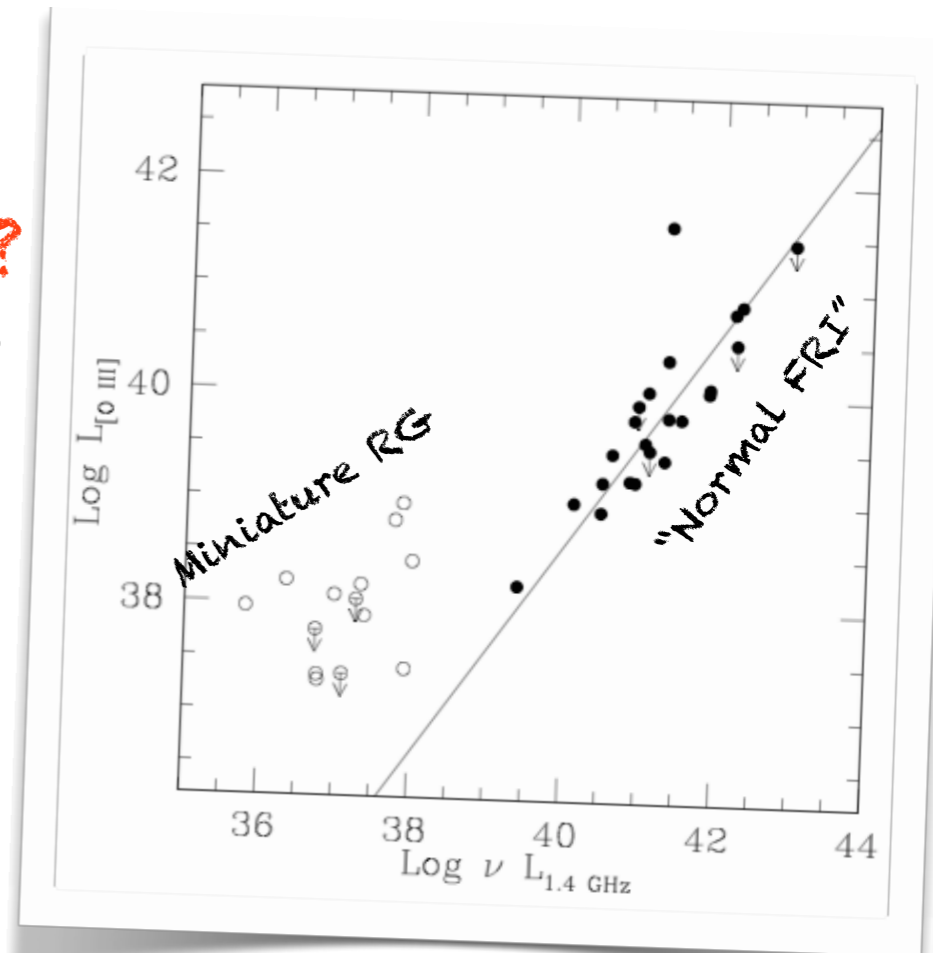
Parent population
(misaligned EHBL)?

$$L_r \lesssim 10^{40} \text{ erg s}^{-1}$$



"Miniature"
radio-galaxies?

Baldi & Capetti 2009



Outlook

Blazar jets: emission region(s)
dissipation/acceleration process(es) (magnetic, shock, turb.)

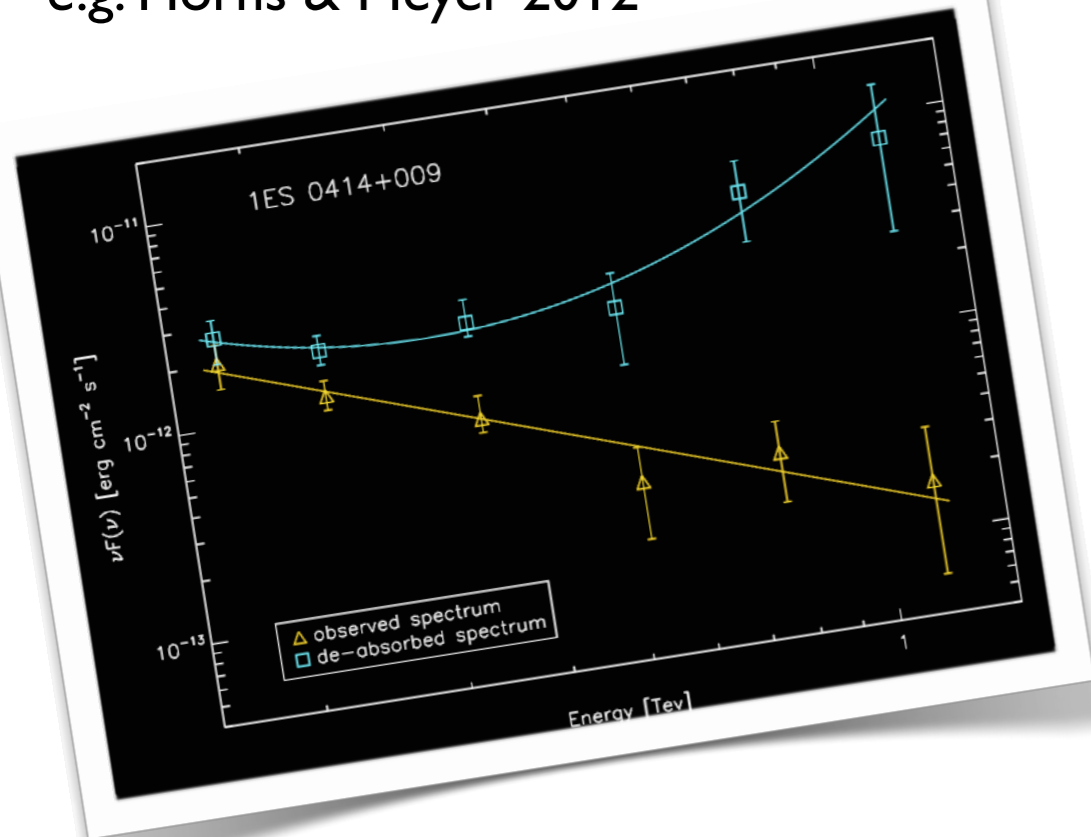
Structured jets: radiogalaxies
neutrinos
powerful jets (FR II)?

