

Lecture 4

Photon capture radius
Schutz, Chap 11

Let's compute the trajectory of a photon near a black hole:

p_0 and p_φ are constant, $\tilde{E} \equiv -p_0$, $\tilde{L} \equiv p_\varphi$

We know that $E_\infty = h\nu_{\text{obs}} = -p_0 = \tilde{E}$

Assume: $\vartheta = \frac{\pi}{2}$, $p^\theta = 0$, λ is the affine parameter of geodesic

$$p_0 = -\tilde{E}, \quad p^0 = \frac{\tilde{E}}{1 - \frac{2GM}{r}}$$

$$p_r = \frac{1}{1 - \frac{2GM}{r}} \frac{dr}{d\lambda}, \quad p^r = \frac{dr}{d\lambda}$$

$$p_\varphi = \tilde{L}, \quad p^\varphi = \frac{\tilde{L}}{r^2}$$

$$\vec{p} \cdot \vec{p} = 0 = -\frac{\tilde{E}^2}{1 - \frac{2GM}{r}} + \left(\frac{dr}{d\lambda}\right)^2 \frac{1}{1 - \frac{2GM}{r}} + \frac{\tilde{L}^2}{r^2}$$

$$\Rightarrow \left(\frac{dr}{d\lambda}\right)^2 = \tilde{E}^2 - \left(1 - \frac{2GM}{r}\right) \frac{\tilde{L}^2}{r^2} = \tilde{E}^2 - \tilde{V}^2(r)$$

$$\tilde{V}^2(r) = \frac{\tilde{L}^2}{r^2} \left(1 - \frac{2GM}{r}\right) \quad \begin{array}{l} r \rightarrow \infty, \quad \tilde{V}^2 \rightarrow 0 \\ r = 2GM, \quad \tilde{V}^2 = 0 \end{array}$$

Extreme: $\frac{d\tilde{V}^2}{dr} = 0 \Rightarrow r = 3GM$ independent of \tilde{L}

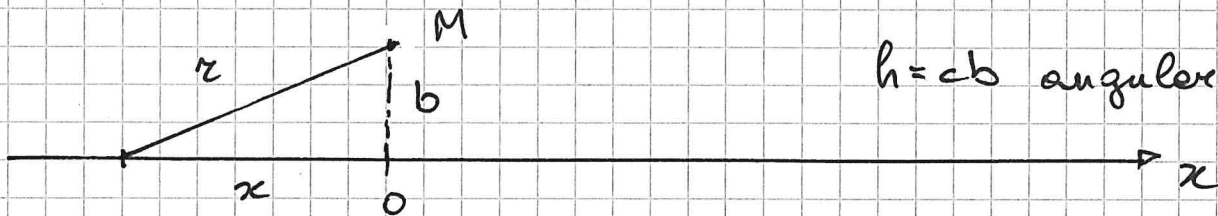
$$\tilde{V}_{\text{max}}^2 = \frac{\tilde{L}^2}{27(GM)^2}$$



There is an unstable orbit at $r = 3GM$

Photon trajectory in Newtonian dynamics

4.2



$h = cb$ angular momentum

$$x = ct, \quad r^2 = b^2 + x^2 = b^2 + c^2 t^2 \Rightarrow 2r \frac{dr}{dt} = 2c^2 t$$

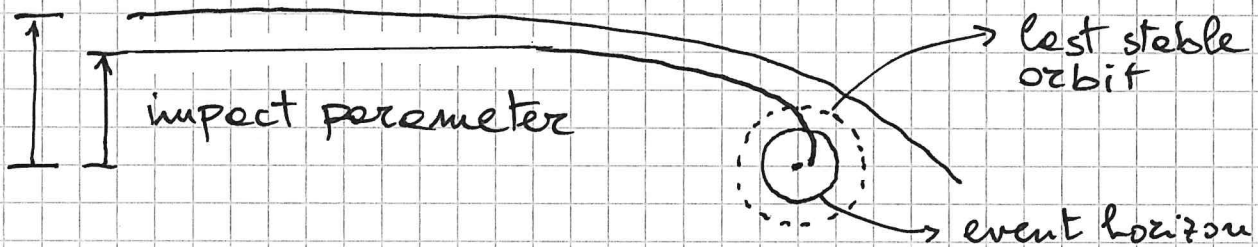
$$\Rightarrow \left(\frac{dr}{dt} \right)^2 = c^2 \left(1 - \frac{b^2}{r^2} \right), \quad V_{\text{eff}}(r) = \frac{c^2 b^2}{r^2} = \frac{h^2}{r^2}$$

the straight path of the photon implies a centrifugal barrier

→ Impact parameter for a particle around a black hole at infinity:

$$\tilde{L} = pb = h, \quad \tilde{E} = p \Rightarrow b = \frac{\tilde{L}}{\tilde{E}} = \frac{L}{E_\infty}$$

→ Capture radius of a black hole:



Impact parameter is what is seen from a distant observer.
Photon trajectory:

$$\left(\frac{dr}{d\lambda} \right)^2 = E_\infty^2 \left[1 - \left(1 - \frac{2GM}{r} \right) \frac{b^2}{r^2} \right]$$

$$\tilde{V}^2(r) = E_\infty^2 \left(1 - \frac{2GM}{r} \right) \frac{b^2}{r^2}$$

Photon capture radius: the smallest value of b before the photon falls on the BH

→ Orbit pericenter: $\frac{dr}{d\varphi} = 0$

$$\frac{dr}{d\varphi} = \frac{dr}{d\lambda} \frac{d\lambda}{d\varphi} = \frac{p^r}{p^\varphi}, \quad p^\varphi = \frac{b E_\infty}{r^2}$$

$$\left(\frac{dr}{d\varphi} \right)^2 = \frac{r^4}{b^2} \left[1 - \frac{b^2}{r^2} \left(1 - \frac{2GM}{r} \right) \right] = 0$$

$$\Rightarrow b^2 (r_{\text{peri}} - 2GM) = r_{\text{peri}}^3$$

but photons with $r_{\text{peri}} \leq 3GM$ are lost

$$b^2 \geq \frac{(3GM)^3}{(3GM - 2GM)} = 27(GM)^3$$

This defines the photon capture radius:

$$b \geq \sqrt{27} GM$$