Extreme BL Lac: why are they so different?
UHECR sources?
parent population?

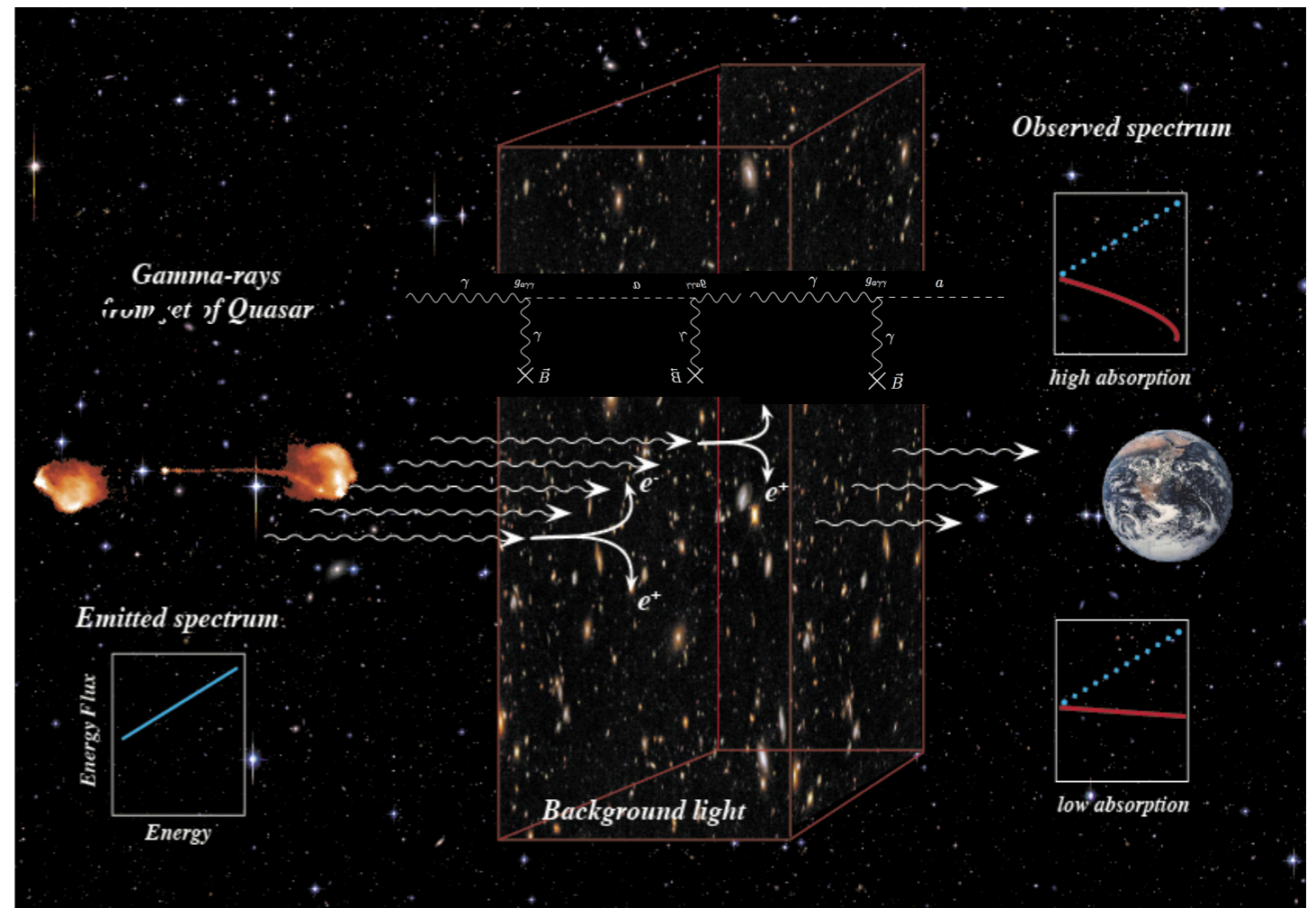
THANK YOU!

Cosmic opacity anomaly: ALP

e.g. Horns & Meyer 2012



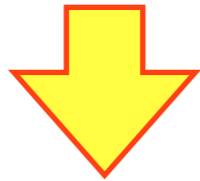
e.g. De Angelis et al. 2011



Cosmic opacity anomaly: LIV

LIV induces an effective mass for the photon

$$\beta_\gamma = 1 - \left(\frac{E_\gamma}{M_{LVn}} \right)^n \quad ; \quad m_\gamma^2 = -\frac{E_\gamma^{2+n}}{M_{LVn}^n},$$



Modification of threshold for pair production at high E

LIV induces suppression of EBL-opacity

Fairbairn+ 2014

ASTRI/CTA
Mini-Array

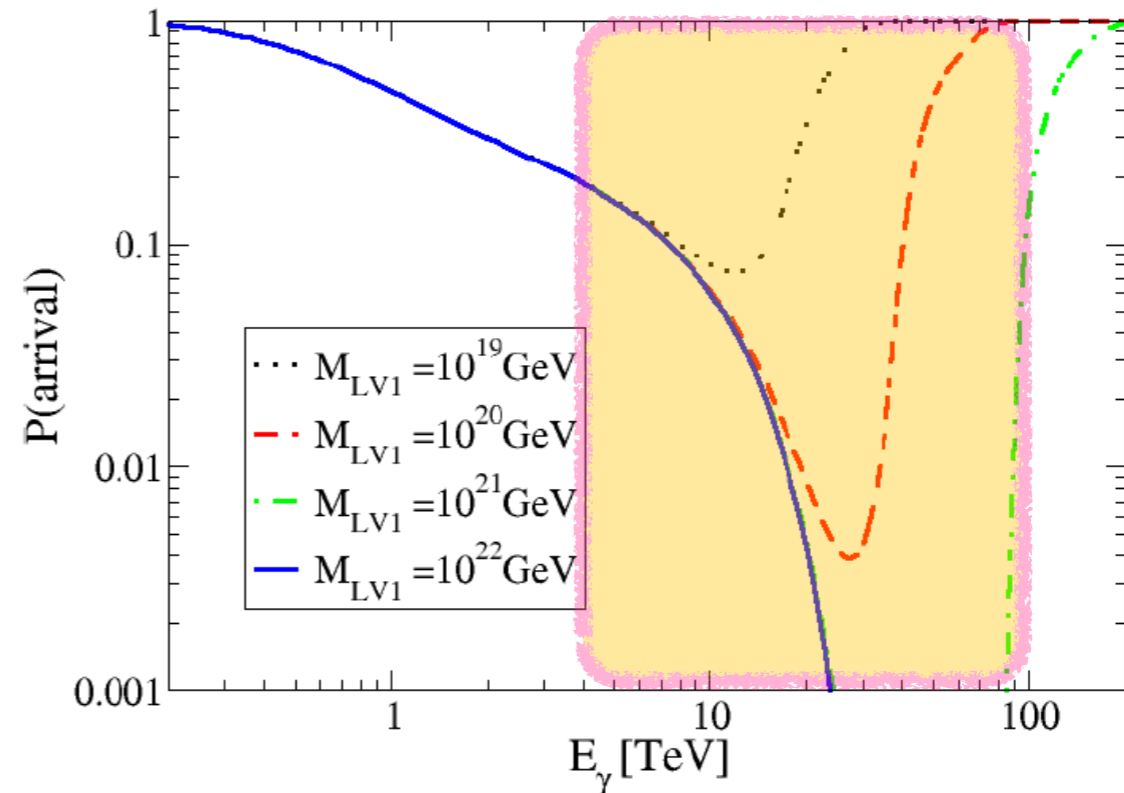
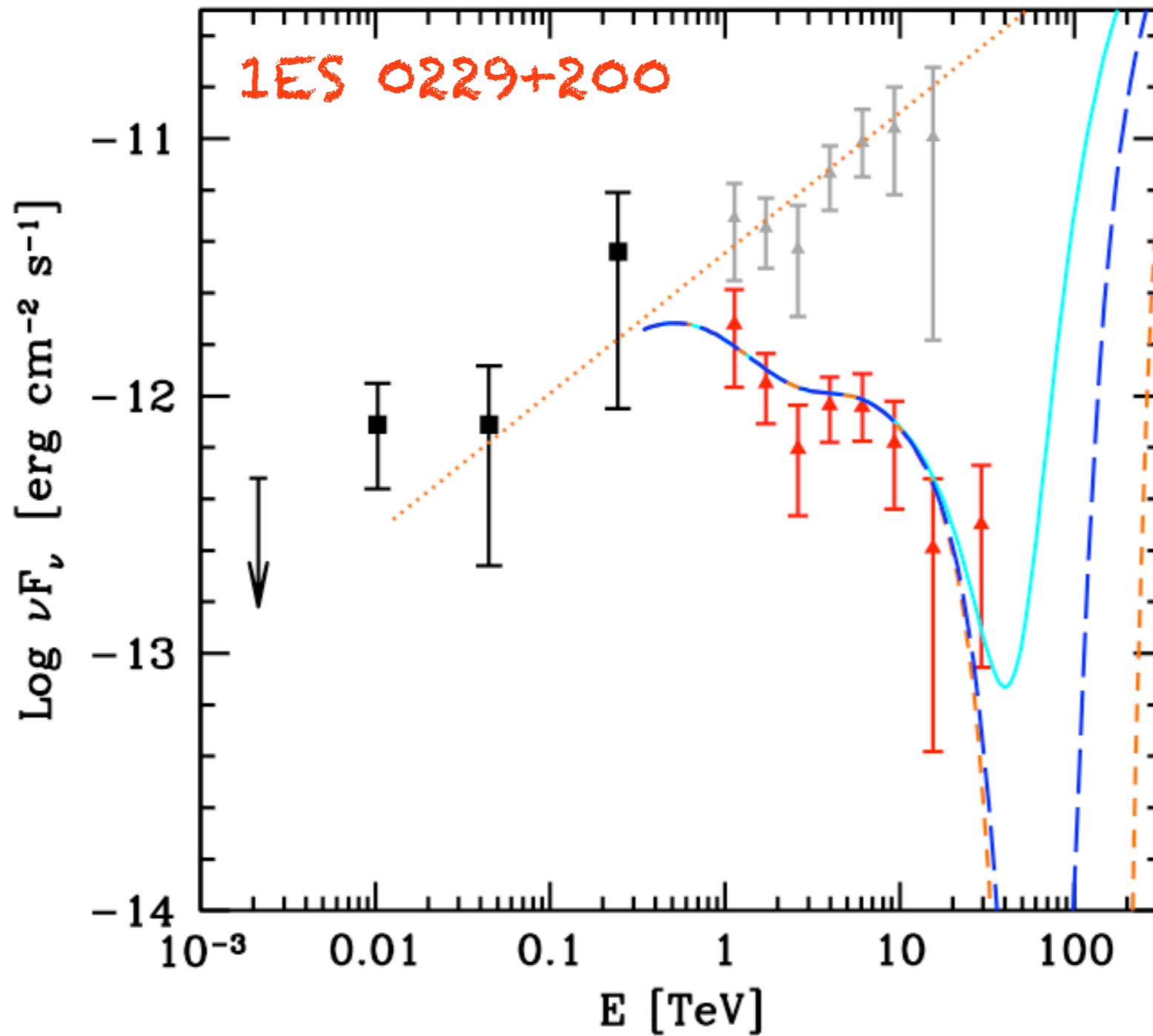


Figure 2. The arrival probability of a photon emitted from a hypothetical source at redshift $z = 0.05$ as a function of energy. The different curves represent different values of the Lorentz-violating scale M_{LV1} . VHE photons with energies $\gtrsim 100$ TeV can travel through the CMB effectively unimpeded.

Cosmic opacity anomaly: LIV



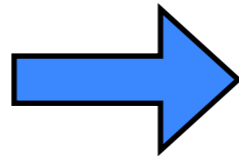
Intermezzo

neutrino production in a nutshell

relativistic protons

$$p + \gamma \rightarrow n + \pi^+$$

$$p + \gamma \rightarrow p + \pi^0$$



pion decay

$$\pi^+ \rightarrow \mu^+ + \nu_\mu \rightarrow e^+ + \bar{\nu}_e + \bar{\nu}_\mu + \nu_\mu$$

$$\pi^0 \rightarrow 2\gamma$$

$$\sigma_{p\gamma} \simeq 3.4 \times 10^{-28} \text{ cm}^2$$

$$E_{\text{th}} = \frac{2m_p m_\pi + m_\pi^2}{4\epsilon} \simeq 7 \times 10^{16} \left(\frac{\epsilon}{\text{eV}} \right)^{-1} \text{ eV}$$

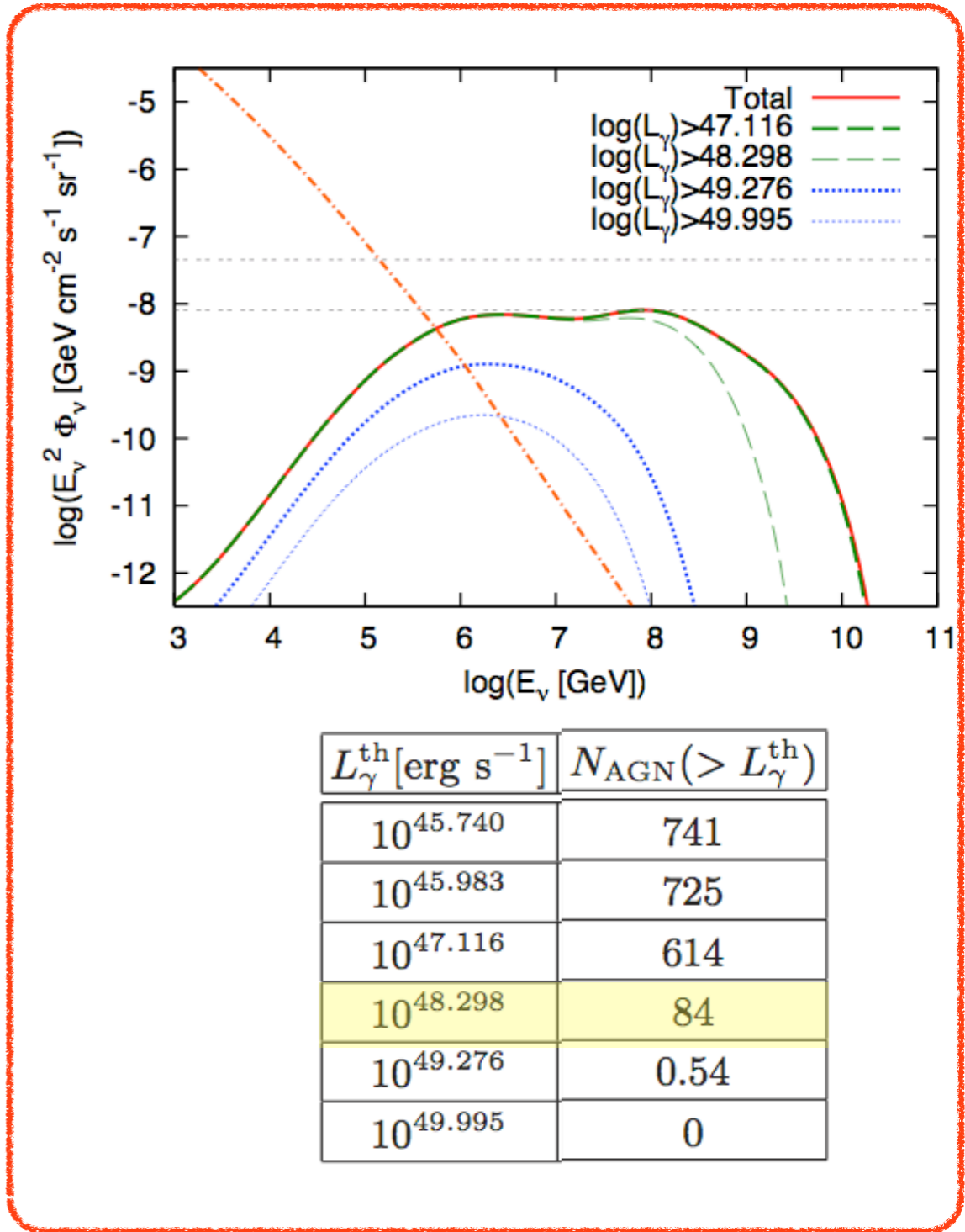
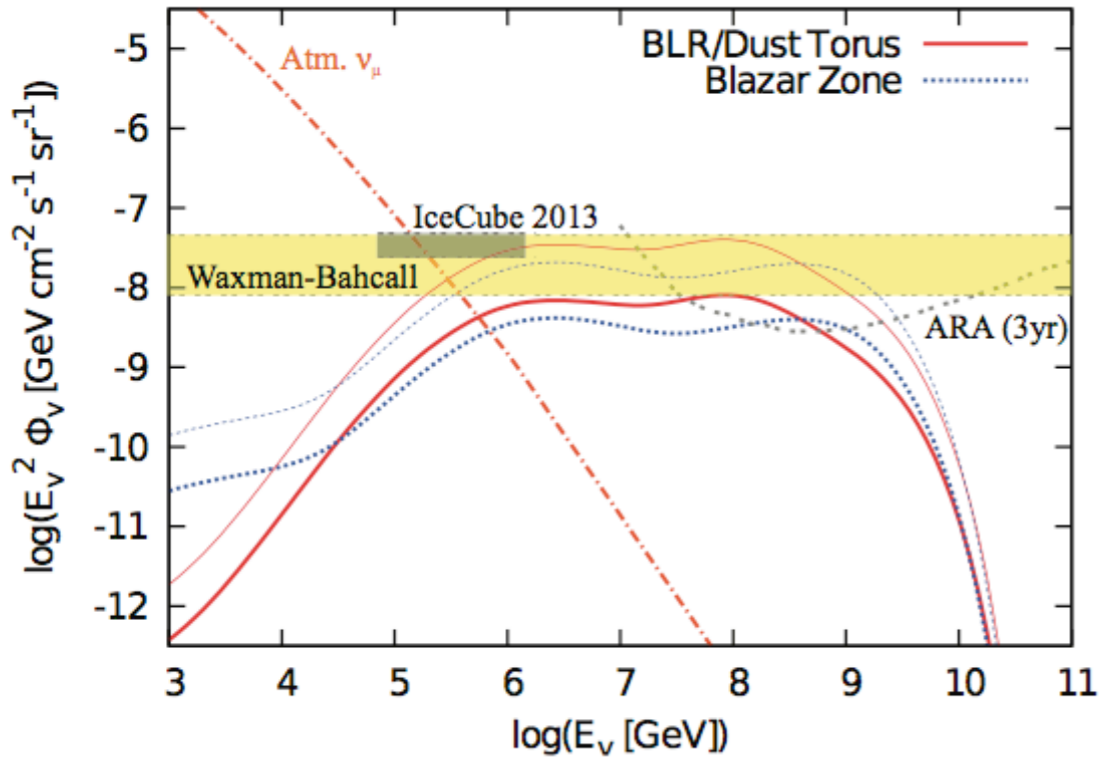
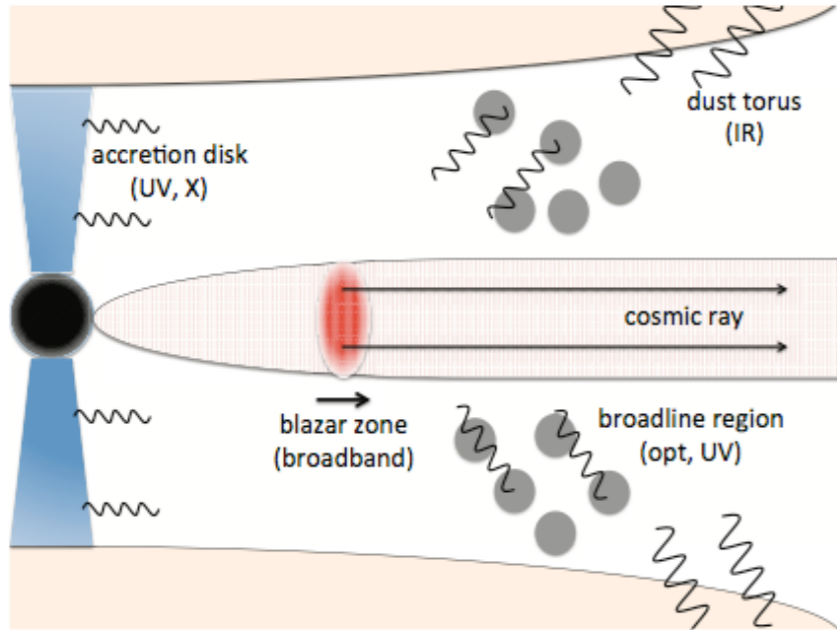
$$E_\nu \simeq 0.05 E_p$$

$$E_\gamma \simeq 0.1 E_p$$

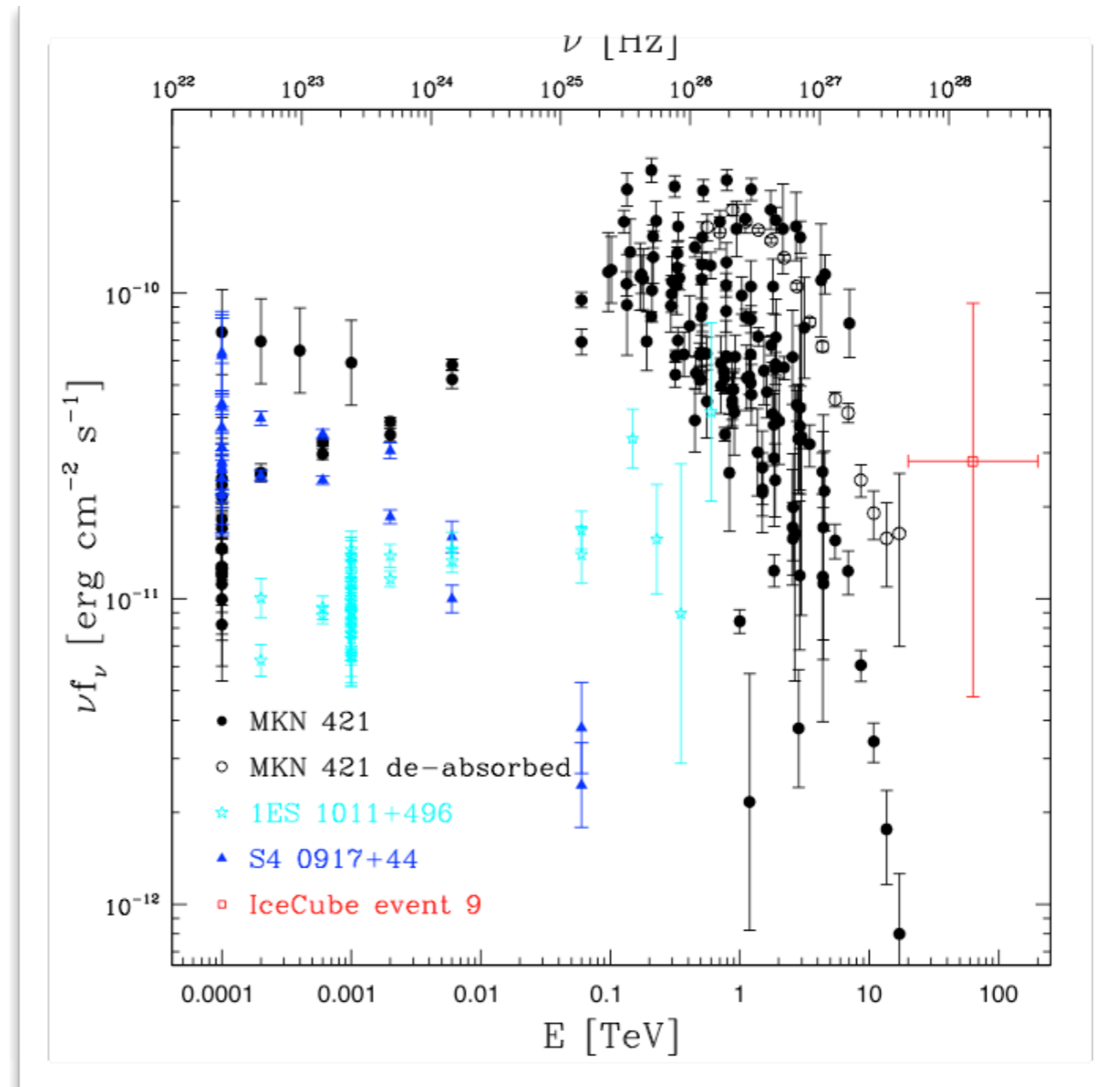
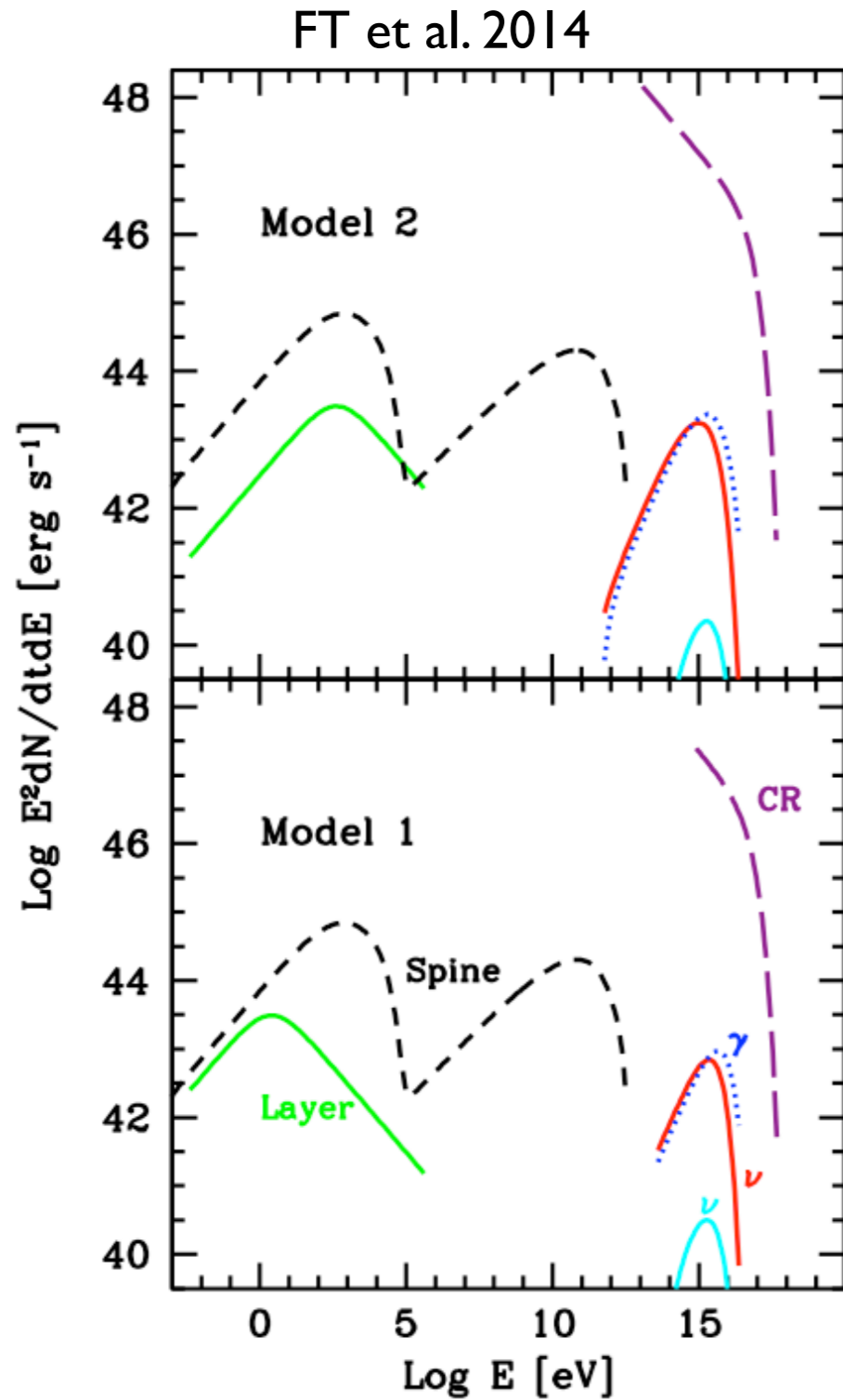
Neutrinos from FSRQ?

Strong correlation expected!

Murase, Inoue & Dermer 2014

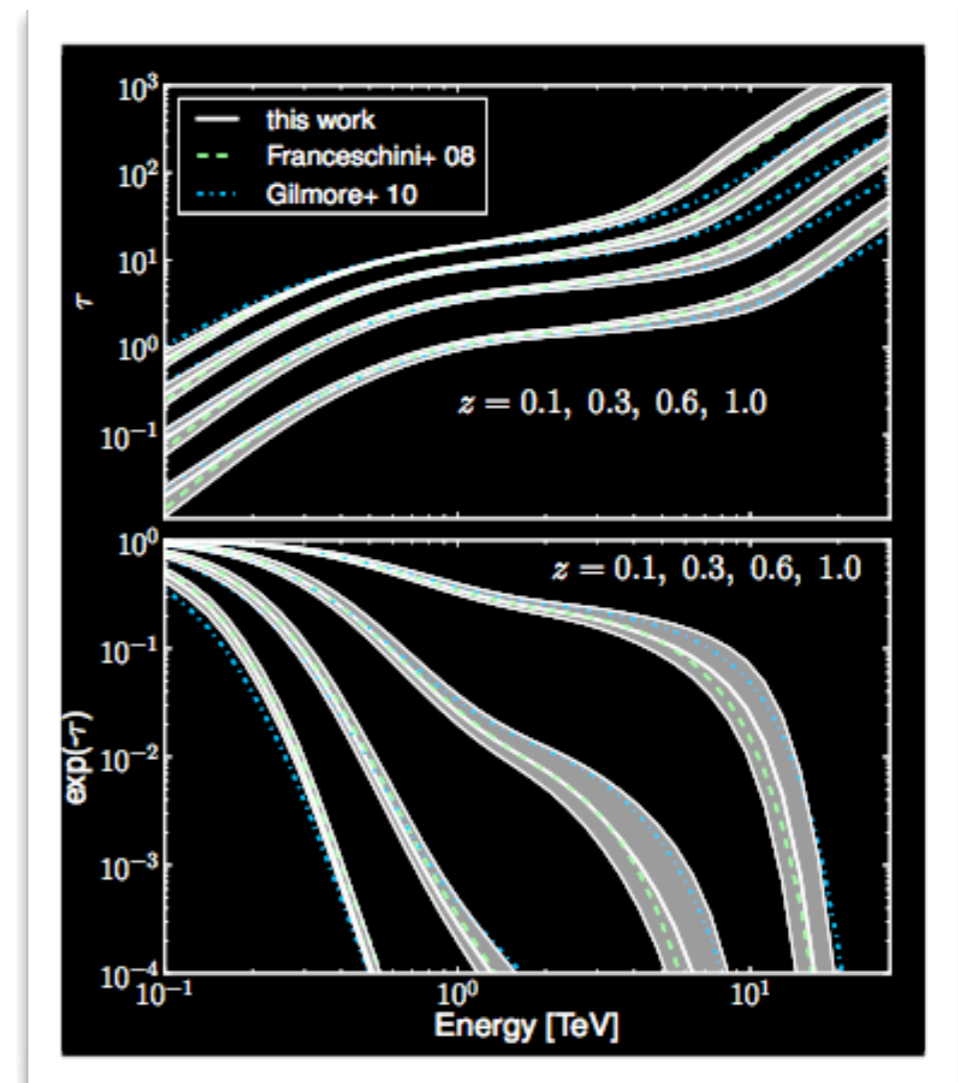
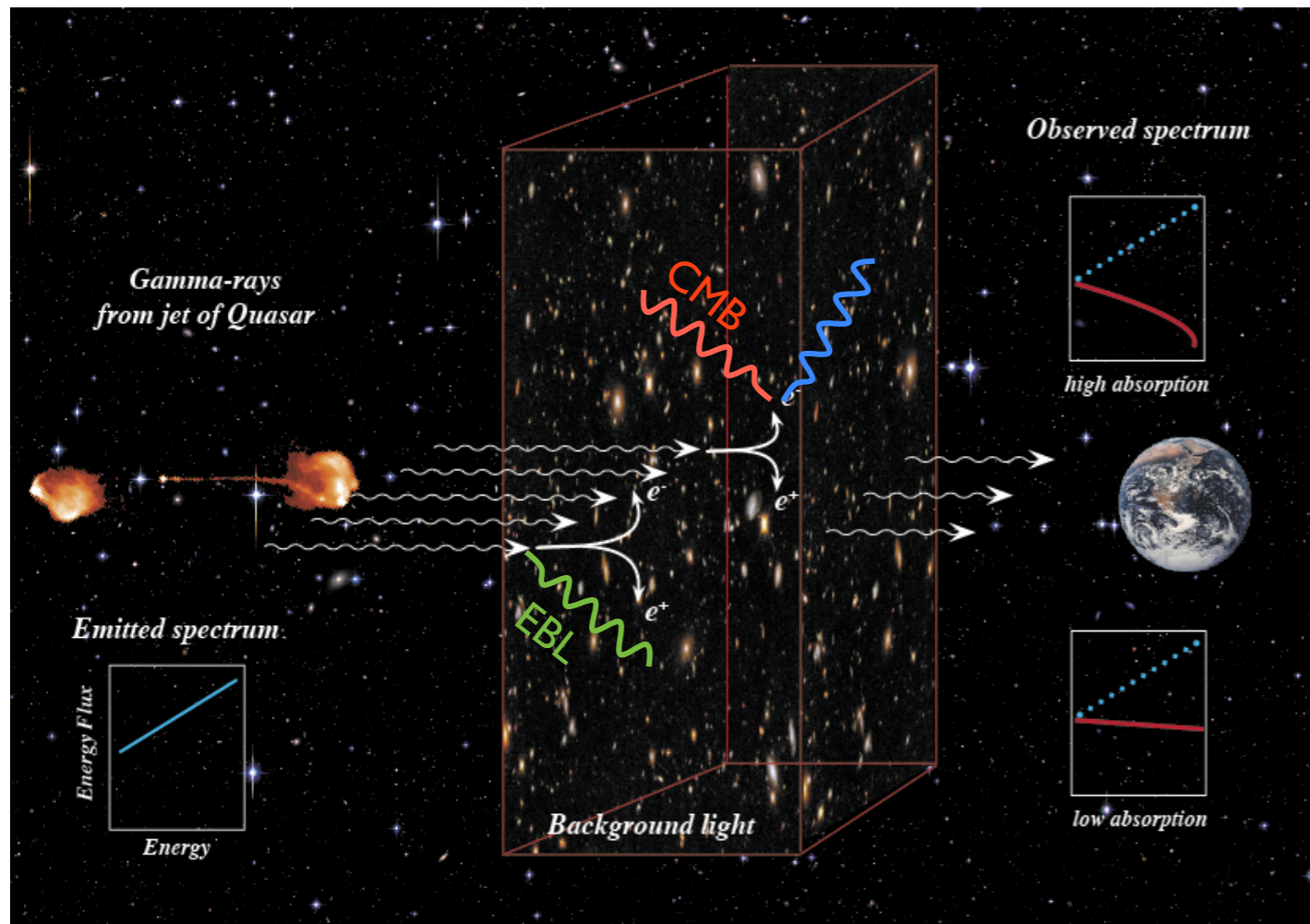


Structured jets: neutrinos



Padovani & Resconi 2014

Propagation: EBL absorption



Dominguez et al. 2011

Structured jets

$\theta < 5^\circ$

$\Gamma = 15-20$



$\theta_{\text{view}} \sim 15^\circ - 20^\circ$
for M87

Layer
 $\theta \sim 15^\circ$
 $\Gamma = 3-5$

